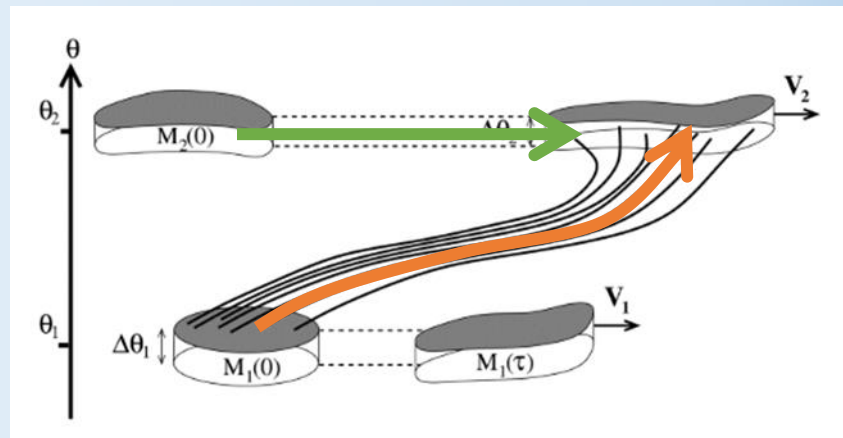


# Revisiting the isentropic view of PV modification in warm conveyor belts

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John Methven, Oscar Martinez-Alvarado

# Outline

- There are two (equally-valid?) frameworks for interpreting diabatic PV changes
  - Air parcel trajectories
  - Isentropic trajectories
- Why might isentropic trajectories be useful?
- An illustration: Cyclone Vladiana



Methven (2015)

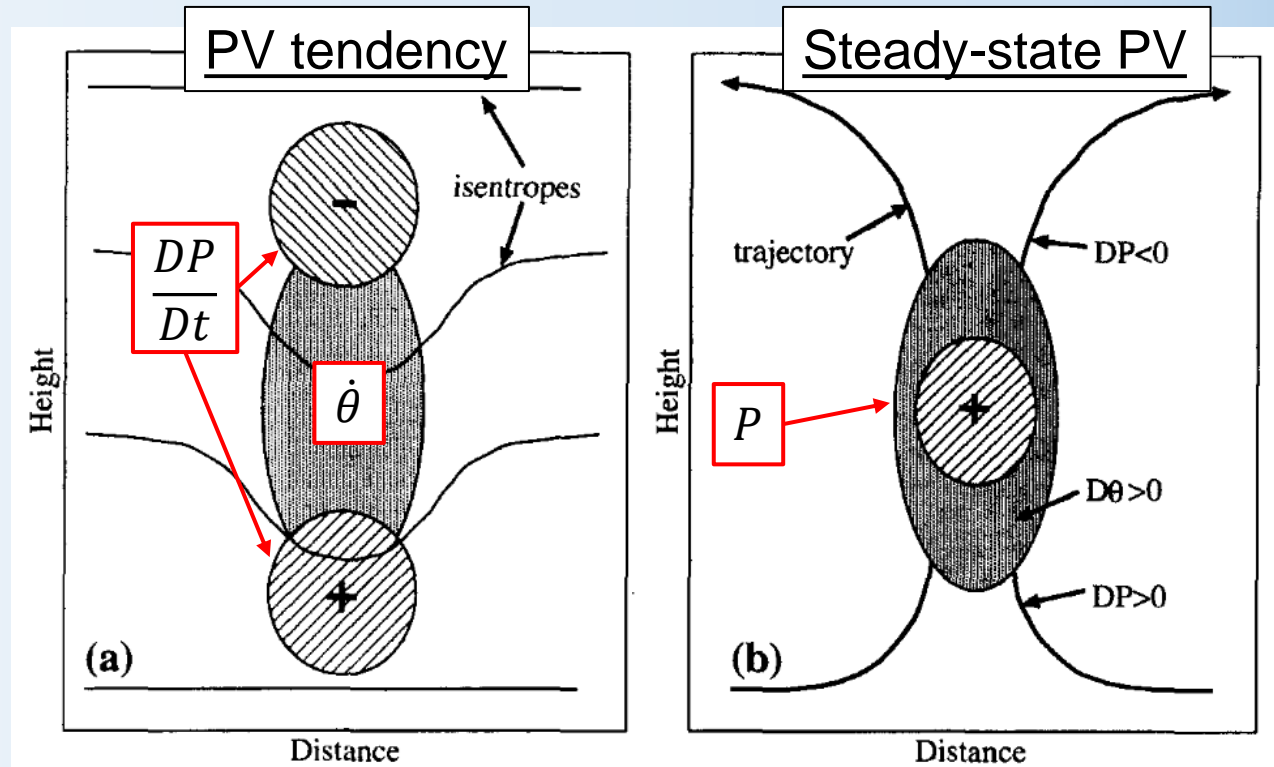
# PV modification along air parcel trajectories

Lagrangian PV equation:

$$\rho \frac{DP}{Dt} = \zeta \cdot \nabla \dot{\theta}$$

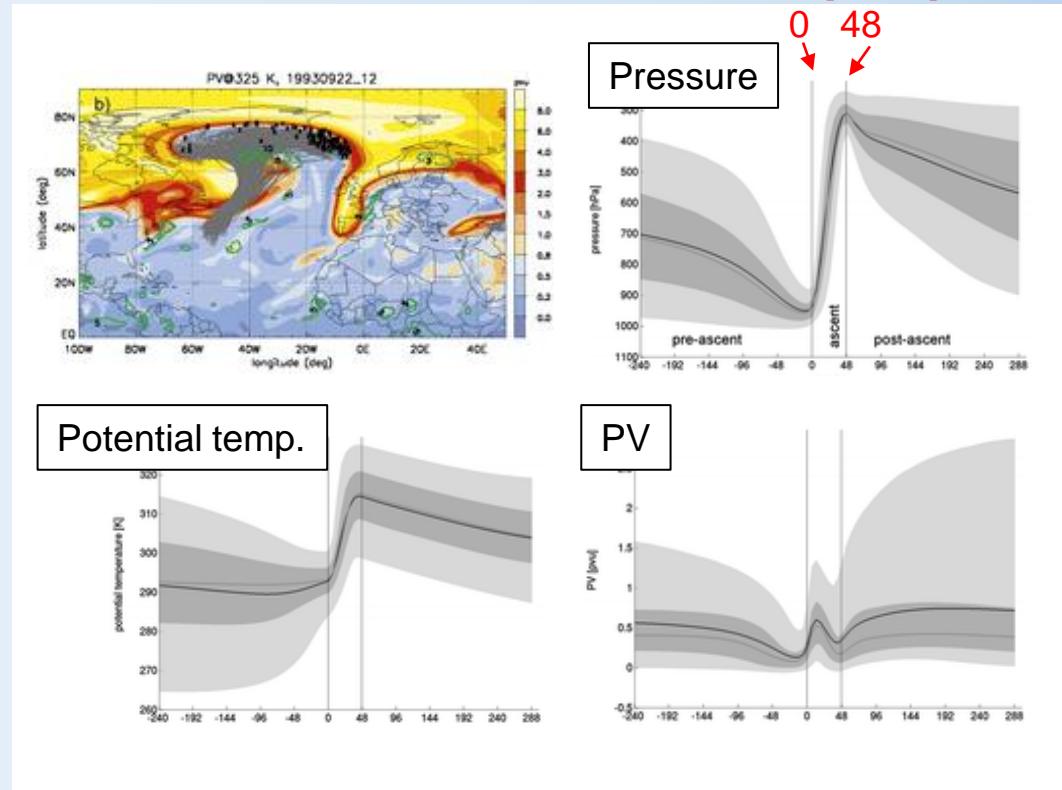
Absolute vorticity

Diabatic heating  
 $(\frac{D\theta}{Dt} = \dot{\theta})$

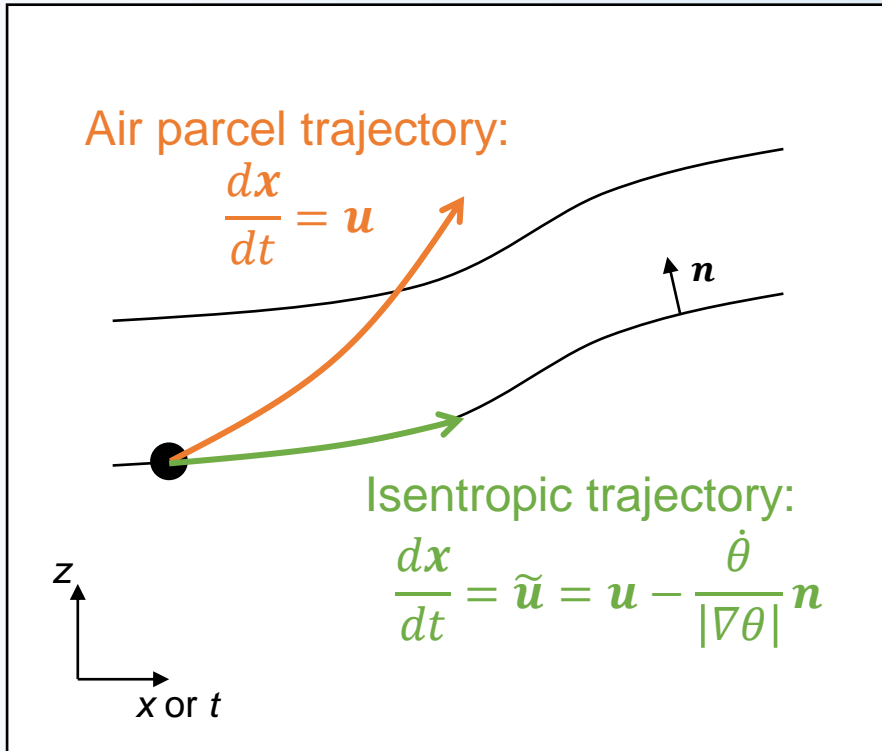


# PV modification along air parcel trajectories

- WCB trajectories: ascend by 600 hPa in 48 hours
- Composite evolution from ERA-Interim (North Atlantic, DJF)
- All trajectories warm (mean change = +20K)
- PV evolution is more complex: maximum value occurs mid-ascent



# Air parcel vs isentropic trajectories



- Air parcel trajectories follow the (resolved) wind
- Isentropic trajectories follow the flow along isentropic surfaces
- These are identical unless there is diabatic heating
- Note: isentropic framework fails if statically unstable

# PV modification along isentropic trajectories

Key Idea: Instead of thinking of PV as being attached to fluid parcels, think of it as being attached to isentropic surfaces:

$$\frac{\tilde{D}}{Dt} = \partial_t + \tilde{\mathbf{u}} \cdot \nabla \quad \text{with} \quad \tilde{\mathbf{u}} = \mathbf{u} - \frac{\dot{\theta}}{|\nabla\theta|} \mathbf{n}$$

The PV equation becomes:

$$\rho \frac{\tilde{D}P}{Dt} = P \frac{\partial}{\partial z} \left( \frac{\rho \dot{\theta}}{\theta_z} \right) - \theta_z \mathbf{k} \cdot \nabla_\theta \times \left( \mathbf{v}_z \frac{\dot{\theta}}{\theta_z} \right)$$

E.g. Haynes and McIntyre (1987)

Can we learn anything from isentropic trajectories that we don't already know from air parcel trajectories?

Suggest 3 **potential benefits** of using isentropic trajectories:

1. More natural **physical interpretation**

→ retain circulation / mass ideas from adiabatic dynamics

2. More **monotonic evolution** of PV following trajectories

→ a cleaner attribution of physical processes?

3. Less variation of **trajectory positions** with resolution

→ a fairer comparison across models?

→ a simpler evaluation of convection schemes?

# PV modification along isentropic trajectories

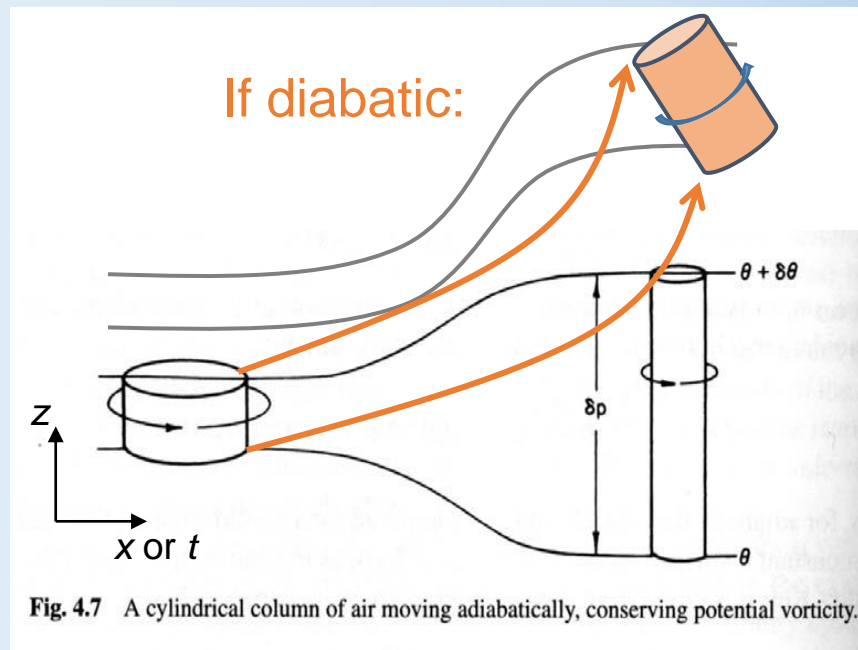
PV represents the mass-weighted circulation on an isentropic surface:

$$\langle P \rangle = \frac{c \delta\theta}{\mathcal{M}}$$

→ An exact, bulk formula

If adiabatic:

- Mass is conserved because we're following a material volume
- Circulation is conserved due to Kelvin's circulation theorem



Holton (2004)



# PV modification along isentropic trajectories

If instead we follow **isentropic trajectories**:

- Circuit remains aligned with isentropes (by definition)
- Relationship between  $\langle P \rangle$  and  $\mathcal{C}$  remains (+physical interpretation)

But...No longer following air parcels. Two issues:

- Mass (and circulation) are not conserved
- Conceptually more difficult?

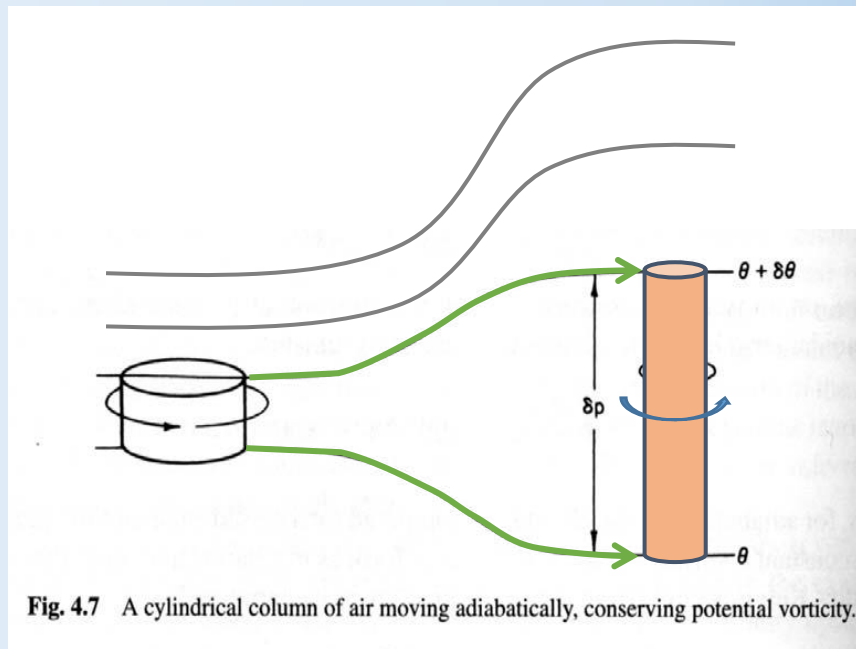


Fig. 4.7 A cylindrical column of air moving adiabatically, conserving potential vorticity.

Holton (2004)

# PV modification along isentropic trajectories

$$\langle P \rangle = \frac{c \delta\theta}{\mathcal{M}} \quad \rightarrow \quad \mathcal{M} \frac{d\langle P \rangle}{dt} = -\langle P \rangle \frac{d\mathcal{M}}{dt} + \delta\theta \frac{dc}{dt}$$

[recall isentropic PV equation:  $\rho \frac{\tilde{D}P}{Dt} = P \frac{\partial}{\partial z} \left( \frac{\rho \dot{\theta}}{\theta_z} \right) - \theta_z \mathbf{k} \cdot \nabla_{\theta} \times \left( \mathbf{V}_z \frac{\dot{\theta}}{\theta_z} \right)$ ]

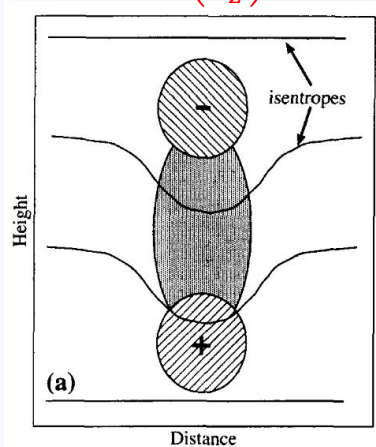
Two distinct physical mechanisms modifying PV:

1. PV concentration/dilution [diabatic mass flux convergence]
2. Diabatic circulation source/sink [Kelvin's circulation theorem]

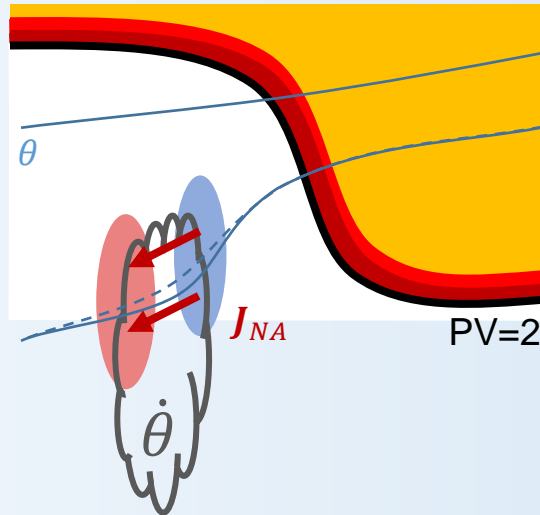
## PV dilution/concentration

- A vertical dipole: concentration below heating, dilution above
- Proportional to  $P$  so can't turn the PV negative

$$P \frac{\partial}{\partial z} \left( \frac{\rho \dot{\theta}}{\theta_z} \right)$$



$$-\theta_z \mathbf{k} \cdot \nabla_{\theta} \times \left( \mathbf{v}_z \frac{\dot{\theta}}{\theta_z} \right)$$



## Diabatic circulation source/sink

- Requires heating in the presence of vertical wind shear
- *Can* act to turn the PV negative
- Can be written as  $\nabla \cdot \mathbf{J}$  (the “non-advective PV flux”)
- $\mathbf{J}$  is always directed ‘down the isentropic slope’
- On an isentropic surface, there is an exact dipole: large-scale PV field does not ‘feel’ this term
- Scaling: importance grows at small scales – dominates on convective scale

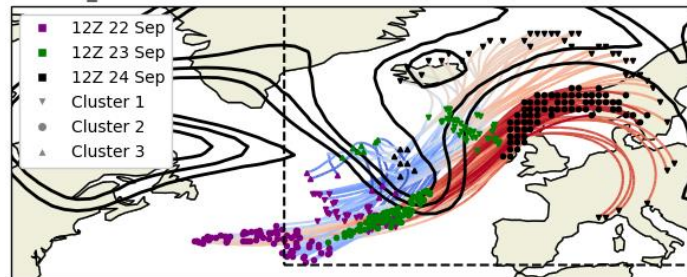
# Illustration: Cyclone Vladiana (Sep 2016; NAWDEX IOP3)

- N96 MetUM simulation of a 'fairly typical' cyclone ( $\approx 150$  km grid spacing)
- Store hourly model output including all diabatic and frictional tendencies

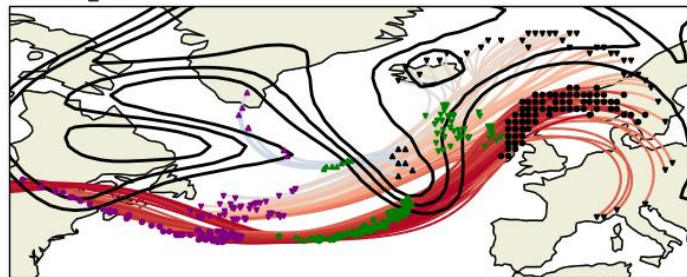
## Method:

1. Identify WCB trajectories in the standard way (ascend 500 hPa in 48 hours): *MAT\_WCB*
2. Compute isentropic trajectories backwards from the outflow region: *ISEN\_OUT*
3. Compute isentropic trajectories forwards from the inflow region: *ISEN\_IN*

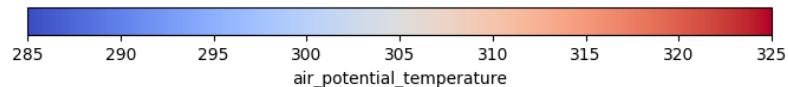
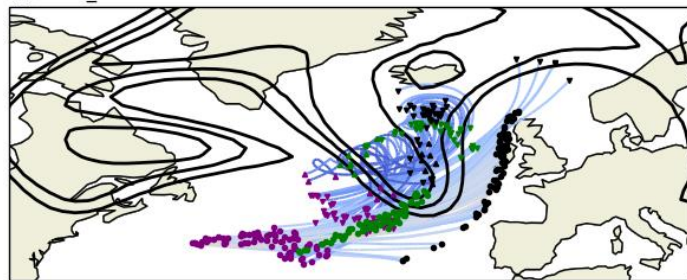
(a) MAT\_WCB



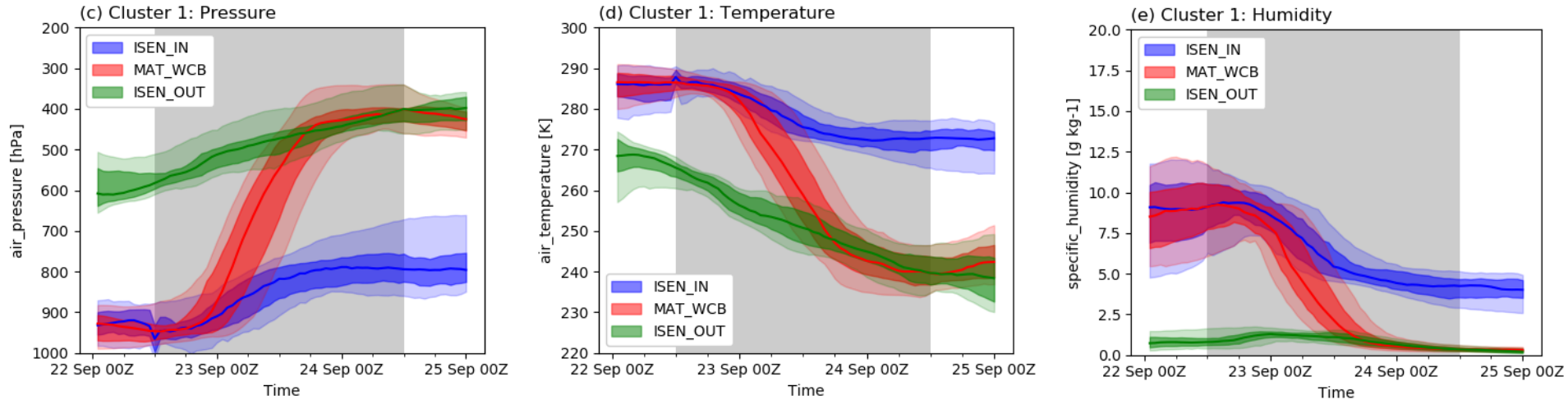
(b) ISEN\_OUT



(c) ISEN\_IN

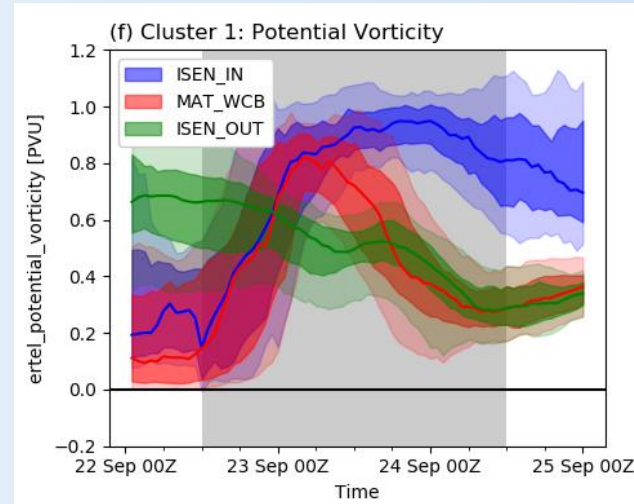
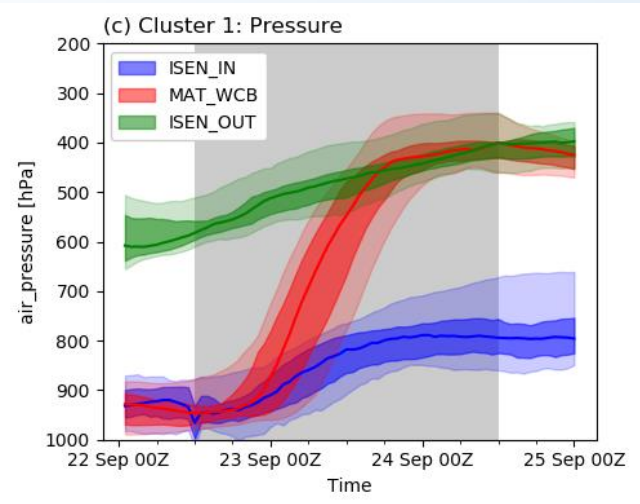


# Evolution of physical properties along trajectories



- Strong ascent and drying along *MAT\_WCB* trajectories (net heating=20K)
- Both *ISEN\_IN* and *ISEN\_OUT* trajectories still ascend (cf dry baroclinic wave) but by 200hPa instead of 600hPa

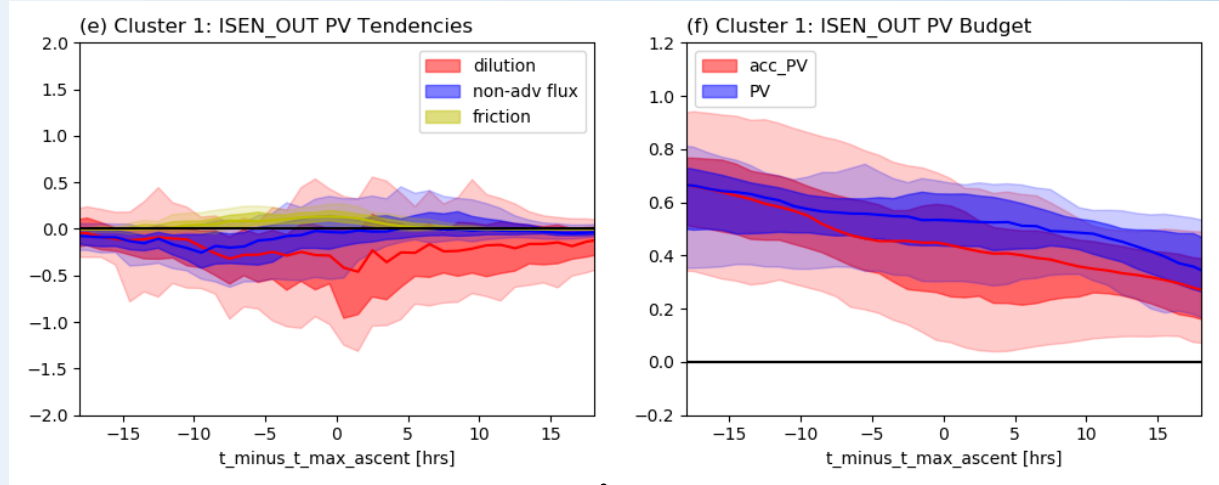
# Evolution of PV along trajectories



- Typical WCB PV evolution along *MAT\_WCB*:  
Increase whilst below 600 hPa, followed by decrease above
- Monotonic PV evolution along isentropic trajectories:  
Gradual decrease along *ISEN\_OUT*, increase along *ISEN\_IN*

# PV budget along isentropic trajectories

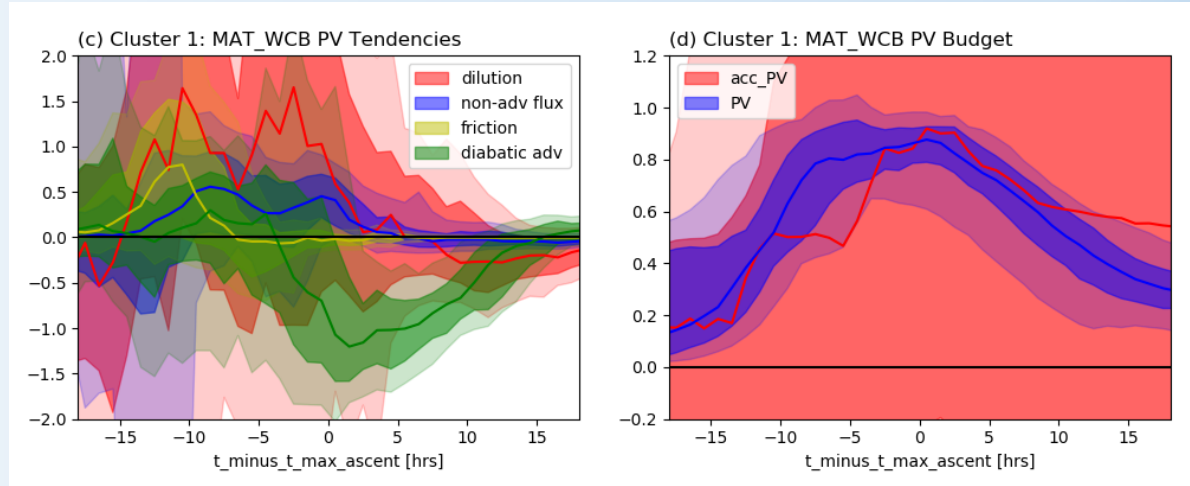
$$\rho \frac{\tilde{D}P}{Dt} = P \frac{\partial}{\partial z} \left( \frac{\rho \dot{\theta}}{\theta_z} \right) - \theta_z \mathbf{k} \cdot \nabla_{\theta} \times \left( \mathbf{V}_z \frac{\dot{\theta}}{\theta_z} \right)$$



- Compute 2 source terms using total  $\dot{\theta}$  (sum of all processes – but dominated by convective param here)
- PV decrease along *ISEN\_OUT* is dominated by **dilution term**
- Budget closes very well for 36 hour trajectories

# PV budget along air parcel trajectories

$$\rho \frac{DP}{Dt} = P \frac{\partial}{\partial z} \left( \frac{\rho \dot{\theta}}{\theta_z} \right) - \theta_z \mathbf{k} \cdot \nabla_{\theta} \times \left( \mathbf{V}_z \frac{\dot{\theta}}{\theta_z} \right) + \frac{\rho \dot{\theta}}{\theta_z} \frac{\partial P}{\partial z}$$



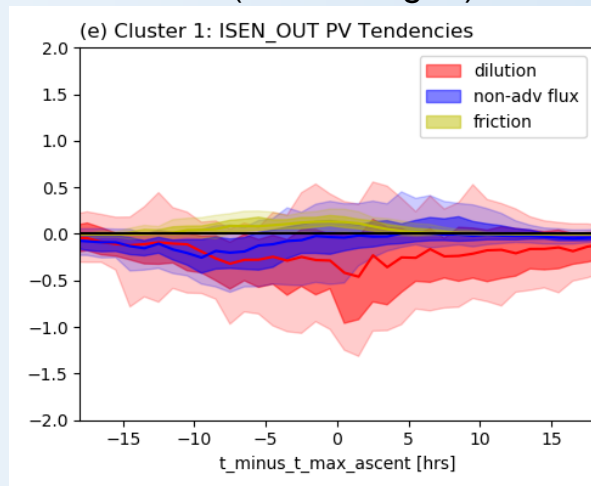
- PV evolution for *MAT\_WCB* more complex:
  - Initial increase due to PV concentration + circulation changes + friction
  - Later decrease dominated by advection across PV gradient
- Traj mean budget closes well, but huge spread whilst at low levels



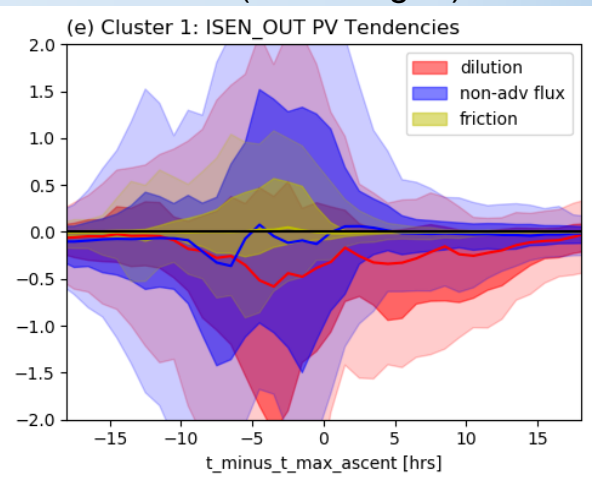
# Impact of model resolution

$$\rho \frac{\tilde{D}P}{Dt} = P \frac{\partial}{\partial z} \left( \frac{\rho \dot{\theta}}{\theta_z} \right) - \theta_z \mathbf{k} \cdot \nabla_{\theta} \times \left( \mathbf{V}_z \frac{\dot{\theta}}{\theta_z} \right)$$

N96 ( $\approx 150$  km grid)



N320 ( $\approx 50$  km grid)



- Much larger spread in circulation tendencies at N320 than N96 [ $\rightarrow$ strong horizontal PV dipoles emerging]
- But the mean changes from all 3 terms are almost identical

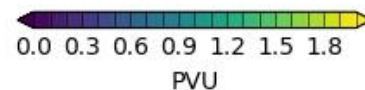
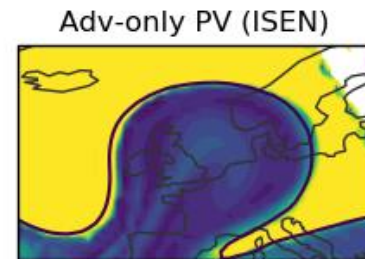
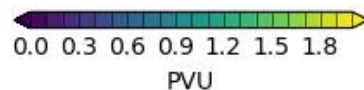
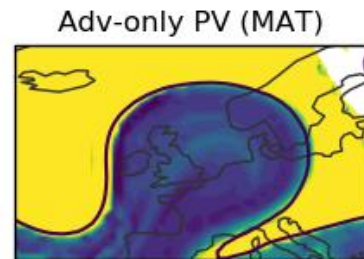
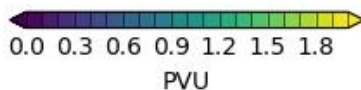
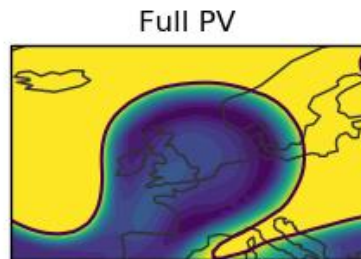
# Aside: What is diabatic PV?

Typically think of it as:

*the part of the PV field generated by diabatic processes during a certain time period*

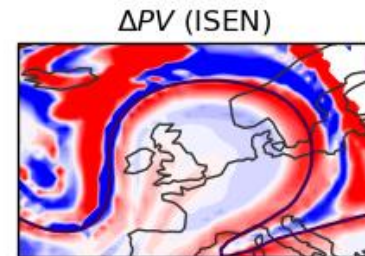
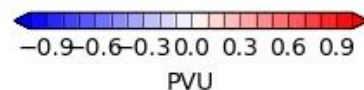
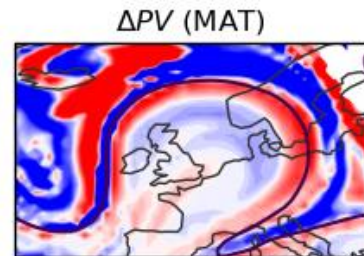
But this depends on the framework used:

**The PV generated by diabatic processes is different if you follow air parcel trajectories or isentropic trajectories**



Reverse domain-filling trajectories released at 12Z on 24 Sep

All panels show 325K



# Summary

- There are two (equally valid?) frameworks for understanding diabatic modification of PV
  - The air parcel view is used a lot in the literature
  - Can we learn anything extra from the isentropic view?
- Identified 3 potential benefits of the isentropic view
  1. A more natural physical interpretation
  2. Expect more monotonic changes in PV
    - allowing a cleaner attribution of physical processes?
  3. Expect isentropic trajectory positions to vary less with resolution
    - providing a fairer comparison across models?  
and/or evaluation of convection parametrisation schemes?