

# The ICON model: actual state and first steps towards a new dynamical core based on the Discontinuous Galerkin method

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## Outline

- The **ICON** dyn. core, recent developments: deep atmosphere
- **Discontinuous Galerkin** scheme as a possible alternative dynamical core for ICON - results from a **2D toy model**
  - Introduction
  - The HEVI approach
  - DG on the sphere
- First steps towards an **ICON-prototype** based on DG: **BRIDGE**

(short break)

## Properties of the dynamical core of ICON

- uses non-hydrostatic, compressible Euler eqns.
  - exactly mass- and tracer mass-conserving
  - true 2<sup>nd</sup> order scheme (as long as parameterizations are switched off)
  - stable in very steep mountainous regions
  - useable both for global and regional applications
  - computationally very efficient and scales well on current parallel computers
- (Zängl et al. (2015) QJRM, Zängl (2012) MWR, Baldauf, Reinert, Zängl (2014) QJRM)



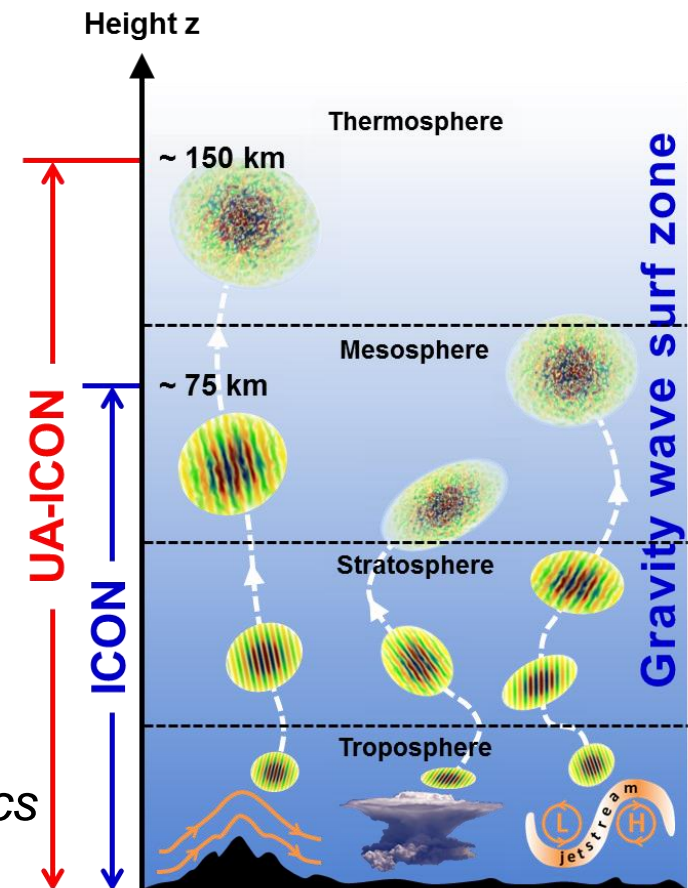
### Some numerical details:

- staggering: horizontal: icosahedral, triangle C-grid, vertical: Lorenz-grid
- mixed finite-volume / finite-difference
- predictor-corrector time-integration, horiz. explicit – vertic. implicit (HEVI)
- several damping mechanisms are used (off-centering, divergence damping, horizontal Smagorinsky diffusion and 4th order artificial diffusion, ...)

# Modification of ICON\* for the deep atmosphere: Motivation

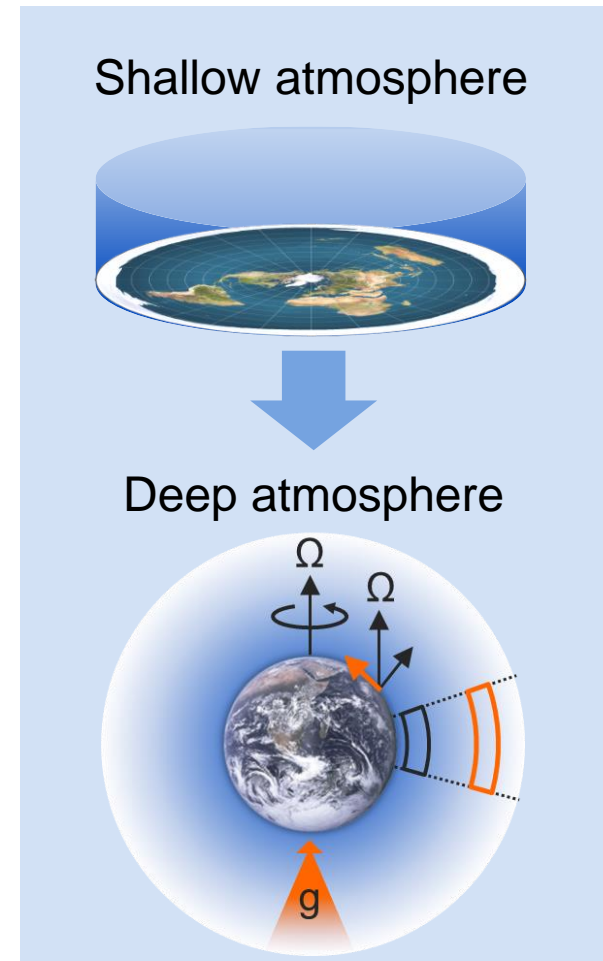
S. Borchert

- Work is part of DFG research group: **Multiscale Dynamics of Gravity Waves** (MS-GWaves\*\*)
- Collaboration with colleagues from Max Planck Institute for Meteorology (Hamburg)
- Goal: simulation of large part of gravity wave life cycle, from sources in Troposphere to wave breaking in upper Mesosphere lower Thermosphere
- Implementation of upper-atmosphere physics package (by MPI-M)
- Implementation of *deep-atmosphere dynamics* (by DWD)

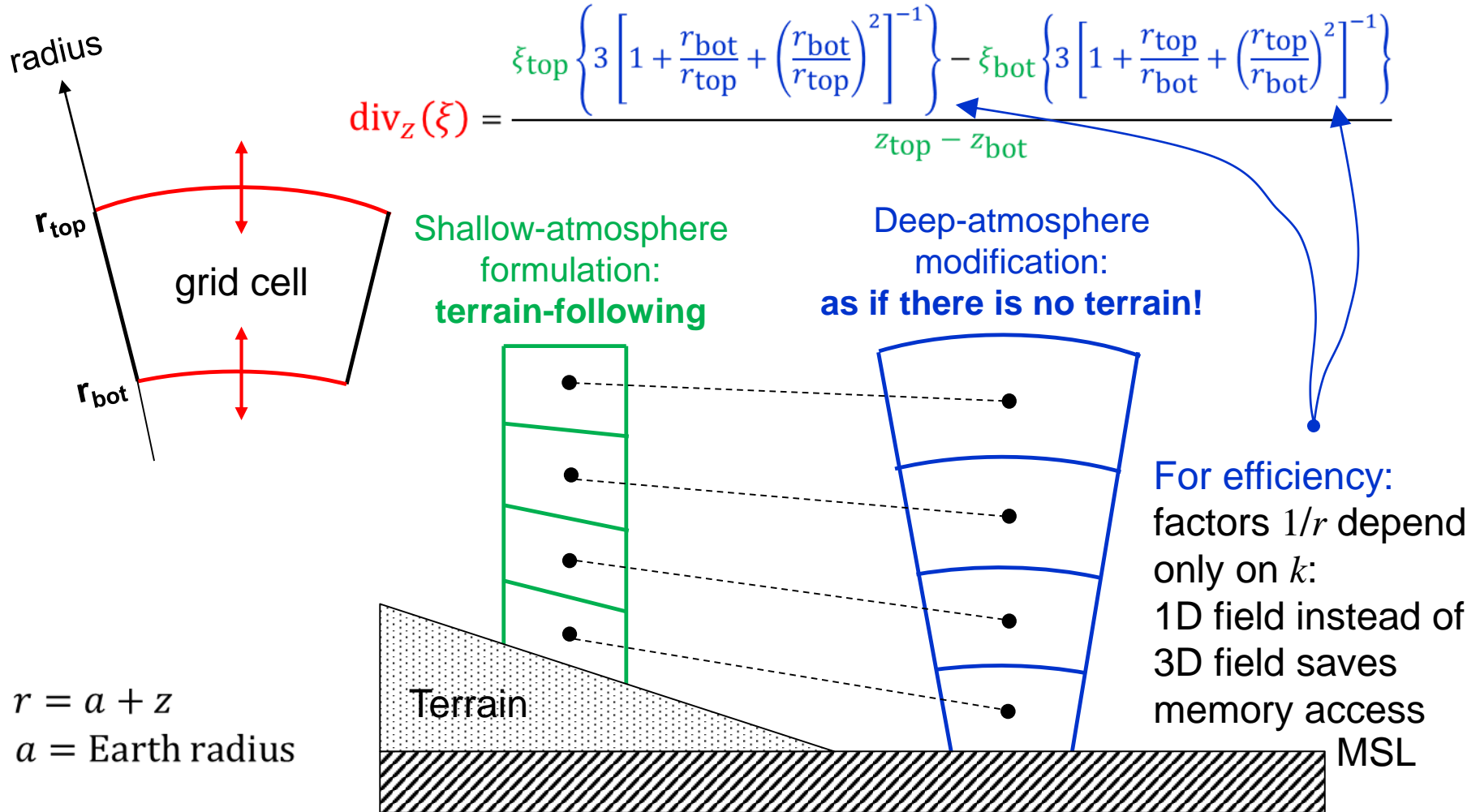


# Modification of ICON for the deep atmosphere: Overview

- **Shallow atmosphere\***
  - Standard configuration in ICON
  - In particular, replace prefactors  $1/r \rightarrow 1/a$
- **Deep-atmosphere modifications**
  - Abandon *shallow-atmosphere approximation*
    - Increase of grid cell extension with height
    - Gravitational field strength  $|g|$  decreases with height
  - Abandon *traditional approximation*
    - Coriolis acceleration due to  $\Omega_h$
    - take all metric terms in advection

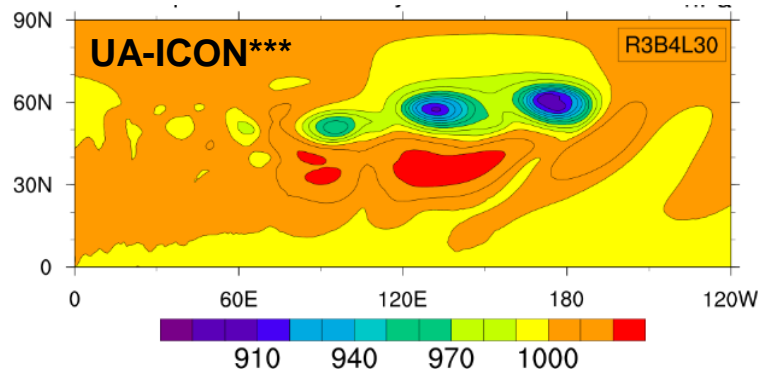
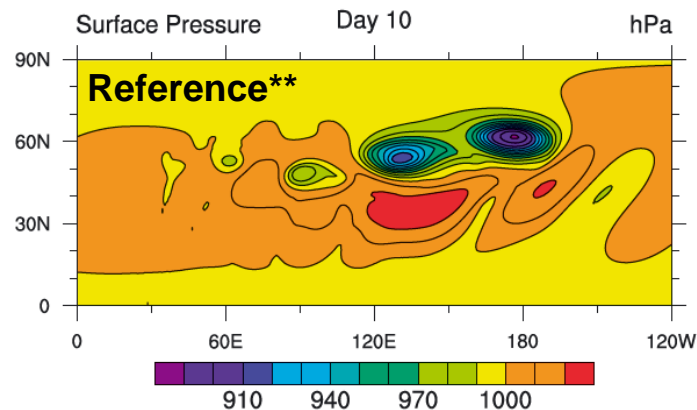


## Example: vertical divergence of some flux $\xi$



## Modification of ICON for the deep atmosphere: Test II

- Jablonowski-Williamson baroclinic instability test case\* in its extension for deep-atmosphere dynamical cores\*\*
- Focus on hydrostatic balance and baroclinic waves as important atmospheric synoptic-scale features
- No analytic solution available: model intercomparison



➔ If you are interested in the upper-atmosphere extension of ICON:  
\*\*\* Borchert, Zhou, Baldauf, Schmidt, Zängl, Reinert (2019) *The upper-atmosphere extension of the ICON general circulation model (version ua-icon-1.0)*, GMD



## Time table (at DWD)

- Jan. 2015: ICON (13km) replaces GME as a global forecast model
- Jan. 2016: ICON-EU-nest (6.5km) replaces COSMO-EU (7km)
- Jan. 2021: ICON-D2 (2.1km) replaces the convection-permitting COSMO-D2 (2.2km)  
similar replacements at COSMO partner countries until ~2023

## Current, ongoing dynamical core developments:

- Replace continuity equation for total mass by those for dry mass  
→ this has a lot of implications in dyn. core itself, in physics-dyn.-coupling, in boundary conditions, ...  
(D. Reinert (DWD), ... KIT)



## Michael Baldauf (DWD)

[illegible]

## Discontinuous Galerkin (DG) methods in a nutshell (I)

$$\frac{\partial q^{(k)}}{\partial t} + \nabla \cdot \mathbf{f}^{(k)}(q) = S^{(k)}(q), \quad k = 1, \dots, K$$

e.g.

Cockburn, Shu (1989) *Math. Comput.*

Cockburn et al. (1989) *JCP*


Hesthaven, Warburton (2008)

1.) weak formulation  $\int_{\Omega_j} dx v(\mathbf{x}) \cdot \dots$

$$\Rightarrow \frac{d}{dt} \int_{\Omega_j} q^{(k)} v dV + \int_{\partial\Omega_j} f^{(k)num,\perp} v da - \int_{\Omega_j} \mathbf{f}^{(k)} \cdot \nabla v dV = \int_{\Omega_j} S^{(k)} v dV$$

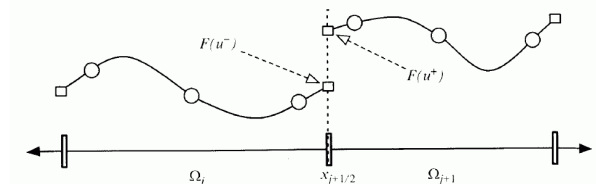
2.) Finite-element ingredient

$$q^{(k)}(x, t) = \sum_{l=0}^p q_{j,l}^{(k)}(t) p_l(x - x_j)$$

Galerkin-idea: identify  $v \equiv p_l$  

Modal base: orthogonal functions e.g. Legendre-Polynomials

Nodal base: interpolation (Lagrange) polynomials



From Nair et al. (2011) in  
'Numerical techniques for global atm.  
models'

## Discontinuous Galerkin (DG) methods in a nutshell (II)

Weak formulation

$$\frac{d}{dt} \int_{\Omega_j} q^{(k)} v dV + \int_{\partial\Omega_j} f^{(k)num,\perp} v da - \int_{\Omega_j} \mathbf{f}^{(k)} \cdot \nabla v dV = \int_{\Omega_j} S^{(k)} v dV$$

### 3.) Finite-volume ingredient:

Replace physical flux by a numerical flux in the surface integral

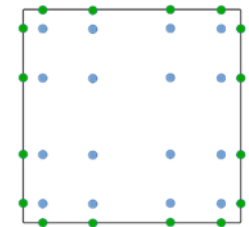
→ couple two neighbouring cells

Often used: simple Lax-Friedrichs flux

$$\mathbf{f}(q) \rightarrow f^{num,\perp}(q^+, q^-) = \frac{1}{2} (\mathbf{f}(q^+) + \mathbf{f}(q^-)) \cdot \mathbf{n} - \frac{\alpha}{2} (q^+ - q^-)$$

### 4.) Gaussian quadrature for the volume and surface integrals

→ ODE-system for  $q^{(k)}_{jl}(t)$



### 5.) Use a time-integration scheme (Runge-Kutta, ...)

- **local conservation** of every prognostic variable
- any **order of approximation (convergence)** possible
- flexible application on **unstructured grids** (also dynamic adaptation is possible, h-/p-adaptivity)
- very good **scalability** on massively-parallel computers (compact data transfer and no extensive halos)
- **separation** between (analytical) equations and numerical implementation
- **boundary conditions** are easily prescribed (fluxes or values in weak form)  
→ coupling with other subcomponents (ocean model, ...) should be easy
- higher accuracy helps to **avoid several awkward approaches** of standard 2<sup>nd</sup> order schemes: staggered grids (on triangles/hexagons, vertically heavily stretched), numerical hydrostatic balancing, grid imprints by pentagon points or along cubed sphere lines, ...
- **unified numerical treatment** of all flux terms and source terms
- **explicit** schemes are relatively easy to build and are quite well understood

- **high computational costs** due to
  - (apparently) **small Courant numbers** → small time steps
  - higher number of **degrees of freedom**
    - variables ,live' both on interior *and* on edge quadrature points
    - this holds additionally for **parabolic problems (diffusion)**
  - HEVI approach leads to **band diagonal matrices** with many bands
- **well-balancing** (hydrostatic, perhaps also geostrophic?) in Euler equations is an issue (can be solved!)
- basically ,only' an **A-grid-method**, however, the ,spurious pressure mode' is very selectively damped!

Current status:

we are still relatively far away from a full-fledged meteorological model,  
only a **toy model for 2D problems exists** with:

- explicit time integration DG-RK (with Runge-Kutta schemes) or horizontally explicit-vertically implicit (DG-HEVI) (with IMEX-Runge-Kutta)
- ‚local DG‘ (LDG) option for PDEs with higher spatial derivatives (e.g. diffusion)
- use of a triangle grid (also on the sphere) is optional

Some examples on the next slides ...

## 2D Euler-equations (non-hydrostatic, compressible) with diffusion (=Navier-Stokes eqns.) with terrain-following coordinates $(x, z')$ on a plane

$$\begin{aligned}
 \frac{\partial}{\partial t} \sqrt{G'} \rho' + \frac{\partial}{\partial x} \sqrt{G'} M^{*x} + \frac{\partial}{\partial z'} \sqrt{G'} \left( \frac{\partial z'}{\partial x} M^{*x} + \frac{\partial z'}{\partial z} M^{*z} \right) &= 0, & M^{*i} &= \rho v^i \\
 \frac{\partial}{\partial t} \sqrt{G'} M^{*x} + \frac{\partial}{\partial x} \sqrt{G'} T^{*xx} + \frac{\partial}{\partial z'} \sqrt{G'} \left( \frac{\partial z'}{\partial x} T^{*xx} + \frac{\partial z'}{\partial z} T^{*xz} \right) &= 0, \\
 \frac{\partial}{\partial t} \sqrt{G'} M^{*z} + \frac{\partial}{\partial x} \sqrt{G'} T^{*zx} + \frac{\partial}{\partial z'} \sqrt{G'} \left( \frac{\partial z'}{\partial x} T^{*zx} + \frac{\partial z'}{\partial z} T^{*zz} \right) &= -\sqrt{G'} g \rho' - \sqrt{G'} \frac{M^{*z}}{\tau}, \\
 \frac{\partial}{\partial t} \sqrt{G'} \eta' + \frac{\partial}{\partial x} \sqrt{G'} H^{*x} + \frac{\partial}{\partial z'} \sqrt{G'} \left( \frac{\partial z'}{\partial x} H^{*x} + \frac{\partial z'}{\partial z} H^{*z} \right) &= 0, & \eta &= \rho \theta \\
 p &= p_{ref} \left( \frac{\eta R_d}{p_{ref}} \right)^{cp/cv},
 \end{aligned}$$

Momentum fluxes:

$$\begin{aligned}
 T^{*xx} &= \frac{M^{*x} M^{*x}}{\rho} + p' - 2K\rho \left( \frac{\partial v^{*x}}{\partial x} + \frac{\partial z'}{\partial x} \frac{\partial v^{*x}}{\partial z'} \right), \\
 T^{*xz} \equiv T^{*zx} &= \frac{M^{*x} M^{*z}}{\rho} - K\rho \left( \frac{\partial z'}{\partial z} \frac{\partial v^{*x}}{\partial z'} + \frac{\partial v^{*z}}{\partial x} + \frac{\partial z'}{\partial x} \frac{\partial v^{*z}}{\partial z'} \right), \\
 T^{*zz} &= \frac{M^{*z} M^{*z}}{\rho} + p' - 2K\rho \frac{\partial z'}{\partial z} \frac{\partial v^{*z}}{\partial z'},
 \end{aligned}$$

Heat fluxes:

$$\begin{aligned}
 H^{*x} &= \frac{\eta M^{*x}}{\rho} - K\rho \left( \frac{\partial \Theta}{\partial x} + \frac{\partial z'}{\partial x} \frac{\partial \Theta}{\partial z'} \right), \\
 H^{*z} &= \frac{\eta M^{*z}}{\rho} - K\rho \frac{\partial z'}{\partial z} \frac{\partial \Theta}{\partial z'}.
 \end{aligned}$$

*strong conservation form* with terrain following coordinates but cartesian base vectors  
(Wedi, Smolarkiewicz (2003) JCP, Schuster et al. (2014) MetZ, appendix, for the sphere)

# Horizontally explicit - vertically implicit (HEVI)-scheme with DG

*Motivation:* get rid of the **strong time step restriction** by vertical sound wave expansion in **flat grid cells** (in particular near the ground)

$$\frac{\partial q^{(s)}}{\partial t} + \underbrace{\nabla \cdot \mathbf{f}_{slow}^{(s)}}_{\text{explicit}} + \underbrace{\nabla \cdot \mathbf{f}_{fast}^{(s)}}_{\text{implicit}} = \underbrace{S_{slow}^{(s)}}_{\text{explicit}} + \underbrace{S_{fast}^{(s)}}_{\text{implicit}} \quad \mathbf{f}_{fast}^{(s)} = f_{z,fast}^{(s)} \mathbf{e}_z$$

$$f_{z,fast}^{(s)} = \sum_{s'} H^{ss'} q^{(s')}$$

- Use of **IMEX-Runge-Kutta** (SDIRK) schemes: SSP3(3,3,2), SSP3(4,3,3) (*Pareschi, Russo (2005) JSC*)
- The implicit part leads to several band diagonal matrices  
→ here a direct solver is used (expensive!)

References:

*Giraldo et al. (2010) SIAM JSC:* propose a HEVI semi-implicit scheme

*Bao, Klöfkorn, Nair (2015) MWR:* use of an iterative solver for HEVI-DG

*Blaise et al. (2016) IJNMF:* use of IMEX-RK schemes in HEVI-DG

*Abdi et al. (2017) arXiv:* use of multi-step or multi-stage IMEX for HEVI-DG





## IMEX-Runge-Kutta

- general stability function for the Dahlquist problem is known
- general order conditions are known
- described by double Butcher tableaus

e.g. SSP3(3,3,2) by *Pareschi, Russo (2005) JSC*:

0	0		
1	1	0	
1/2	1/4	1/4	0
	1/6	1/6	2/3

$a$	$a$		
$1-a$	$1-2a$	$a$	
1/2	$1/2-a$	0	$a$
	1/6	1/6	2/3

$$a = 1 - 1/\sqrt{2}$$

- practically SDIRK schemes are preferred

*Lock, Wood, Weller (2014) QJRM*

*Pareschi, Russo (2005) JSC*: SSP3(3,3,2), SSP3(4,3,3)

*Giraldo et al. (2012) Siam JSC*: ARK2(2,3,2)

*Kang, Giraldo, Bui-Thanh (2020) JCP*: IMEX-RK in hybridiz. DG

## 2D linear advection equation – DG-HEVI

maximum Courant numbers:

for a modal basis with Legendre polynomials of degree  $p$   
(i.e. convergence order  $p+1$ )

RK-IMEX scheme	time order	p=0	p=1	p=2	p=3
RK1-IMEX	1	1.0	0	0	0
Trap2	2	1.0	0.198	0.081	
SSP3(3,3,2)	2	1.25	0.378	0.199	
ARK2(2,3,2)	2		0.219	0.110	
SSP3(4,3,3)	3		0.256	0.131	0.079

more specific:  $U_0^{slow}=1$ ,  $V_0^{slow}=0$ ,  $U_0^{fast}=0$ ,  $V_0^{fast}=10$ ,  $dx=dy=1$



Compare: explicit DG  $p=3$  /4th order RK: CFL=0.136

## A problem with the Euler equations ...

Approximate hydrostatic balance

**pressure gradient** ( $\rightarrow$  flux div.)  $\cong$  **buoyancy term** ( $\rightarrow$  source term)  
is crucial for the Euler equations in the atmosphere.

However, the source term integral contains base polynomials themselves, whereas the flux div. term integral uses *derivatives* of base polynomials:

$$\frac{d}{dt} \int_{\Omega_j} q^{(k)} v dV + \int_{\partial\Omega_j} f^{(k)num,\perp} v da - \int_{\Omega_j} \mathbf{f}^{(k)} \cdot \nabla v dV = \int_{\Omega_j} S^{(k)} v dV$$

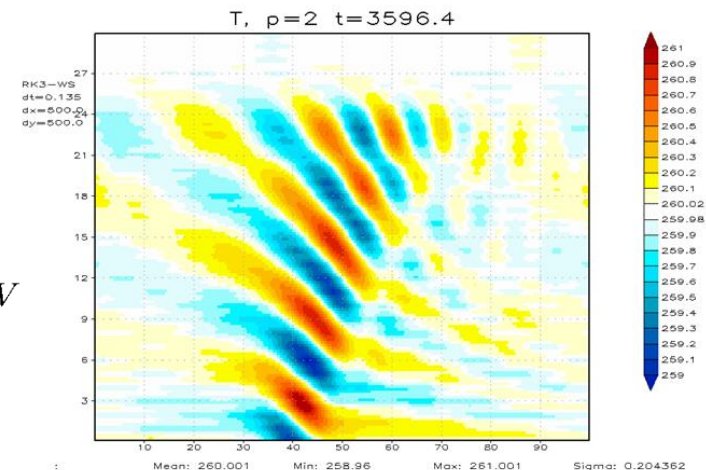
$\rightarrow$  no proper balance possible.

## ... and its solution:

*Blaise et al. (2016) IJNMF, Orgis et al. (2017) JCP:*

use vertically a reduced base (one polynomial degree less in a modal base) for the calculation of the source term.

This leads to an additional filtering procedure in the implicit solver.

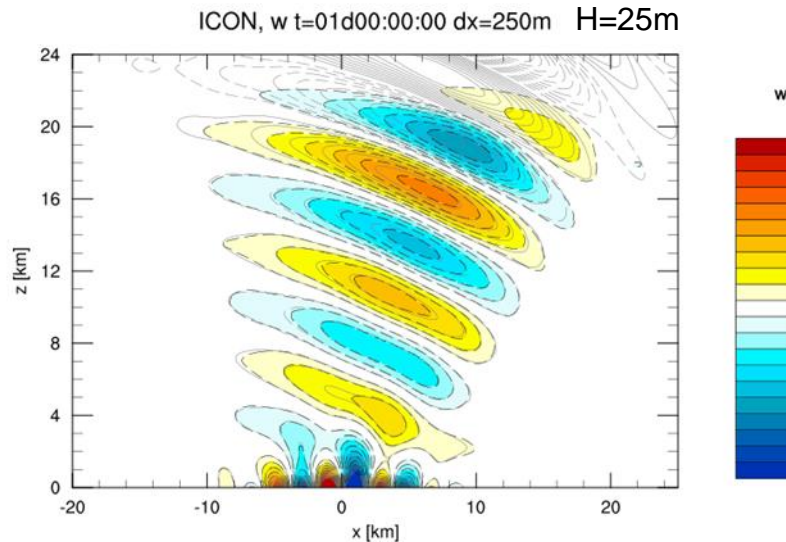


## Flow over mountain with the HEVI-solver

ICON

$\Delta x = 250\text{m}$

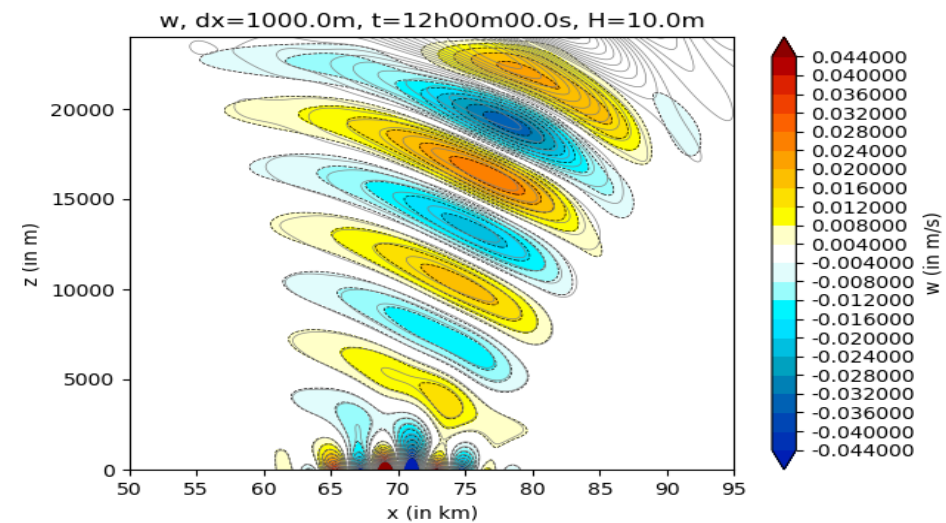
$\Delta z = 150\text{m}$



HEVI-DG 4th order

$\Delta x = 1000\text{m}$

$\Delta z = 600\text{m}$



colors and black dotted lines: model

grey lines: analytic solution (*Baldauf, 2008, COSMO-NewsI.*)

Computing time on 160 processors

(Cray XC40 Broadwell)

for  $t_{\text{total}} = 24\text{h}, 26\text{min}$

→ on 1 processor: 69h

Computing time on 1 processor:

(Intel(R) Core(TM) i7-4790 CPU @ 3.60GHz)

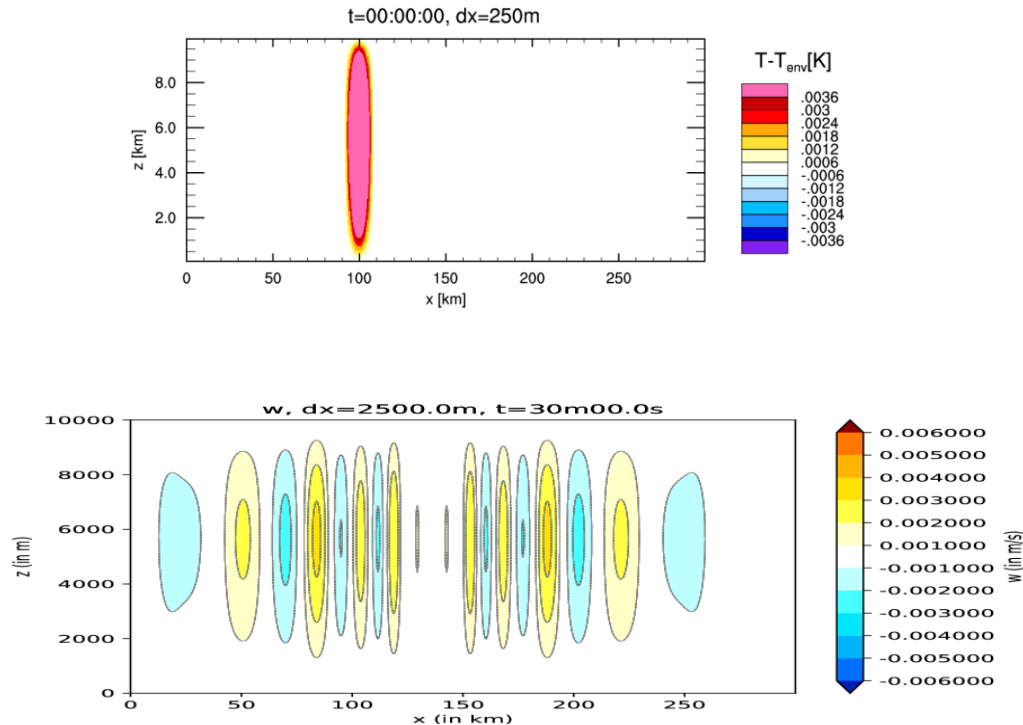
for  $t_{\text{total}} = 24\text{h}, 4\text{th order DG: } 160\text{h}$



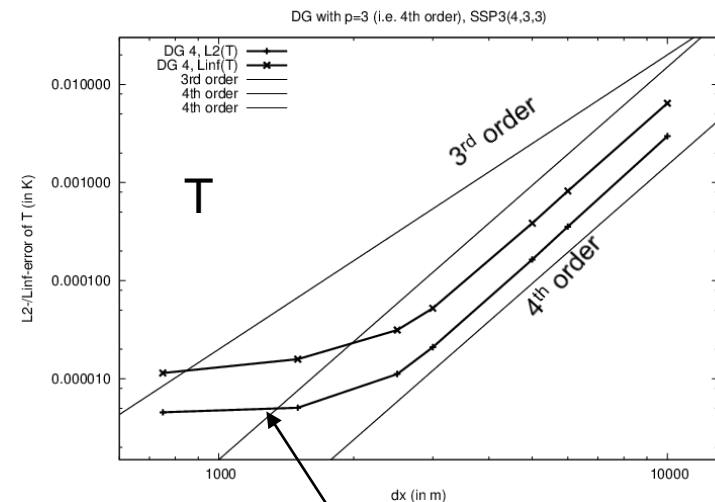
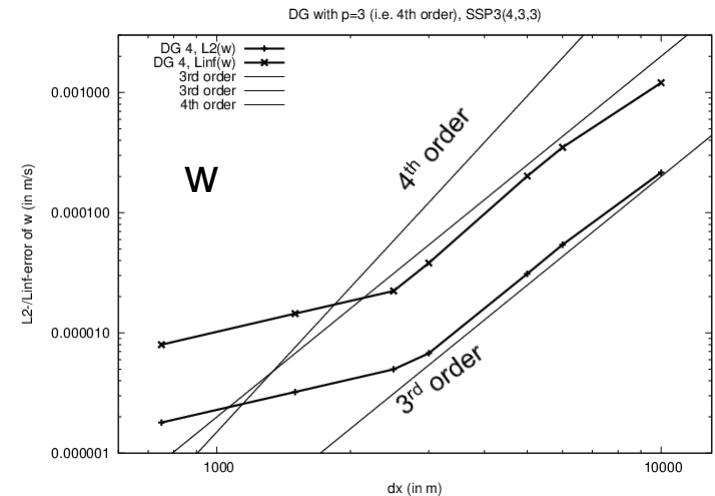
But don't take the computing  
times too serious!

# Linear gravity/sound wave expansion in a channel

setup similar to *Skamarock, Klemp (1994) MWR*



colors : simulation with  $p=3/SSP3(4,3,3)$   
 $dx=2500m, dz=1250m$   
black lines: analytic solution for compressible,  
non-hydrostatic Euler eqns.  
(Baldauf, Brdar (2013) QJRMS)



Nonlinear effect?

## General treatment of diffusion in DG

Simple replacement of a flux  $f(q)$  by  $f(q, \partial q / \partial x)$ , where  $\partial q / \partial x$  is directly calculated from  $q$ , does not work.

Instead (*Bassi, Rebay, 1997*, and similar for 'local DG'):

- define a new variable  $d = \partial q / \partial x$
- treat  $q$  as a flux and apply the DG formalism to this equation, too.
- Replace  $f(q) \rightarrow f(q) + f_{\text{diffus}}(q, d)$

Remark: the numerical flux for  $f_{\text{diffus}}$  does not need additional numerical diffusion.

## 3D-diffusion in terrain-following coordinates

New developments:

- In terrain-following coord. there are several choices for  $d$  possible.  
here: covariant derivatives of the prognostic variables
- If diffusion is treated vertically implicit (i.e. HEVI), too  
→ extension of the band diagonal matrix necessary
- Currently: 'HEVI-diffusion' is done in every sound time-step  
→ expensive! Appropriate time-integration necessary.

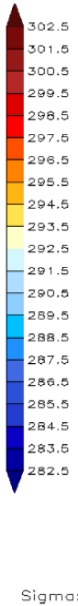
## ICON

## Falling bubble test case

Reference solution  
for  $\theta$   
Straka et al. (1993):

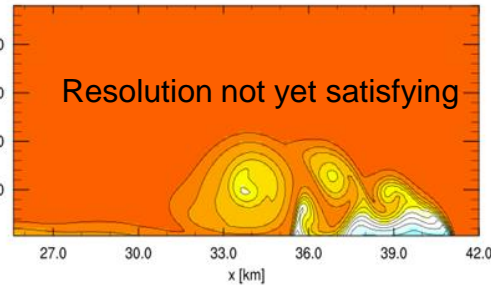
SSP3-4-3-3  
dt=0.07017543859649122  
dx=400.0  
dy=400.0

DG-HEVI

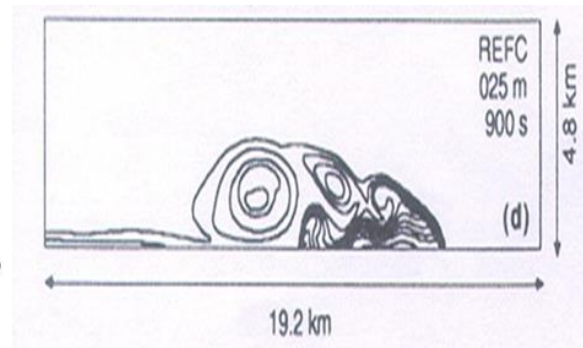
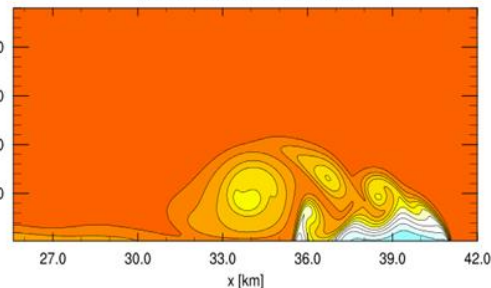


Sigma:

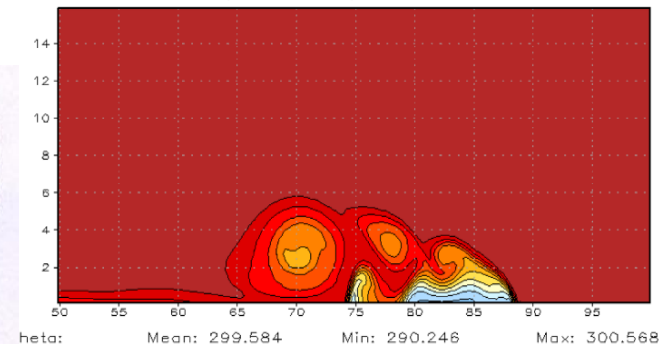
Theta t=00:15:00 dx=100m



Theta t=00:15:00 dx=50m



Theta, ord=4, t=15m00.0s



Computing time on 40 processors

(Cray XC40 Broadwell)

dx=dz=50m: 5min

→ on 1 processor: 3h30min

Computing time on 1 processor

(Intel(R) Core(TM) i7-4790 CPU @ 3.60GHz)

dx=dz=400m, 4th order DG

Euler HEVI, diffusion explicit: 1h40min

Euler + diffusion in HEVI: 3h50min



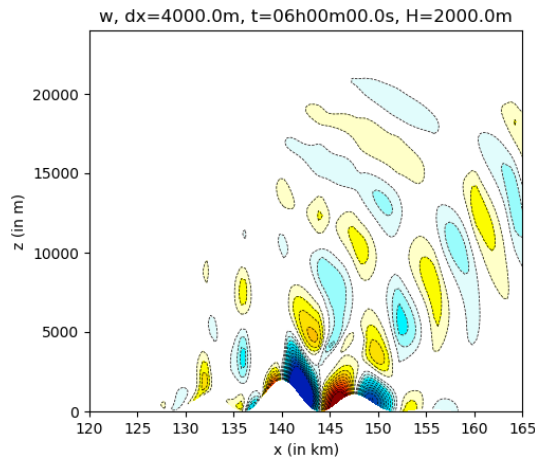
But don't take the computing  
times too serious!



## Flow over mountain with steep slopes and vertical grid stretching

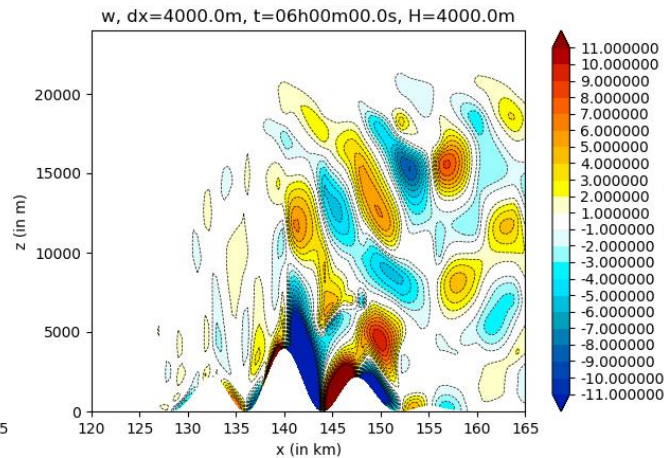
Schaer et al. (2002) MWR, test case 5b:  $U_0=10\text{m/s}$ ,  $N=0.01\text{ 1/s}$ , but  $a=10\text{km}$

$H_{\text{oro}} = 2000\text{m}$ ,  
 $\alpha_{\text{max}} = 38^\circ$



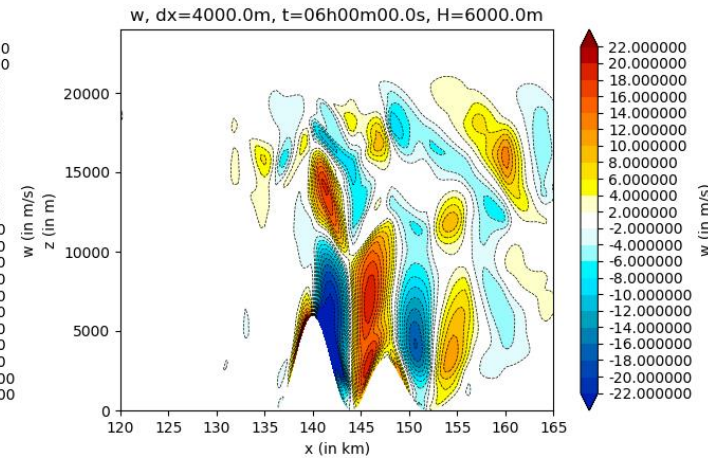
Sim: Min=-17.47014, Max=14.482281

$H_{\text{oro}} = 4000\text{m}$ ,  
 $\alpha_{\text{max}} = 58^\circ$



Sim: Min=-35.865414, Max=33.818314

$H_{\text{oro}} = 6000\text{m}$ ,  
 $\alpha_{\text{max}} = 67^\circ$



Sim: Min=-39.75445, Max=26.139639

$\Delta x = 4\text{ km}$ ; vertical grid stretching:  $\Delta z_{\text{min}} \sim 46\text{m}$ ,  $\Delta z_{\text{max}} \sim 736\text{m}$ ,  $z_{\text{lowest QP}} \sim 10.3\text{m}$

HEVI-DG simulation (4<sup>th</sup> order) remains stable even for steeper slopes!

to avoid instability by strong gravity wave breaking, vertically implicit '3D' diffusion with  $K=100\text{m}^2/\text{s}$  was used



## Efficiency of the implicit solver

The vertically implicit problem leads to the solution of  $J$  **block tridiagonal LES**, i.e  $J$  band diagonal matrices with  $KSMN$  rows and columns and  $2SMN$  diagonals both above and below the main diagonal

$J$  = number of grid cells in the horizontal direction

$K$  = number of grid cells in the vertical direction

$S$  = number of variables

$M$  = number of horizontal base functions

$N$  = number of vertical base functions

- Use of a **direct band diagonal solver**:
  - LU decomposition: complexity  $O( JK(SMN)^2 )$   
→ at most useable with low order DG  
efficiency gain: only once every several dozen time steps
  - Solution of the LES: one matrix-vector-mult. per column in every RK-substep: complexity  $O( JKSMN )$ 
    - compare with a tridiagonal solver in our 'standard' FD/FV codes:  $O( JM KN )$
- Use of an **iterative solver** (*Bao, Klöfkorn, Nair (2015) MWR*)

## How to bring DG on the sphere

*annoying fact:* the sphere doesn't allow a single coordinate system without singularities → leads to **pole problem** in lat-lon coordinates

→ usual remedies:

- *Yin-Yang* grid (often with quadrilateral cells)
- *cubed-sphere* grid (often with quadrilateral cells)
- *Icosahedron* with hexagonal (e.g. MPAS) or **triangle** cells (e.g. ICON)

additionally: allow higher order (i.e.  $>2^{\text{nd}}$ ) discretization

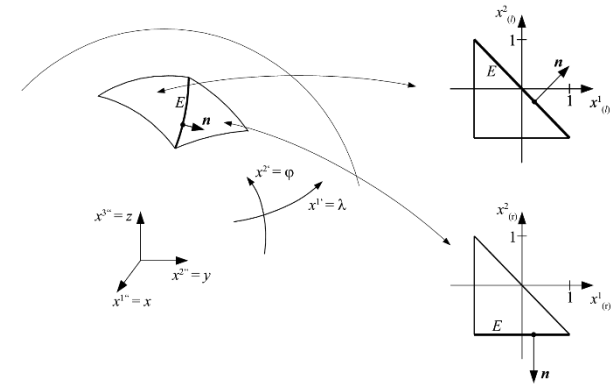
- use **local coordinates** for every (triangle) grid cell,  
= locally rotated gnomonial projection (*Läuter, Giraldo, ... (2008) JCP*)
- geometry is treated exactly!

Idea here: fields components are expressed in local base vectors, too

- no pole problem at all
- transform fluxes between neighbouring cells
- use covariant formulation of the equations

## Shallow water equations, covariant formulation

$$\begin{aligned}\frac{\partial \sqrt{g}H}{\partial t} + \frac{\partial}{\partial x^i} \sqrt{g}M^i &= 0 \\ \frac{\partial \sqrt{g}M^i}{\partial t} + \frac{\partial}{\partial x^j} \sqrt{g}T^{ij} &= \sqrt{g} \left( -g_{grav}H g^{ij} \frac{\partial h_B}{\partial x^j} + F_{Cor}^i - \Gamma_{jk}^i T^{jk} \right) \\ T^{ij} &= \frac{M^i M^j}{H} + \frac{1}{2} g^{ij} g_{grav} H^2, \quad M^i = H v^i\end{aligned}$$



$x^1, x^2$  are arbitrary **local coordinates** on each triangle.

This description is valid **on arbitrary 2D manifolds**.

→ extension to an **ellipsoid** is easy and without additional costs.

Formulate DG discretization only in these local coordinates

→ ,standard‘ DG formulation useable: nodal base, Runge-Kutta time integr., ...

**however: Transformation of fluxes** on the edges between neighbouring cells:

$$f_{(r \rightarrow l)}^i = \frac{\partial x_{(l)}^i}{\partial x_{(r)}^j} f_{(r)}^j, \quad f_{(l \rightarrow r)}^i = \frac{\partial x_{(r)}^i}{\partial x_{(l)}^j} f_{(l)}^j$$

Baldauf, M. (2020): *Discontinuous Galerkin solver for the shallow-water equations in covariant form on the sphere and the ellipsoid*, J. Comp. Phys. 410

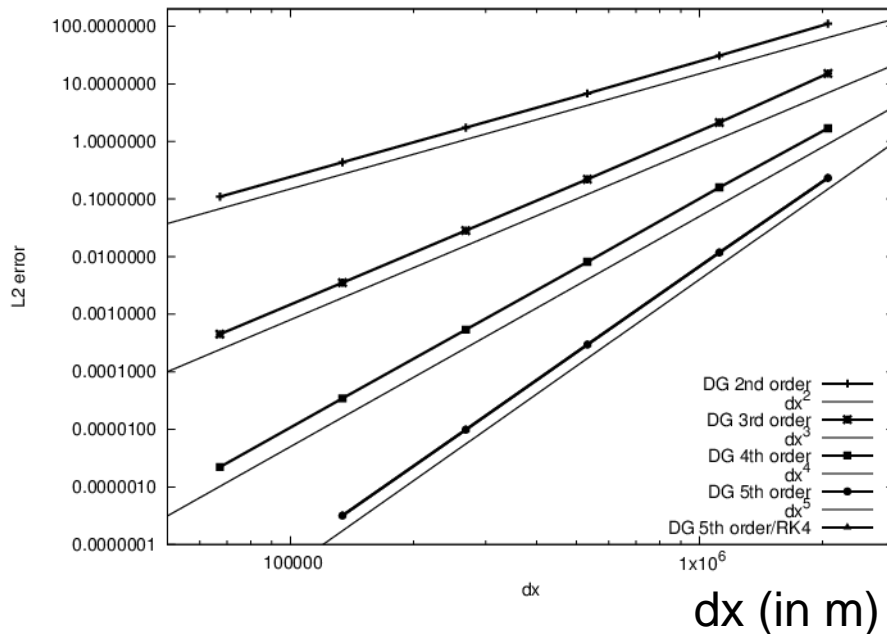
## Unsteady solid body rotation

Läuter et al. (2005) JCP

Exact analytic solution available! → calculate error measures  $L_2$  (=RMSE)

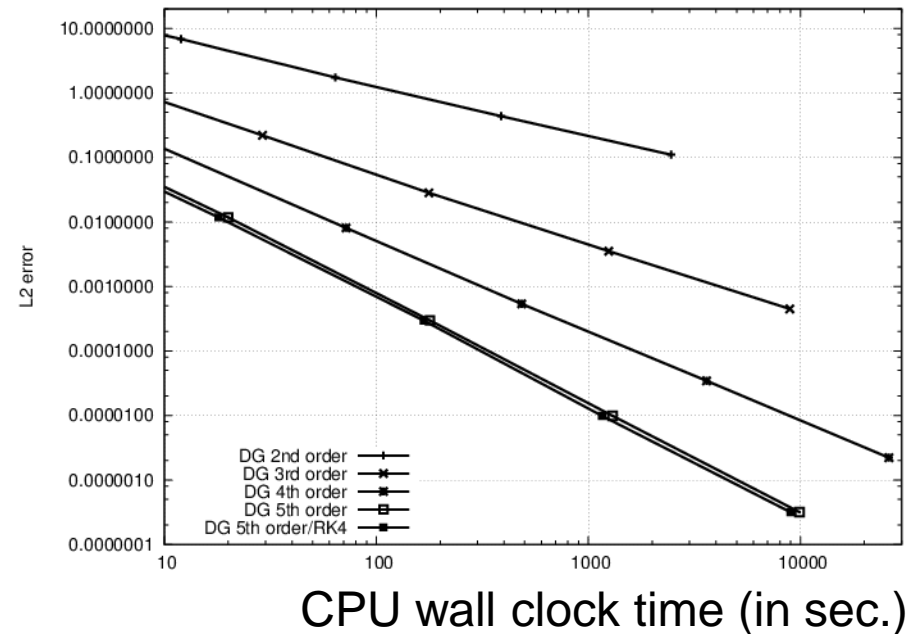
### Convergence plot:

Convergence rate for unsteady solid body rotation test



### Work-precision diagram:

Work-precision diagram for the unsteady solid body rotation test



## Barotropic instability test

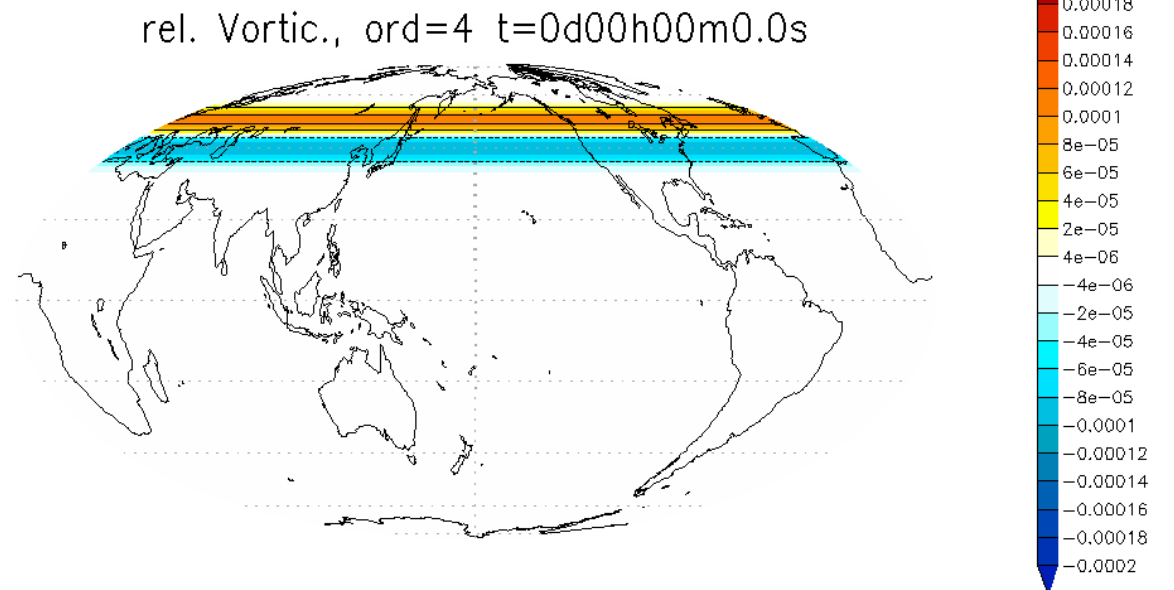
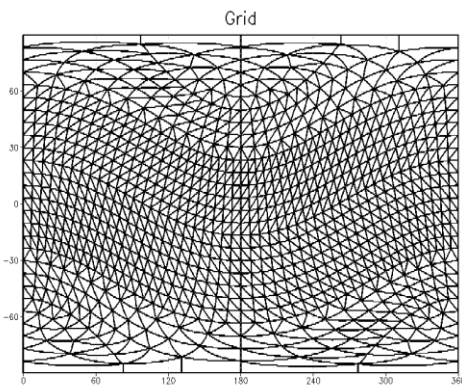
*Galewsky et al. (2004)*

### 4th order DG scheme

without additional diffusion

$dx \sim 67$  km,  $dt = 15$  sec.

simple triangle grid  
on the sphere  
 $dx \sim 500$  km:



relVort:

Mean:  $7.53016e-07$  Min:  $-9.8335e-05$  Max:  $0.000112421$  Sigma: 2

GRADS: COLA/IGES

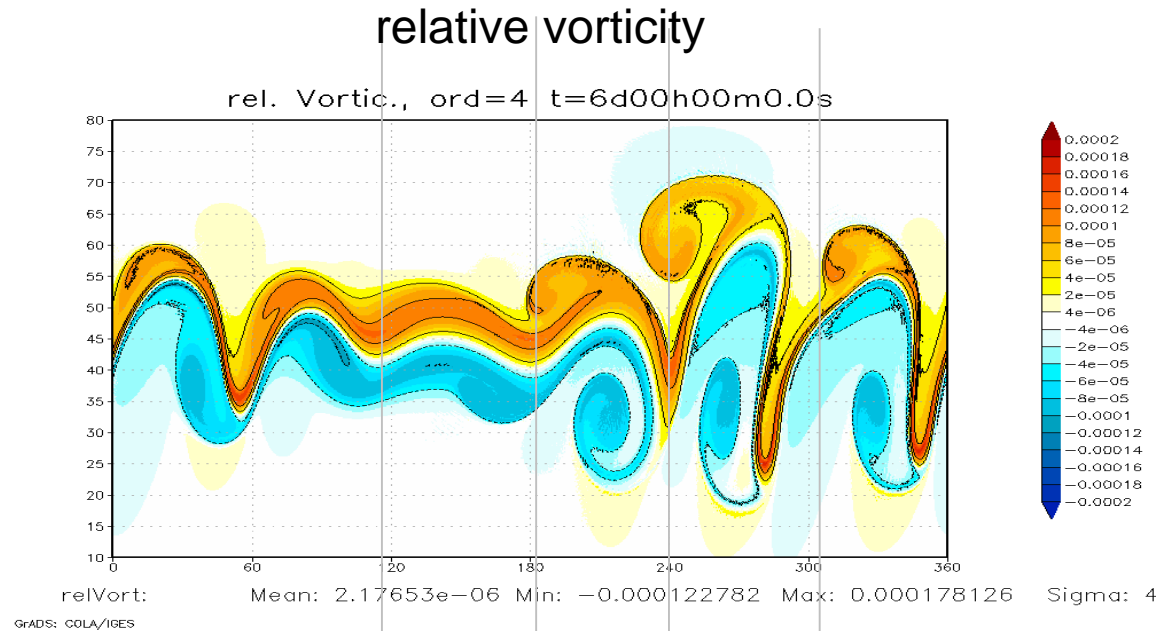
2019-08-22-13:44

GRADS: COLA/IGES

2019-09-03-17:03

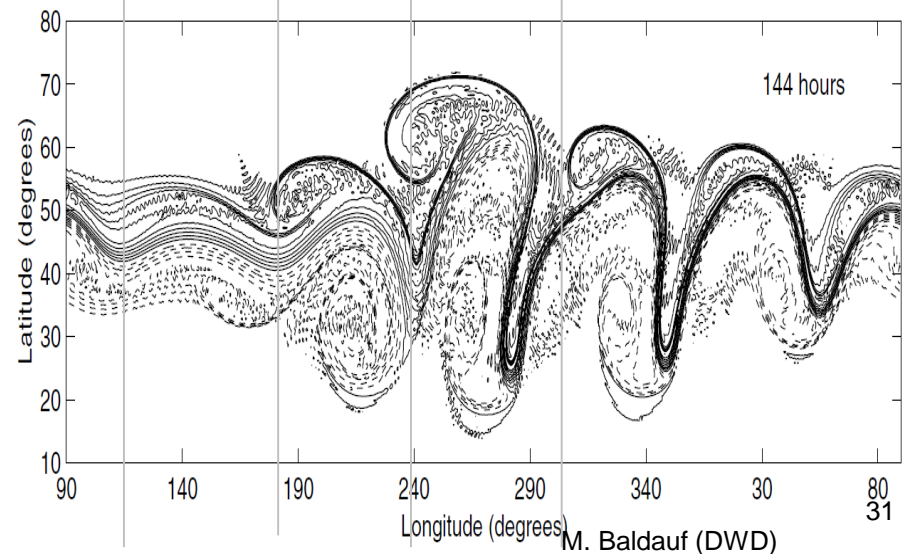
## Barotropic instability test *Galewsky et al. (2004)*

**4th order DG scheme**  
without additional diffusion  
dx~67 km, dt=15 sec.



**FMS-SWM** (Geophys. Fl. Dyn. Lab.)  
without additional diffusion  
dx~60 km (T341), dt=30 sec.

Fig. 4 from *Galewsky et al. (2004)*



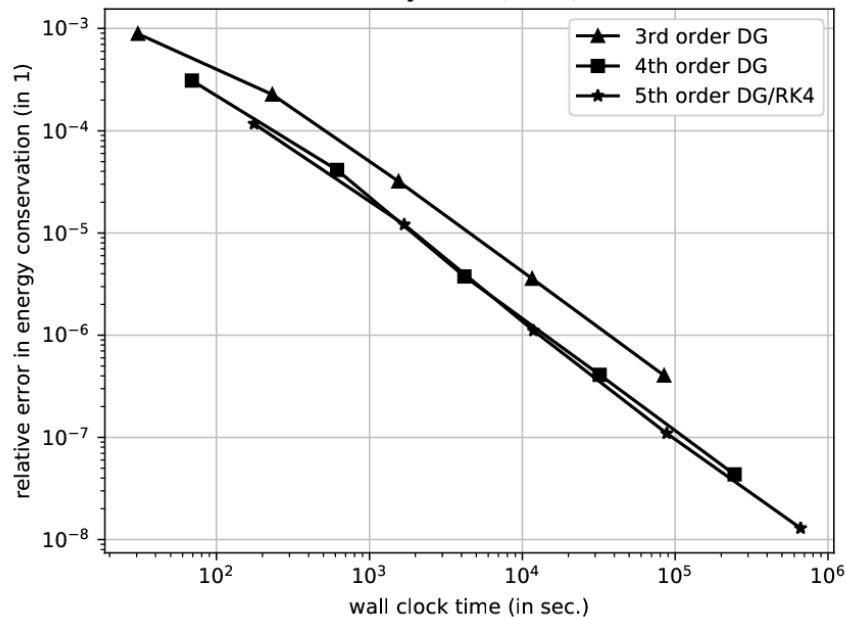
## Barotropic instability test

*Galewsky et al. (2004)*

Work-precision-diagrams

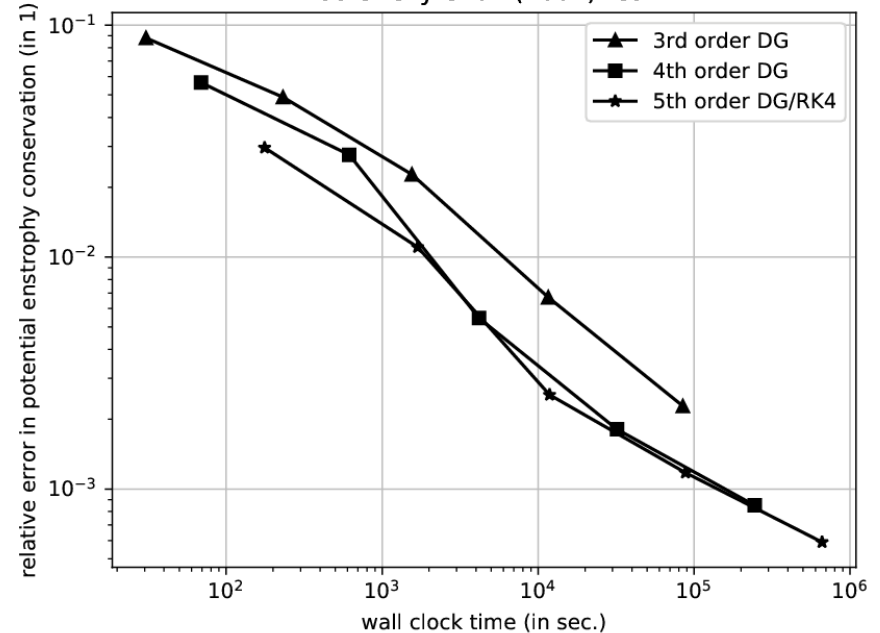
rel. error in energy conservation

Galewsky et al. (2004) test



rel. error in pot. enstrophy conservation

Galewsky et al. (2004) test



## Barotropic instability test *Galewsky et al. (2004)*

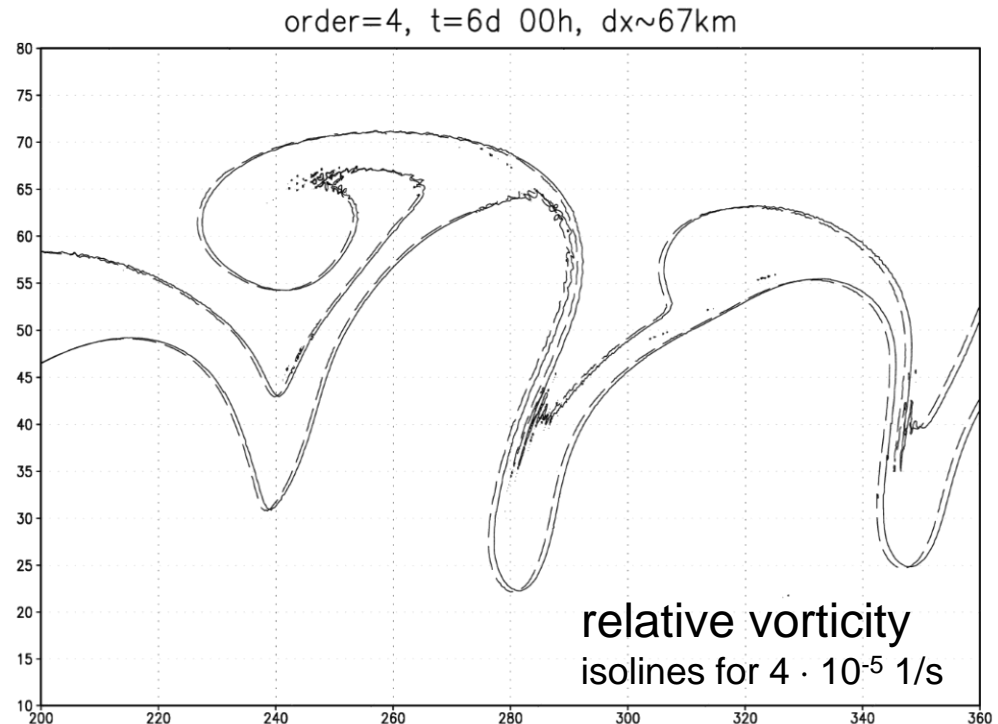
### 4th order DG scheme

without additional diffusion  
 $dx \sim 67$  km,  $dt = 15$  sec.

solid line: sphere  
 $R = 6371.22$  km

dashed line: ellipsoid  
 $a = 6378.137$  km  
 $c = 6356.752$  km  
→ numer. excentr. = 0.082

## Comparison between the sphere and the ellipsoid



- ellipsoidal solution shows westward phase shift of  $\sim 1^\circ$  after 6 days
- is in qualitative agreement with *Bénard (2015) QJRMS*



## Flow over mountain

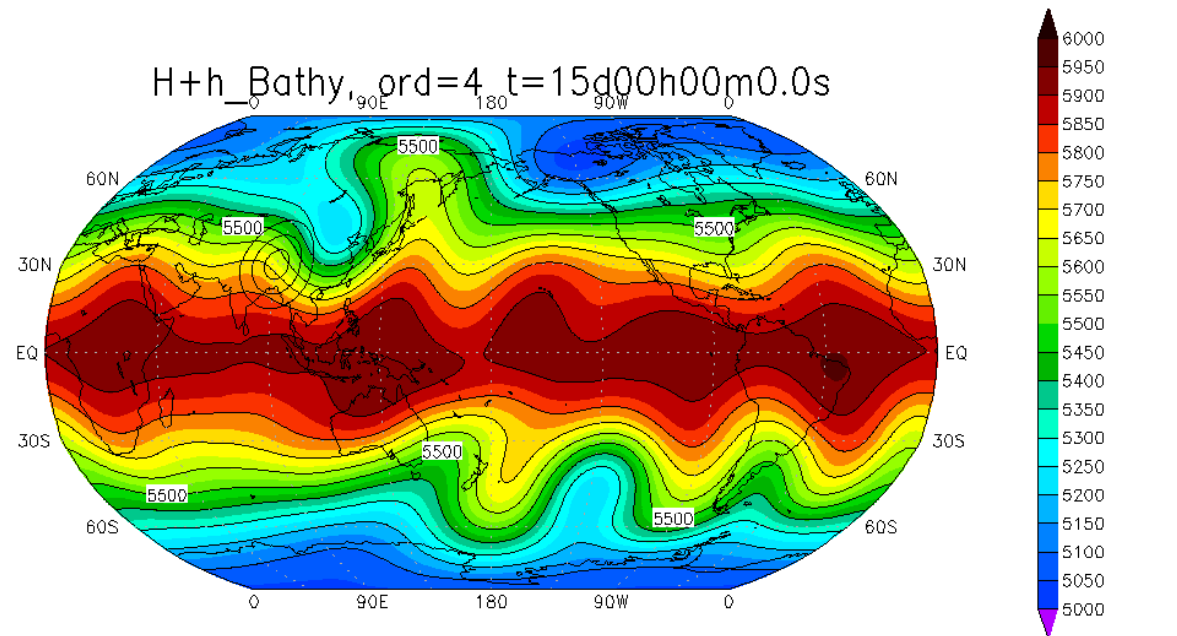
Test case 5 of *Williamson et al. (1992)*

init.: solid body rotation velocity field ( $u_{\max}=20$  m/s) in geostrophic balance with  $H$ ;  
mountain height 2000m

### 4th order DG scheme

without additional diffusion

$dx \sim 67$  km,  $dt = 15$  sec.



H+h\_bathy: Mean: 5501.36 Min: 5032.07 Max: 5953.86 Sigma: 3

GRADS: COLA/IGES

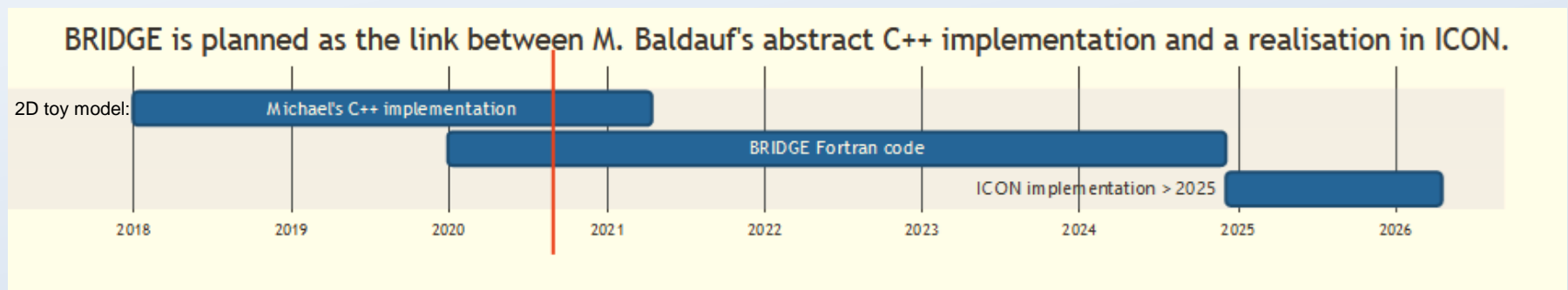
2019-10-10-10:01

# "BRIDGE" - Basic 3D Test Code

3D test code with as little infrastructure overhead as possible

Florian Prill (DWD)

- Discontinuous Galerkin discretizations
- Test platform for DSLs and new infrastructure libraries
- Benchmark "dwarfs"



**DG Working Group** - It's just starting ...

BRIDGE is application-oriented research –  
and it's extra work.

M. Baldauf • F. Prill • D. Reinert • S. Borchert • ...



Image by Brian Marks - IMG\_8527.jpg, CC BY 2.0  
<https://commons.wikimedia.org/w/index.php?curid=60689515>

- Fortran 2003
- MPI parallelization
- Modules and interfaces closely resemble ICON

*Example:*

Source tree: `parallel_infrastructure`, `shared`, `shr_horizontal`, ...

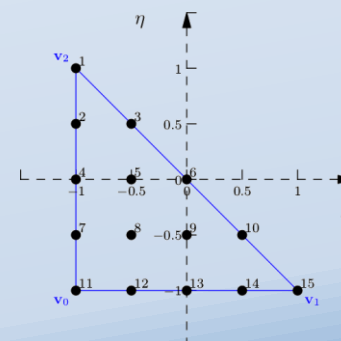
Data types and objects: `wp`, `t_patch`, `t_grid_cells`, `t_var_list`, ...

- ICON triangular grid
- Storage layout: MPI parallel domain decomposition, arrays of 2D triangulations are `nproma`-blocked

**2.5D:** Restrict to quadrature points and node sets which are built from 2Dx1D tensor products.

Consequently, restriction to expansions

$$q(r(x^1, x^2), s(x^3), t) = \sum_{l=1}^M \sum_{m=1}^N Q_{l,m}(t) \phi_l(x^1, x^2) \psi_m(x^3) .$$



- MPI parallelization: no "no-MPI" mode, only worker tasks present (no detached PEs)
- operates on a single domain only
- no "local parent patch" but parent grid may have enlarged halo
- only one precision kind **wp**
- only rudimentary field meta-data
- no restarting capability
- no refinement of cells in a column (eg. small cells at the bottom, coarse at the top)

## Libraries

*... as little infrastructure overhead as possible:*

- Use of supporting libraries like **YAXT**  
The communication library YAXT (DKRZ, Hamburg) abstracts the communication on MPI level from the application.
- better separation between scientific code and infrastructure:  
two auxiliary libraries have been created: **libftnbasic** and **libcbasic**.

## F2003 object-oriented features

The code uses F2003 objects to a far greater extent than ICON:

- to avoid global variables
- to make the data flow in the model more transparent  
*which variable is touched by whom, and when ...*

Vectors (FE coefficients, evaluated shape fct., ...) remain **REAL (wp)** vectors  
= no derived type abstraction.

The user takes care of the index ordering  
or the fact if they are global or local, synched or not.

## DG and Parameterisations

**General principle:** spatial transport (advection, sedimentation, diffusion, ...) must be treated in the DG-scheme!

Otherwise we lose conservation.

**Box-models** (e.g. **cloud physics, chemistry/aerosol-packages**):

are evaluated and deliver tendencies in every quadrature point

→ at the first place no adaptations necessary!

Nevertheless, the 'classical' physics/dynamics-coupling questions remain:  
overall time integration scheme? how to achieve positive definiteness?

### **Turbulence:**

Remark: diffusion needs special treatment in DG (local DG, compact DG, ...)

Advantage of local DG: derivatives of fields are directly available for turbulence modeling!

...

## DG and physics perturbations in ensembles

Some recommendations ...

... to keep conservation properties of the DG scheme:

- in the **transport terms**, only (physical) fluxes should be perturbed.
- in the **source terms**: e.g. moisture var., perturb in a way that  $\rho_{dry} + \rho_v + \rho_r + \dots + \rho_g$  is unchanged (while keeping positive def.)

## DG and data assimilation

At least adaptations in the forward operators necessary:

- by the modified output grid; better say: to the position of the I/O-grid points (these are probably the quadrature points in the triangle grid)
- different prognostic variables (conserved var.)

## Summary

- **2D toy model** for
    - explicit DG-RK (unstructured grids, triangle or quadrilateral grid cells) and
    - **HEVI DG-IMEX-RK**now works for *Euler equations* with *terrain-following coordinates* and optionally with *3D diffusion* (explicit or HEVI): several tests show correct convergence behaviour, well-balancing problems solved, ...
  - Efficiency problems with band diagonal matrix solver strongly improved: the whole implicit part (build coeff. matrix, *LU decomposition*, matrix-vector-mult.) takes ~60% of total run time.
  - **DG on the sphere** on a triangle grid possible by the use of local (rotated gnomonial) coordinates and the covariant formulation of the equations.  
(Baldauf (2020) JCP)
- With respect to the pure dynamical core (=solver for the Euler equations), no show-stopper occurred until now  
However, total efficiency is still an issue!



## Outlook

- Current tasks in the **2D toy model**
  - Consolidation of what has been achieved so far:  
further testing; what are the true limits of the method?

Further **milestones** (for the next years!)

- Development of a **3D prototype**, DG-HEVI on the sphere  
**BRIDGE** (**B**asic **R**esearch for **I**CON with **DG** **E**xtension) (start ~mid 2020)
  - **Further design decisions:** nodal vs. modal?,  
local DG vs. interior penalty vs. ...?, allow non-conformal grids?,  
efficient data layout, ...
  - Coupling of **tracer advection** (mass-consistency)?
  - Develop coupling ideas for parameterizations  
time-integration, preserve pos. def., ...
- Implementation into **ICON** (start ~2024)
  - choose optimal convergence order  $p$  and grid spacing  
estimated:  $p_{\text{horiz}} \sim 4 \dots 6$ ,  $p_{\text{vert}} \sim 3 \dots 4$  ( $p_{\text{time}} \sim 3 \dots 4$ )  
currently I favor:  $p_{\text{horiz}}=4$ ,  $p_{\text{vert}}=4$ ,  $p_{\text{time}}=3$

Thank you very much for your attention!