

# Radio Frequency Interference in the SMAP Radiometer

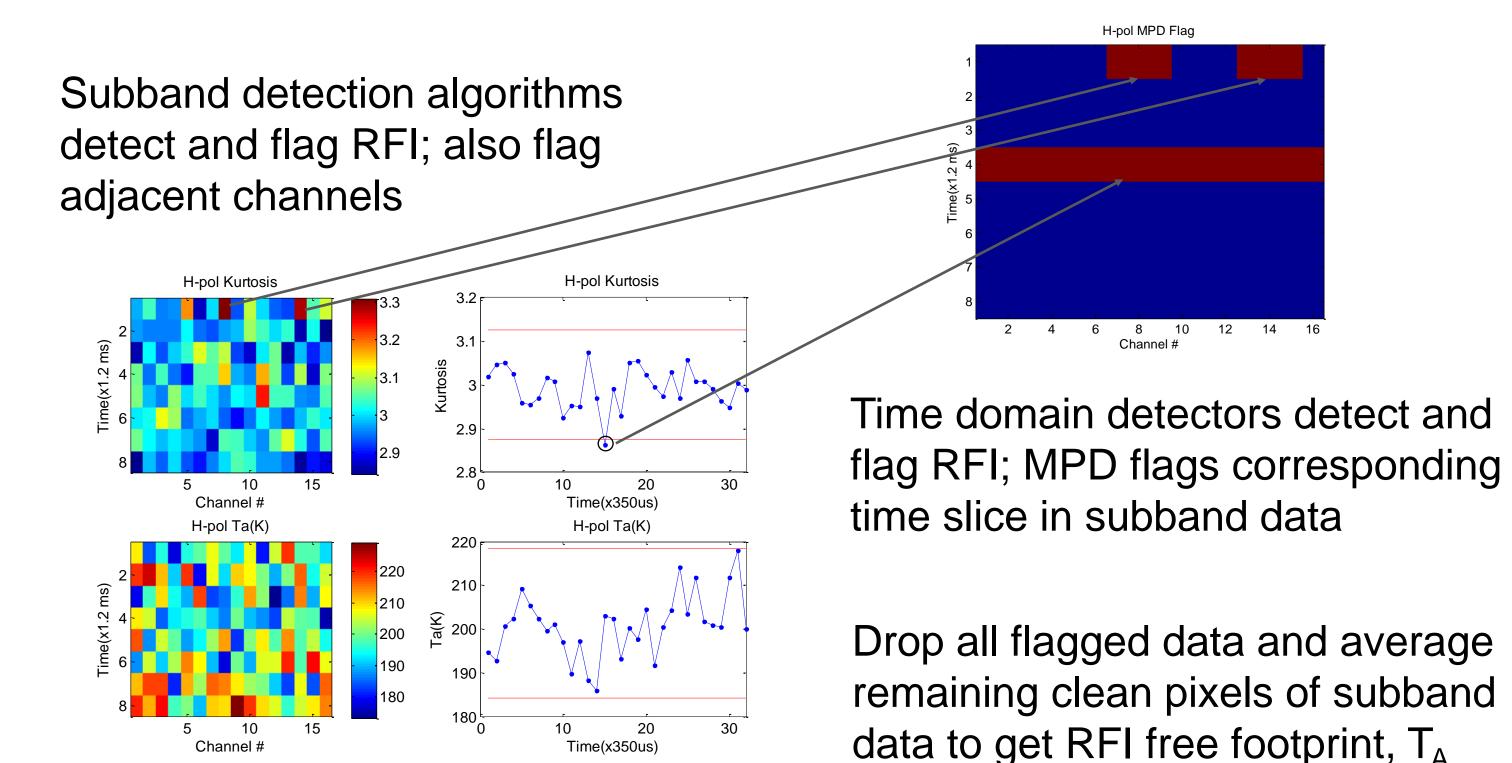
Priscilla N. Mohammed<sup>1, 2</sup>, Alexandra Bringer<sup>3</sup>, Sidharth Misra<sup>4</sup>, Paolo de Matthaeis<sup>1, 5</sup> Joel T. Johnson<sup>3</sup>, Jeffrey R. Piepmeier<sup>1</sup>, Melanie Brunner<sup>6</sup>

<sup>1</sup>NASA's Goddard Space Flight Center, <sup>2</sup>Morgan State University, Baltimore, MD 21251 USA,, Greenbelt, MD 20771 USA, <sup>3</sup>The Ohio State University, Columbus OH 43210, Jet Propulsion Laboratory, California Institute of Technology, Pasadena, CA 91106 USA, <sup>5</sup>University of Maryland Baltimore County, Baltimore MD 21250 USA, <sup>6</sup>NASA Glenn Research Center, Cleveland OH 44135, USA

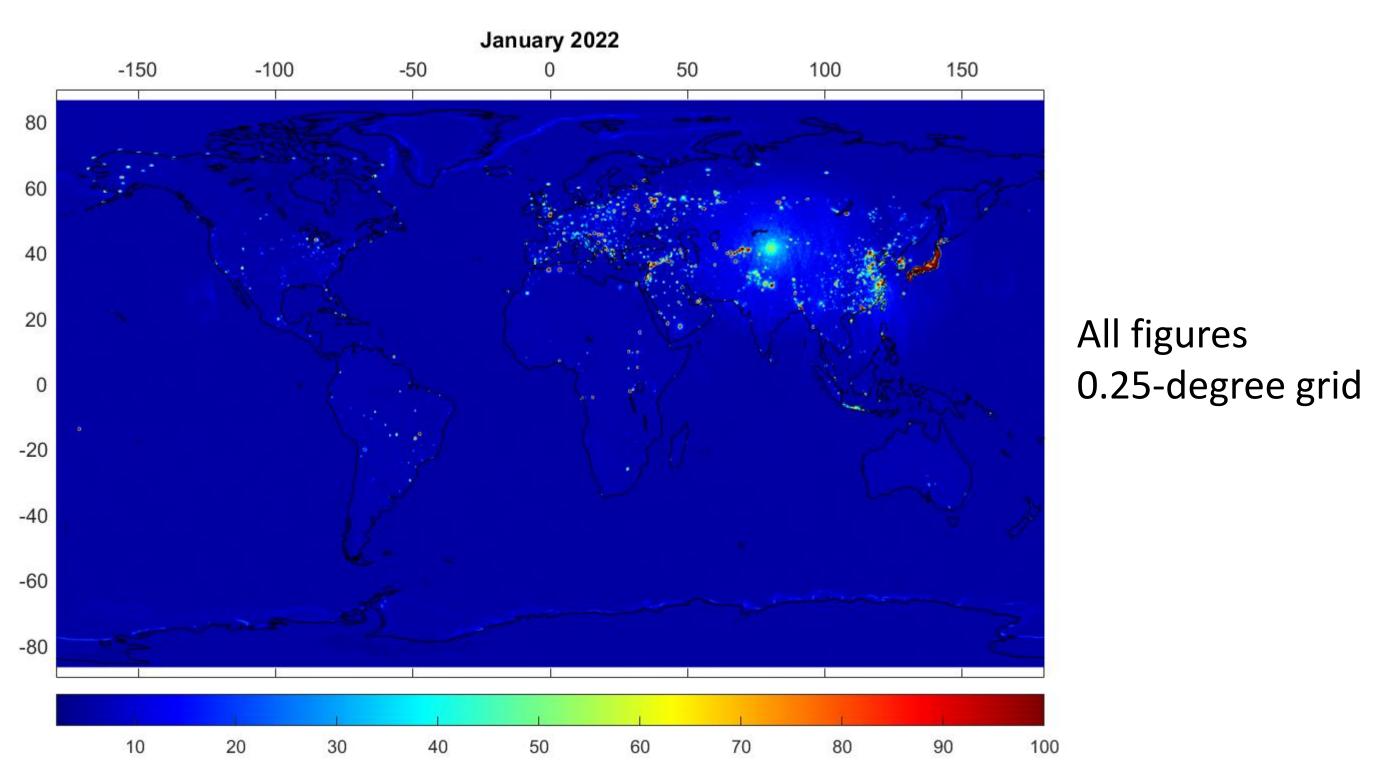
#### Introduction

- SMAP (Soil Moisture Active Passive) was launched by NASA January 31, 2015, to measure soil moisture of the Earth's land surface
- The SMAP radiometer operates in the L-band protected spectrum (1400-1427 MHz) that is known to be vulnerable to radio frequency interference (RFI)
- SMOS and Aquarius provided a good indication of the RFI environment at L-band
- On orbit results show that RFI is indeed a problem
- RFI increases brightness temperatures
   Can lead to dry biases in soil moisture retrieval
- Can lead to dry biases in soil moisture retrievals if undetected
- SMAP radiometer includes a digital backend enabling multiple RFI detection and mitigation capabilities; detection and mitigation processing performed on ground

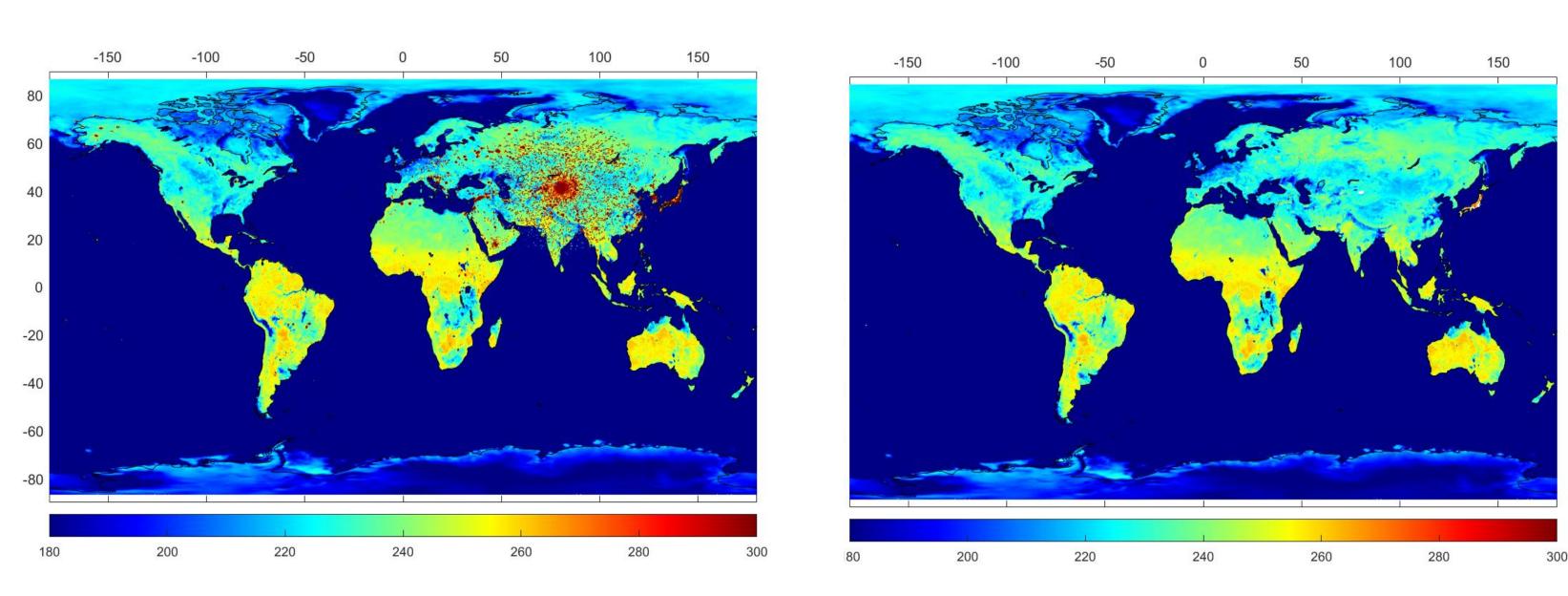
## SMAP RFI Detection and Filtering



## **SMAP RFI Detection**

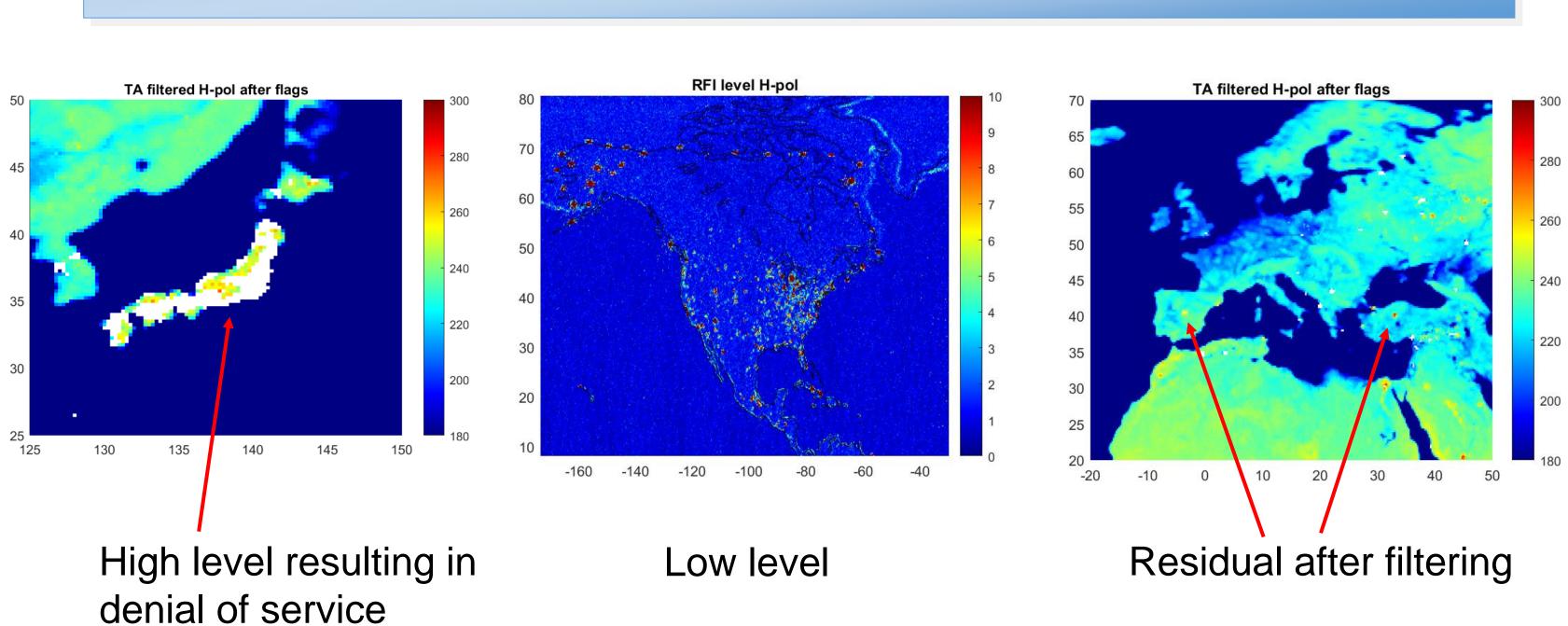


**SMAP RFI Detection Rate** 



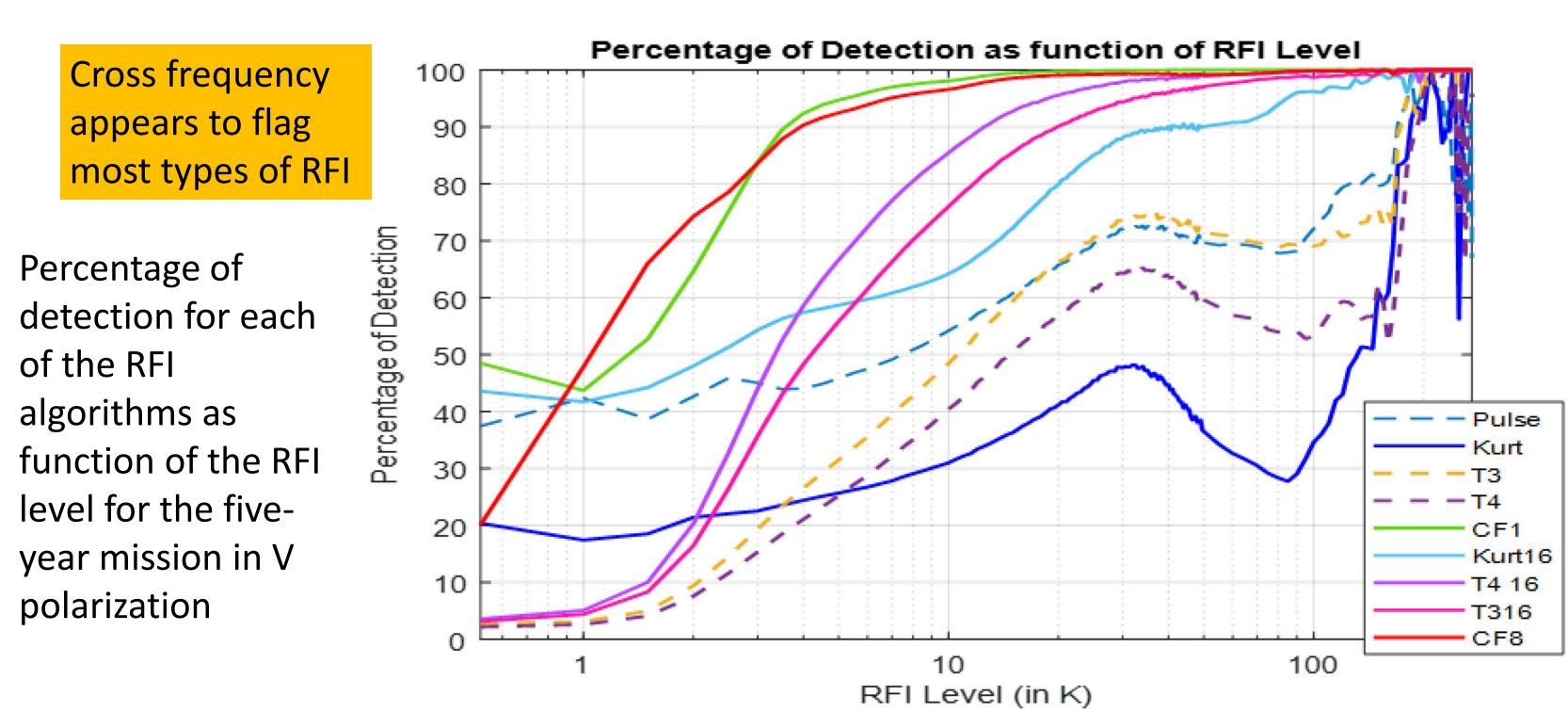
**TA H-pol before and after filtering**Peak hold data 02/02/2022 to 02/08/2022

# Different Types of RFI



#### Peak hold data 02/02/2022 to 02/08/2022

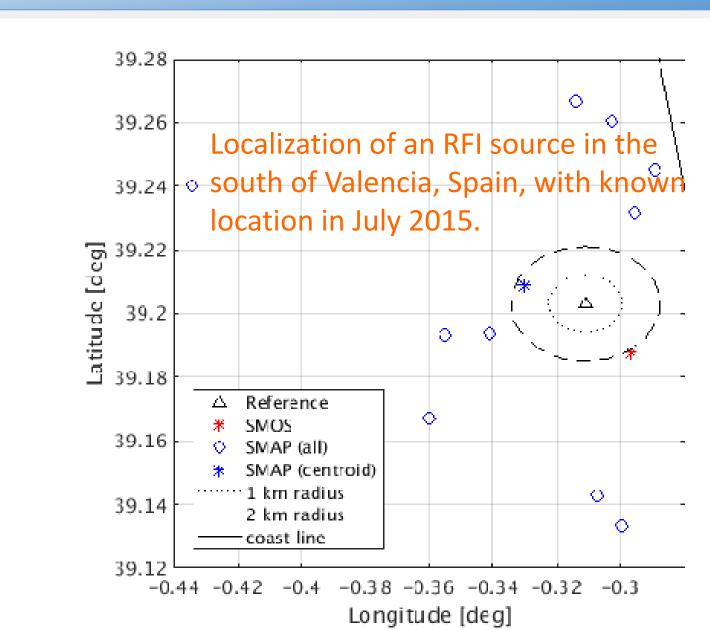
### **SMAP RFI Detectors**



A. Bringer *et al.*, "Properties of the RFI Environment at 1400–1427 MHz as Observed by the Soil Moisture Active/Passive Mission Microwave Radiometer," in *IEEE Journal of Selected Topics in Applied Earth Observations and Remote Sensing*, vol. 14, pp. 7259-7267, 2021, doi: 10.1109/JSTARS.2021.3092996.

## Locating Sources with SMAP Data

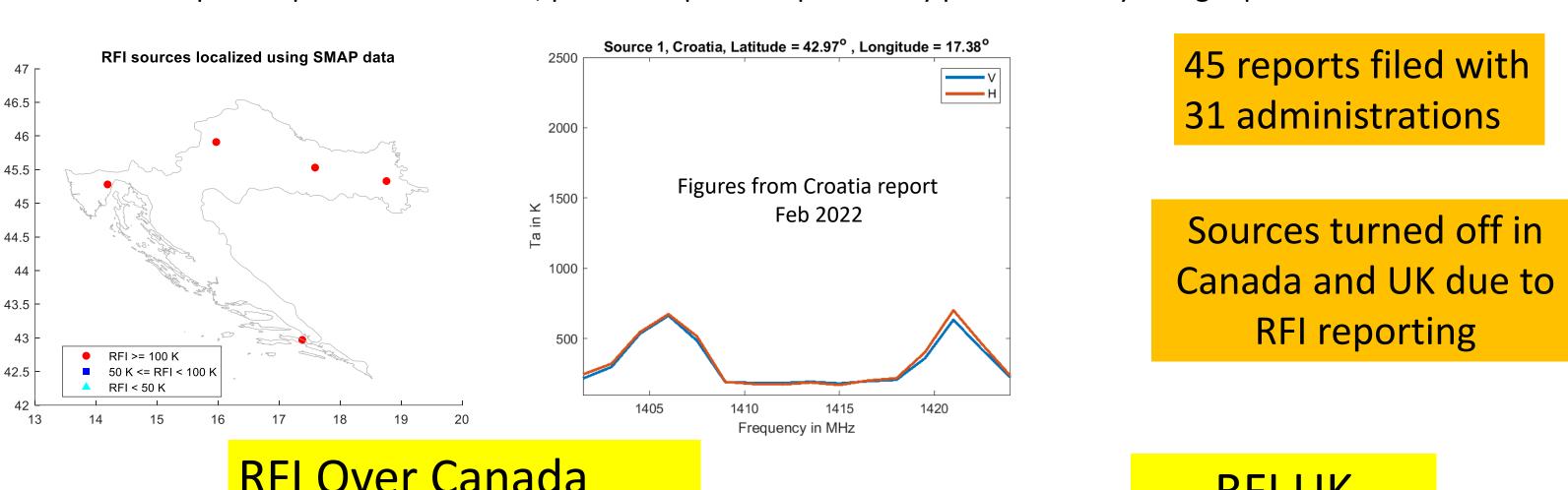
- Each blue diamond is found by clustering RFI footprints from a single pass (half orbit)
- Centroid (blue star) is found using all the clusters from several passes
- clusters from several passes
   The black triangle is the known source location
- Algorithm used to find sources and geolocated sources are included in RFI reports filed
- Y. Soldo *et al.*, "Location of Radio-Frequency Interference Sources Using the SMAP L-Band Radiometer," in *IEEE Transactions on Geoscience and Remote Sensing*, vol. 56, no. 11, pp. 6854-6866, Nov. 2018, doi: 10.1109/TGRS.2018.2844127.

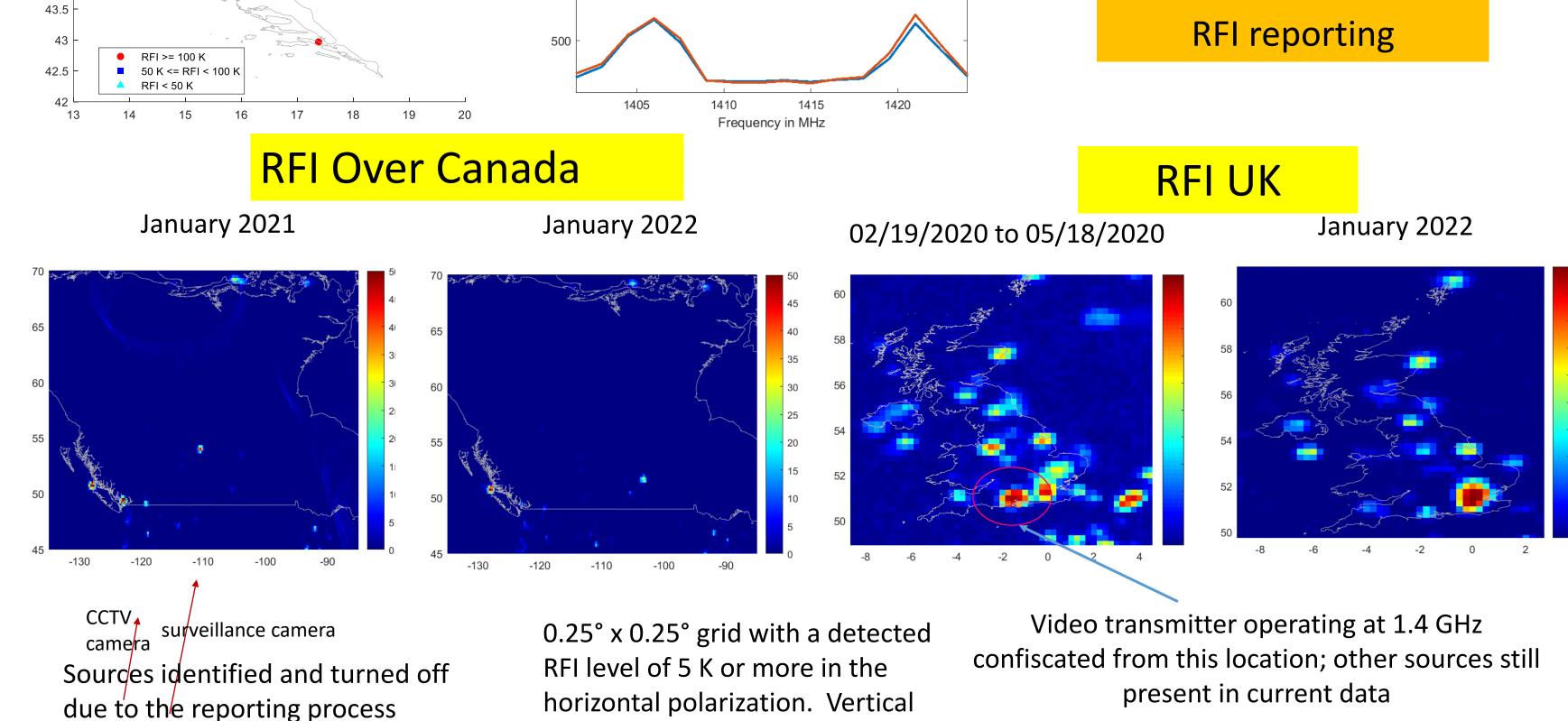


**Probability Plot** 

## RFI Reporting

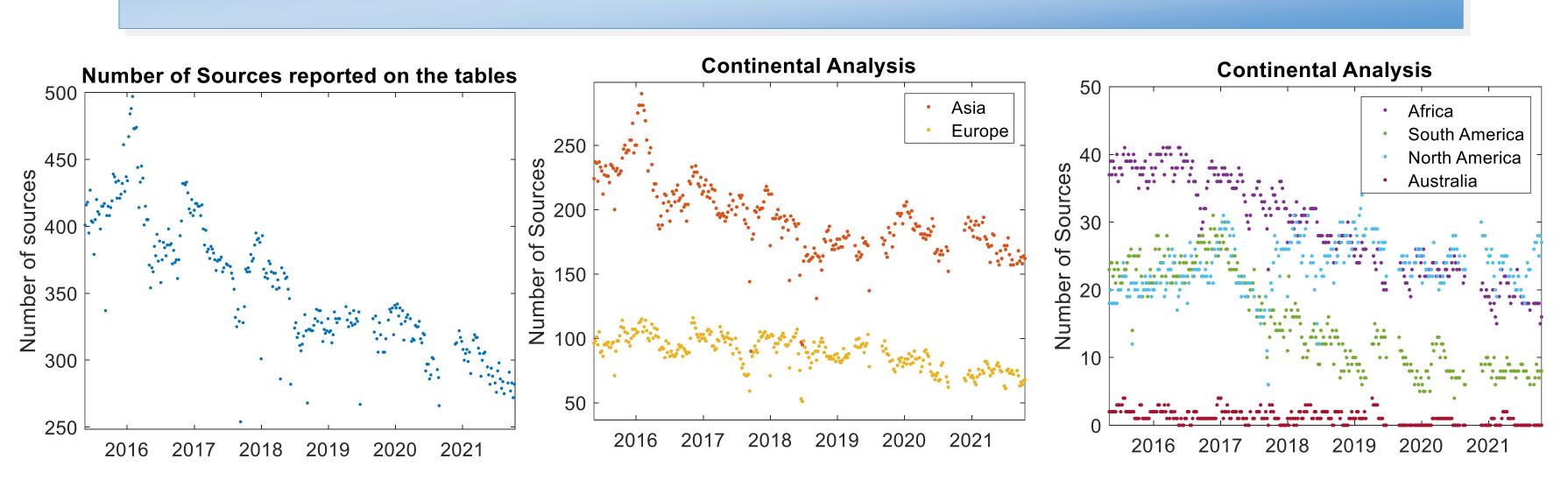
- RFI reported to authorities through NASA spectrum office
- Report a different country/region every month
- SMAP and SMOS agree to report on the same country every month
- Report for each source contains
  - Location coordinate
    Brightness temperature in K
  - Brightness temperature in KEstimate of EIRP of transmitting source
  - Estimate of EIRP of transmitting source
    # of observations of source over analysis period
- # of observations of source
   Date source was last seen
- Spectral plot of each source, peak hold plot and probability plot of country being reported





#### **SMAP RFI Environment**

polarization shows similar results.



•Because the number of RFI sources are large, 2 criteria were defined to select the sources that have been reported:

- o RFI level > 10 K
- o Persistent in time i.e., present in at least 25% of SMAP overpasses during a month
- •A table is generating every week using the information of the 4 previous weeks:
- Over the 6 years of SMAP missions, 300 tables were generated
  Their analysis allow to track RFI changes temporally and spatially