### TRAINING COURSE

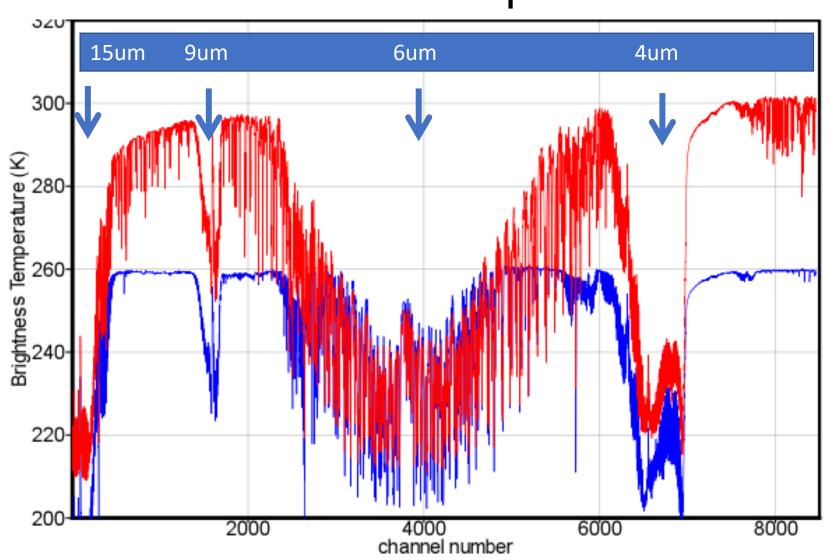
EUMETSAT/
ECMWF
NWP-SAF
satellite data
assimilation



# ECMWF/EUMETSAT NWP-SAF Satellite data assimilation Training Course

The infrared spectrum, measurement and information content

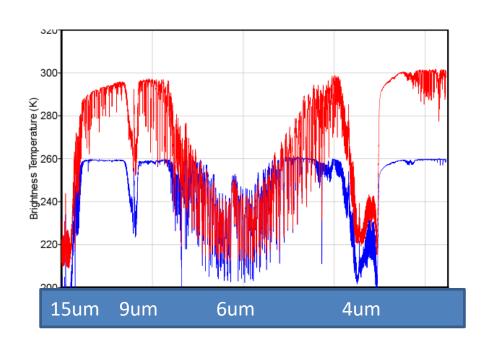
## The IR spectrum in a Tropical and Polar atmosphere

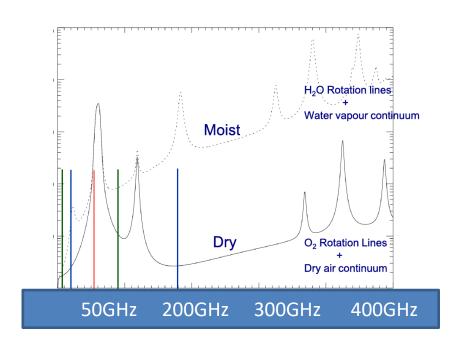


#### Why infrared ...?

....high spectral resolution

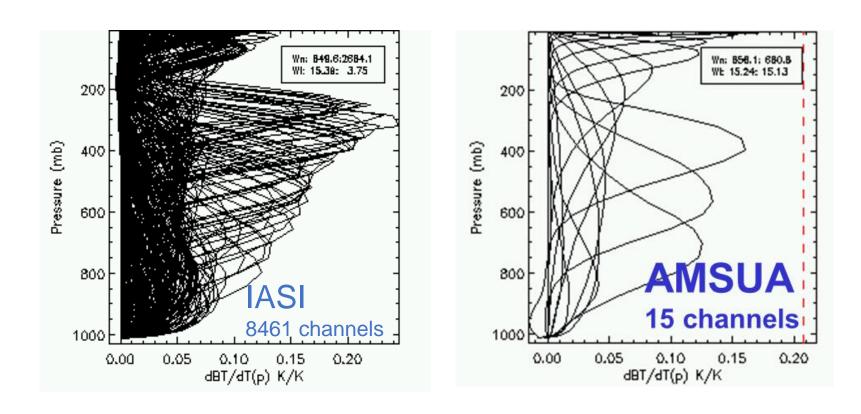
#### **Infrared v Microwave**





In the infrared, the many thousands of distinct resolvable spectral lines allows us to measure atmospheric radiation many thousands of channels – in the microwave there are only a handful on spectral lines!

#### Infrared v Microwave



Each channel has a slightly different weighting function – providing information on a slightly different part of the atmosphere

#### Why high spectral resolution..?

...high vertical resolution

#### ...a helpful linear analogue ...

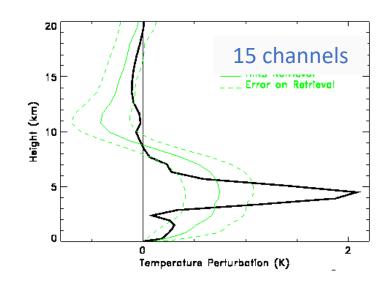
It can be shown that the state that minimizes the cost function is equivalent to a linear correction of the background using the observations:

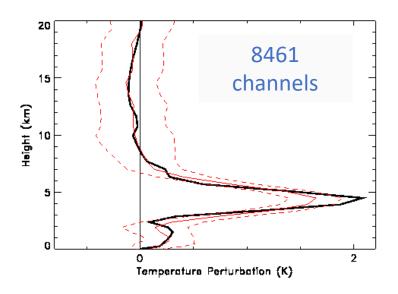
$$x_a = x_b + [\mathbf{H}\mathbf{B}]^T [\mathbf{H}\mathbf{B}\mathbf{H}^T + \mathbf{R}]^{-1} (y - \mathbf{H}x_b)$$
correction term

...and the improvement can be quantified in terms of the key parameters of the assimilation...(i.e. B, R, H)

$$S_a = B - [HB]^T [HBH^T + R]^{-1} HB$$
improvement term

#### **Infrared v Microwave**

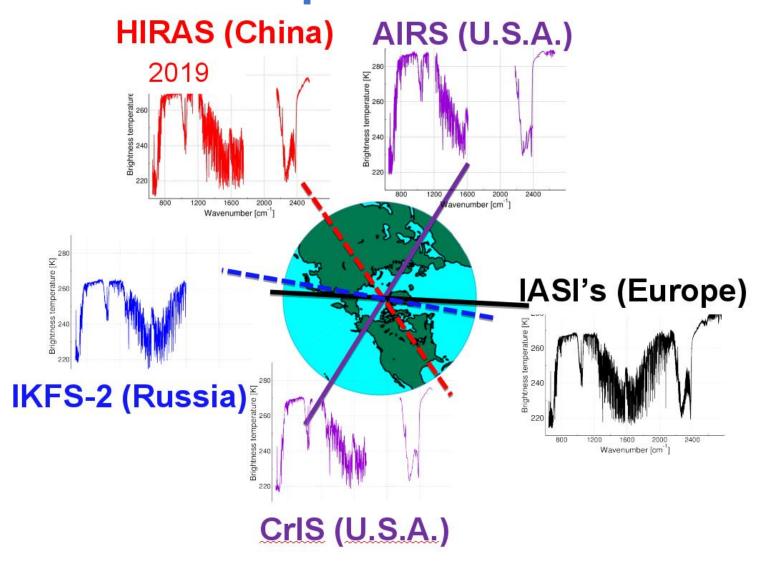




With only a few channels fine vertical feature in the atmosphere are not resolvable – but become more visible with more channels

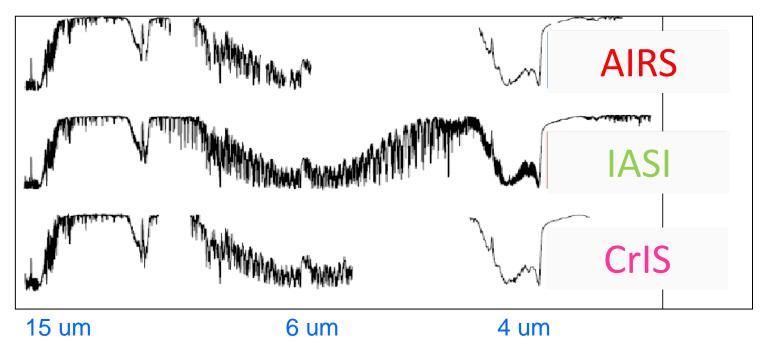
# High Spectral Resolution IR sounders on Polar Spacecraft

### High Spectral Resolution IR sounders on Polar Spacecraft



#### **Operational High Spectral Resolution IR sounders**

| Instrument/ Satellite/ | No. of Channel | Spectral<br>Range        | Spectral Res.                 | IFOV       | Type/<br>Orbit                    |
|------------------------|----------------|--------------------------|-------------------------------|------------|-----------------------------------|
| AIRS/<br>Aqua(EOS-PM)/ | 2378           | 650-2760cm <sup>-1</sup> | ~1cm <sup>-1</sup>            | 13.5k<br>m | Grating<br>Spectrometer/<br>Polar |
| IASI/<br>MetOp/        | 8461           | 645-2760cm <sup>-1</sup> | 0.5cm <sup>-1</sup>           | 12km       | Interferometer<br>/Polar          |
| CrIS/<br>NPP & JPSS/   | 1400           | 635-2450cm <sup>-1</sup> | 1.125-<br>4.5cm <sup>-1</sup> | 12km       | Interferometer<br>/Polar          |



#### IASI v CrIS

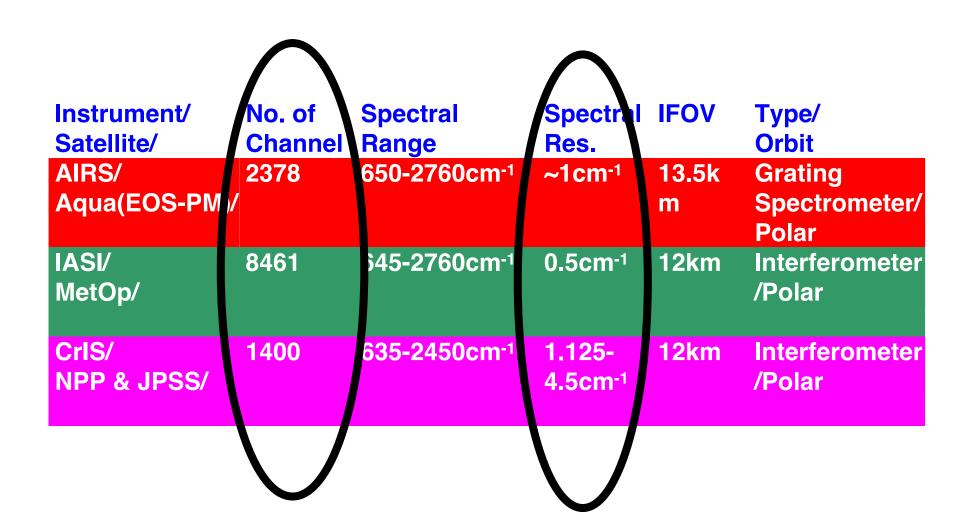
# Cross-track Infrared Sounder (CrIS)

#### CrIS ATMS ±50° Cross track Scans Cris Swath 2200km ATMS Swath 2500km 3x3 Array of CrIS FOVs (Each at 14-km Diameter)

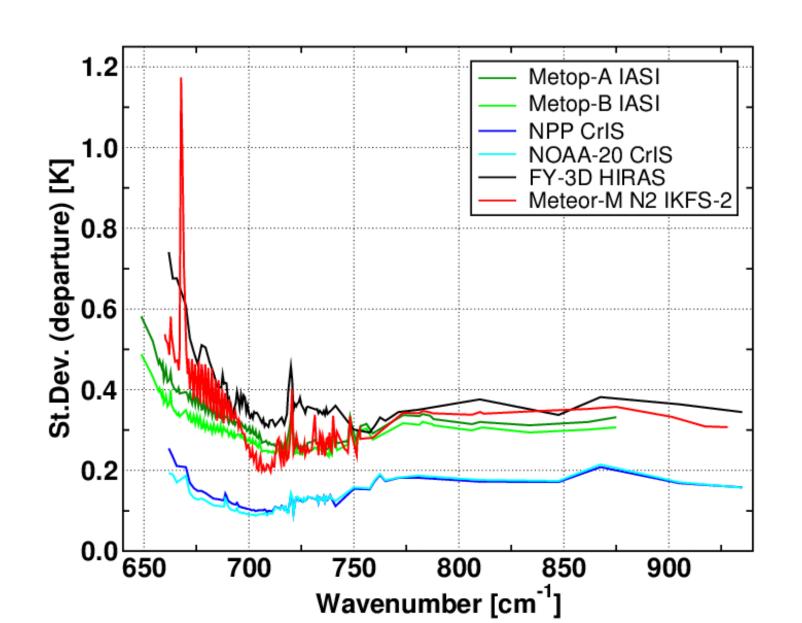
# Infrared Atmospheric Sounding Interferometer (IASI)



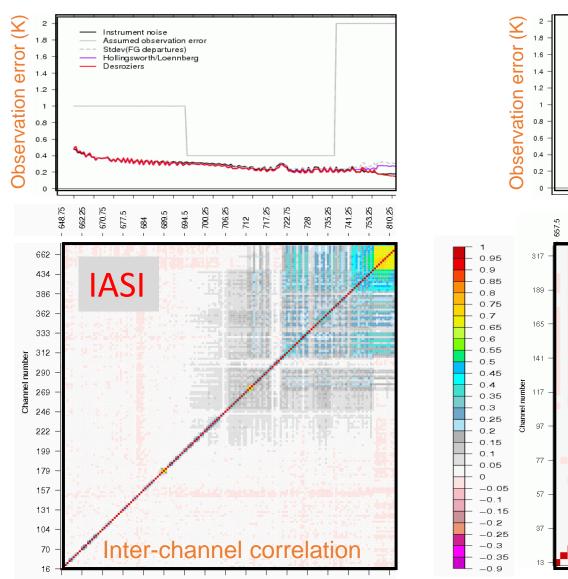
### IASI has higher spectral resolution compared to CrIS

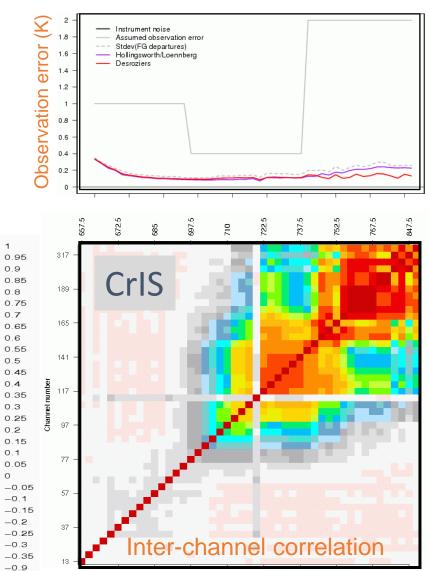


#### IASI has higher noise compared to CrIS



### CrIS has stronger inter-channel correlations than IASI

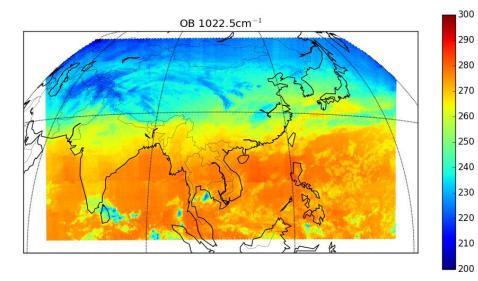




# High Spectral Resolution IR sounders on GEO Spacecraft

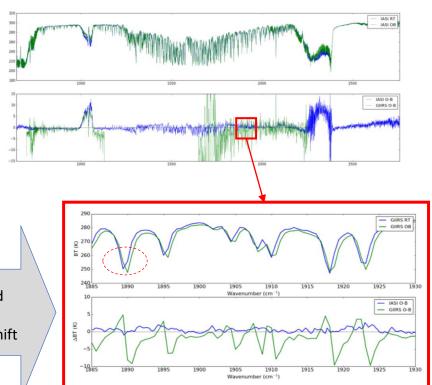
#### First ever GEO High Spectral Resolution IR sounder

FY- 4A GIIRS radiance observations 20190301



Cross checking with model simulations and IASI suggests much of the noise can be explained (and removed) with a spectral shift

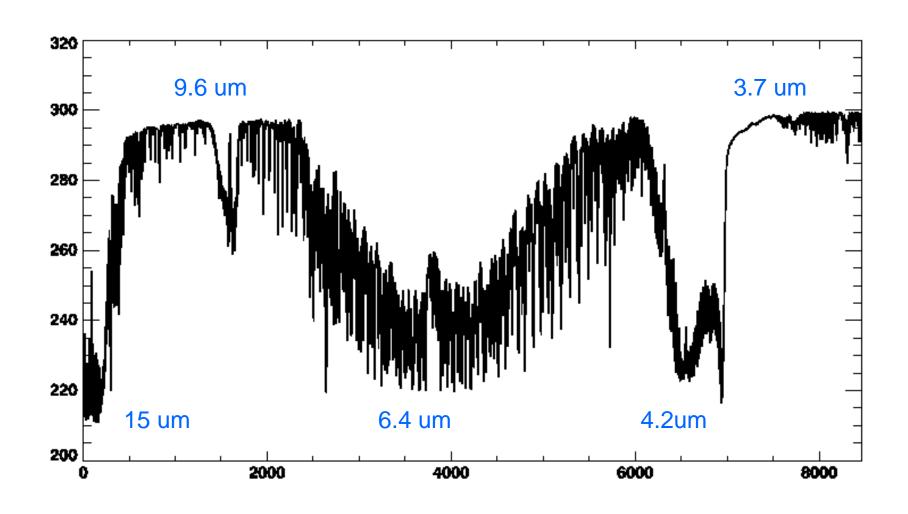
Initial evaluations suggested the GIIRS radiances were significantly more noisy than similar IASI channels



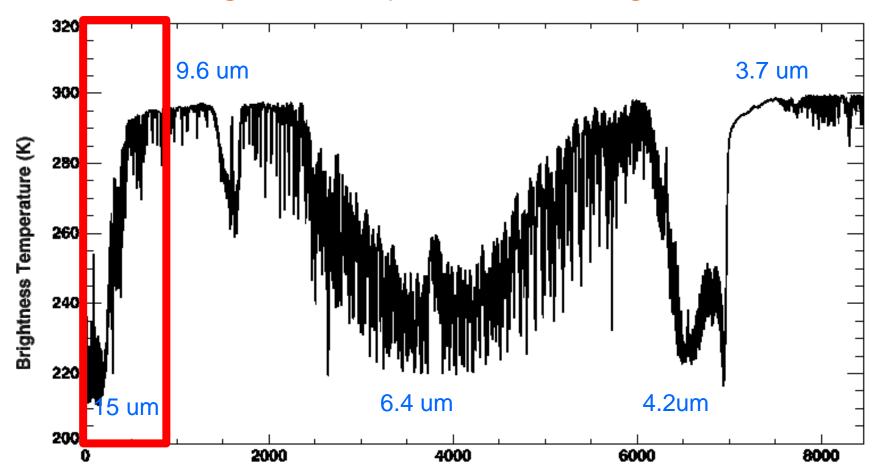
# Case study: The assimilation of IASI

#### What does IASI measure

(and what do we assimilate)

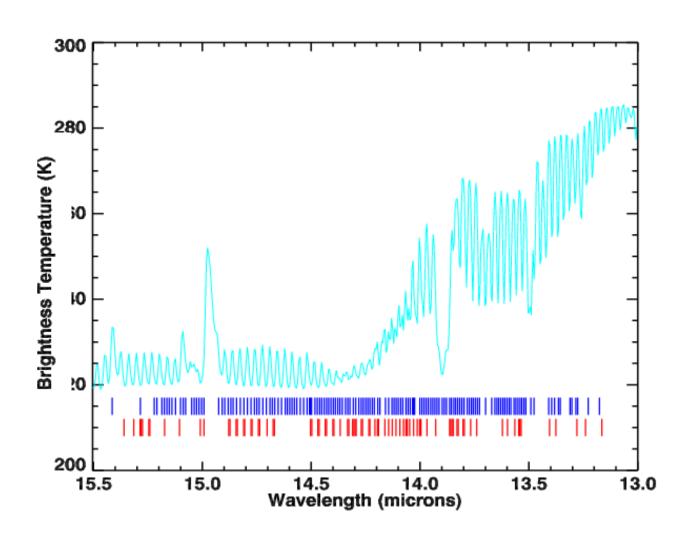


The long-wave temperature sounding band

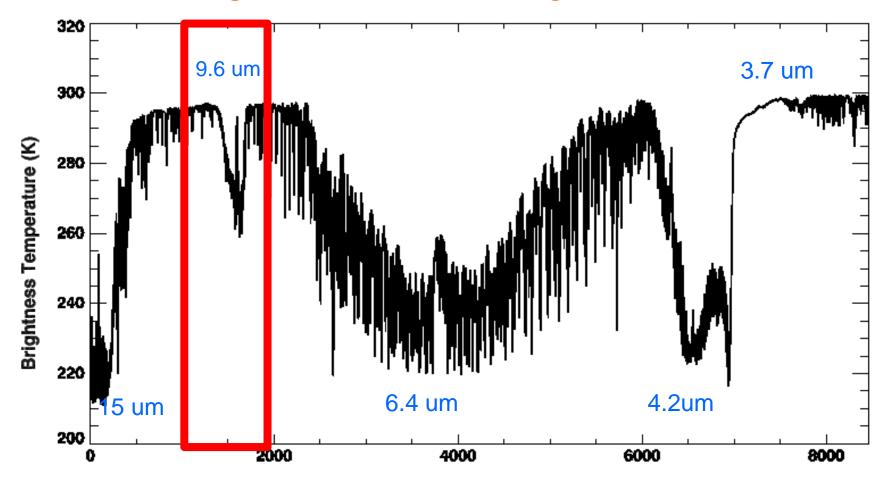


~ 200 channels assimilated

### Zoom of long-wave temperature sounding channel usage for IASI

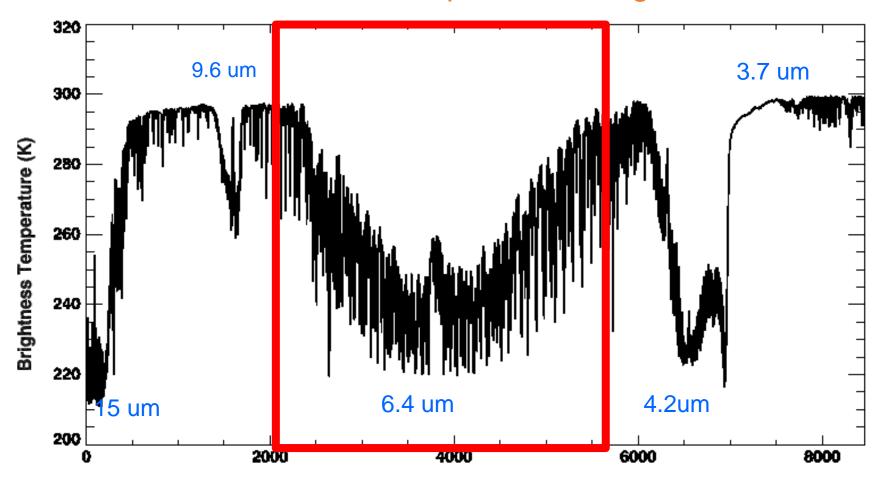


The long-wave ozone sounding band



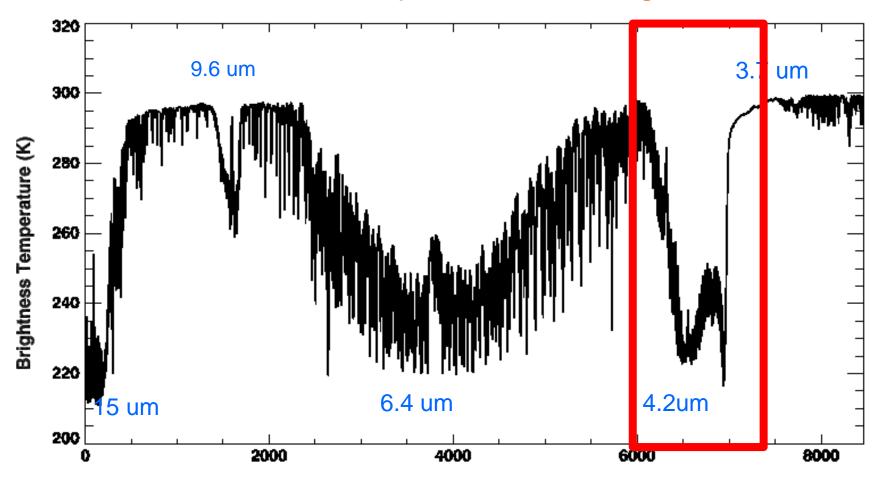
~ 20 channels assimilated

The mid-wave water vapour sounding band



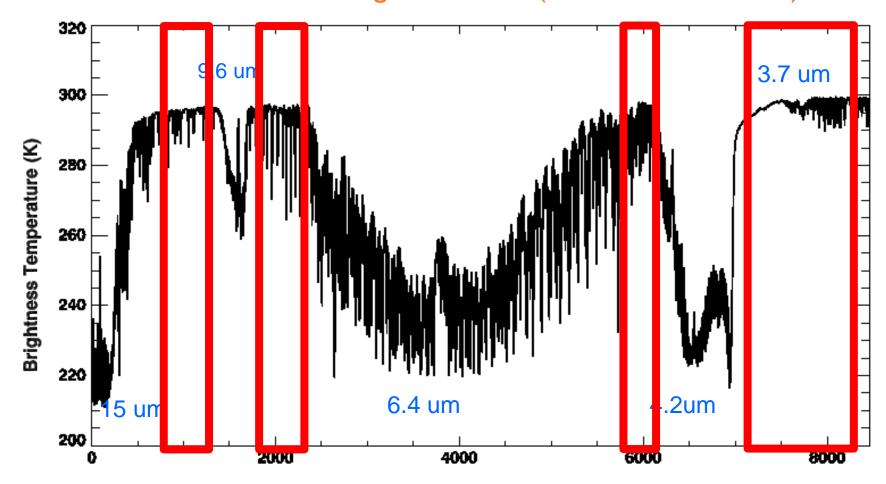
~ 50 channels assimilated

The short-wave temperature sounding band



~ 0 channels assimilated – IASI noise high

The surface sensing channels (window channels)



channels mostly used for cloud detection

## Weighting functions of IASI sounding channels

# Atmospheric sounding channels

...selecting channels where there is **no** contribution from the **surface**....

$$L(\nu) = \int_0^\infty B(\nu, T(z)) \left[ \frac{d\tau(\nu)}{dz} \right] dz + \frac{\text{Surface}}{\text{errorsion}} + \frac{\text{Surface}}{\text{scattering}} + \frac{\text{Cloutzain}}{\text{contrivision}} + \dots$$

#### ATMOSPHERIC TEMPERATURE SOUNDING

If radiation is selected in an atmospheric sounding channel for which

$$L(v) = \int_0^\infty B(v, T(z)) \left[ \frac{d\tau(v)}{dz} \right] dz$$

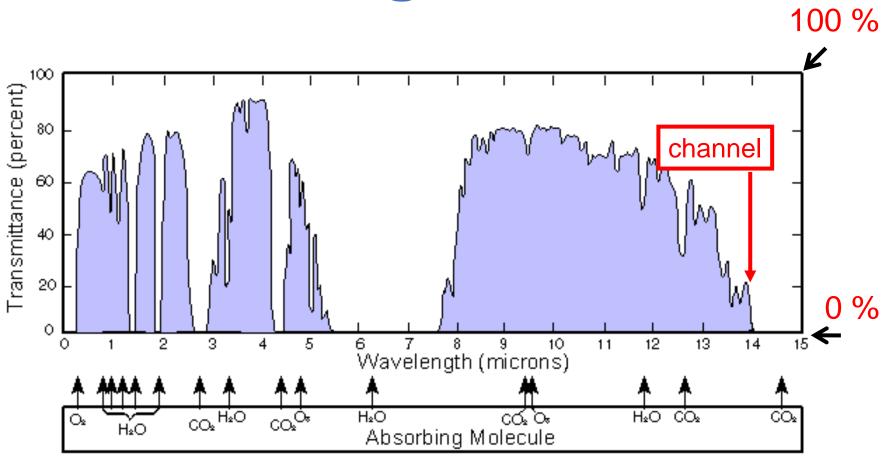
and we define a function 
$$H(z) = \left[\frac{d\tau}{dz}\right]$$

When the primary absorber is a well mixed gas (e.g. oxygen or CO<sub>2</sub>) with known concentration it can be seen that the measured radiance is essentially a weighted average of the atmospheric temperature profile, or

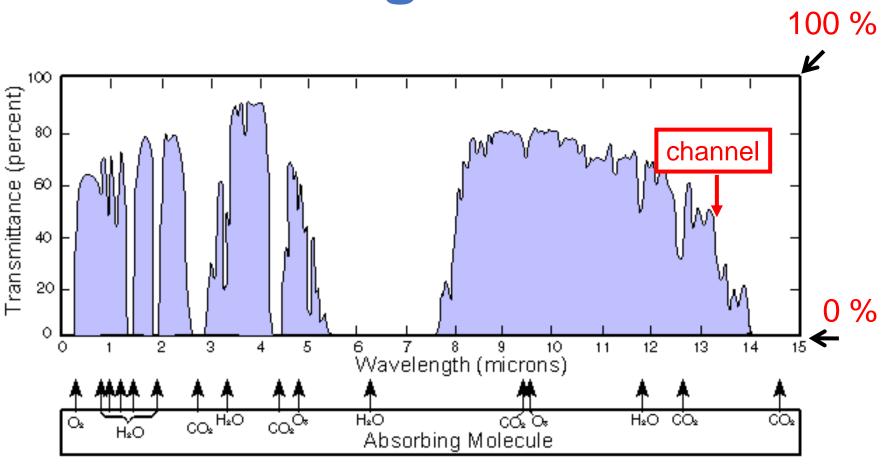
$$L(v) = \int_0^\infty B(v, T(z)) H(z) dz$$

The function H(z) that defines this vertical average is known as a WEIGHTING FUNCTION

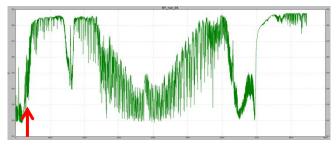
# Strong absorbing sounding channels



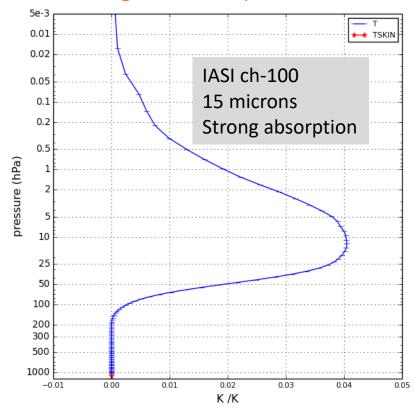
# Weaker absorbing sounding channels

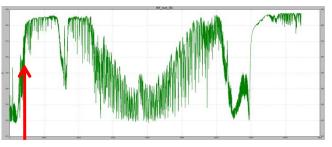


### IASI weighting functions for strong and weak absorption channels

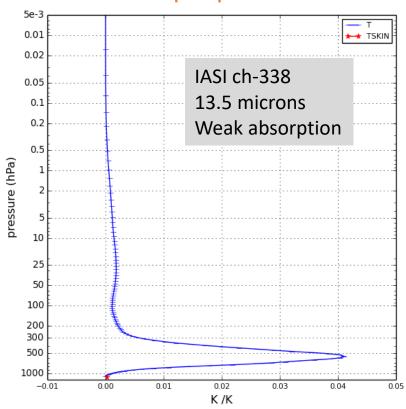


Strong > stratospheric channel



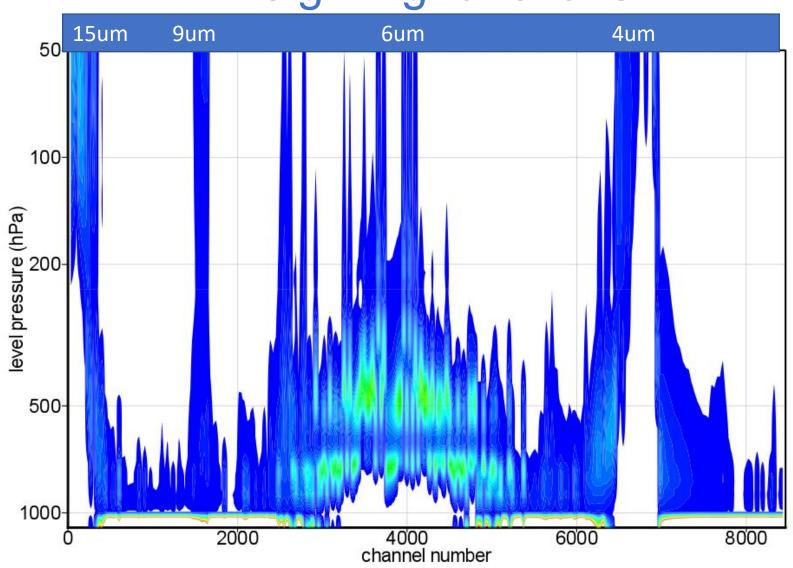


Weak > tropospheric channel

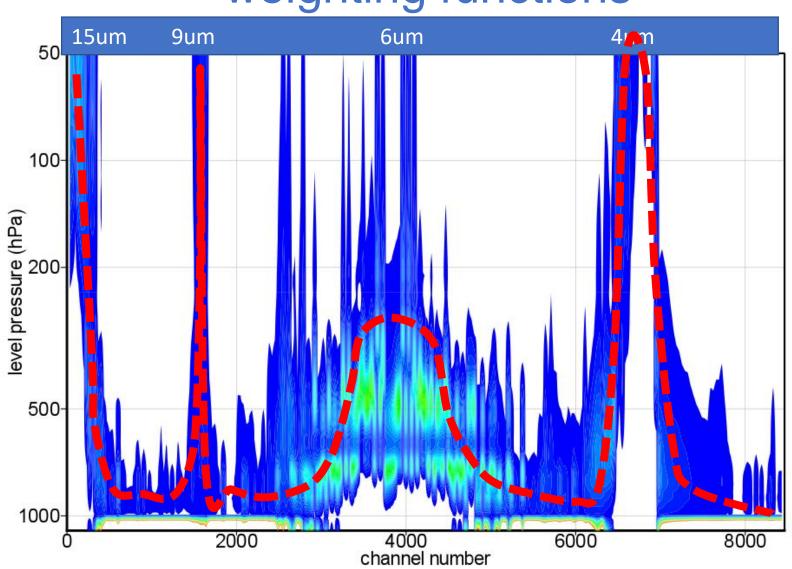


Sampling lines of varying absorption strength IASI provides good vertical coverage of the atmosphere

## Peaks altitudes of IASI channel weighting functions



## Peaks altitudes of IASI channel weighting functions



## Three challenges to the successful assimilation of infrared radiances:

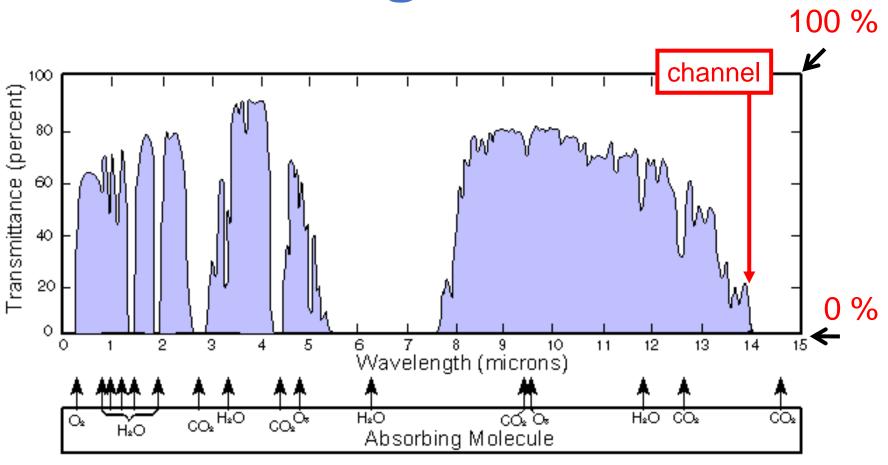
- 1) Sounding the lower atmosphere
- 2) Variable absorbing gasses
- 3) Clouds

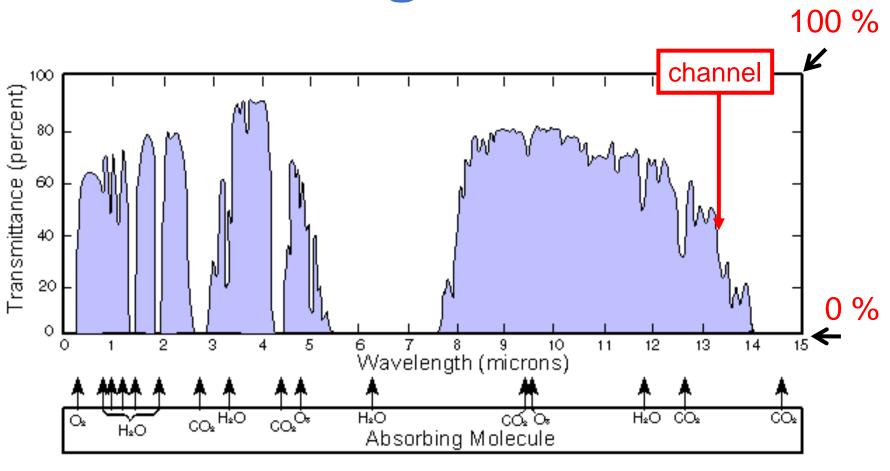
#### Challenge 1: Sounding channels for the lower troposphere ...

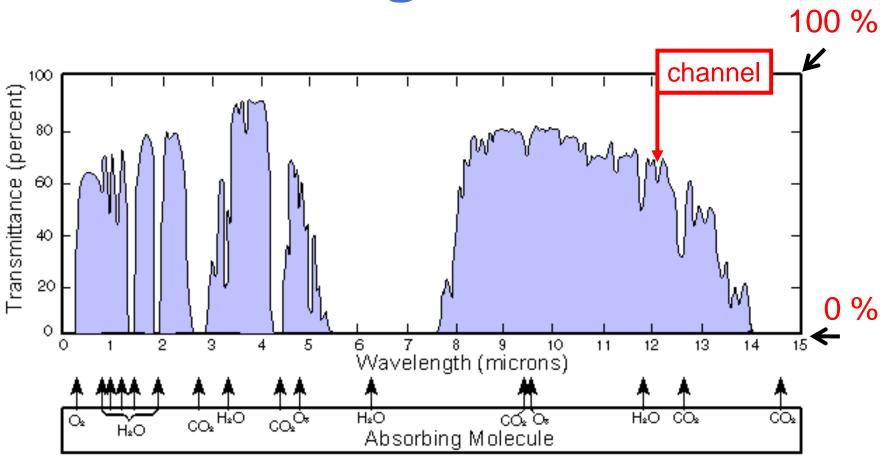
## Atmospheric sounding channels

$$L(\nu) = \int_0^\infty B(\nu, T(z)) \left[ \frac{d\tau(\nu)}{dz} \right] dz + \frac{\text{Surface}}{\text{errorsion}} + \frac{\text{Surface}}{\text{scattering}} + \frac{\text{Cloutzain}}{\text{contrivision}} + \dots$$

# Strong absorbing sounding channels







$$L(\nu) = \int_0^\infty B(\nu, T(z)) \left[ \frac{d\tau(\nu)}{dz} \right] dz + \frac{\text{Surface}}{\text{errorsion}} + \frac{\text{Surface}}{\text{scattering}} + \frac{\text{Cloubtain}}{\text{contribution}} + \dots$$

$$L(\nu) = \int_0^\infty B(\nu, T(z)) \left[ \frac{d\tau(\nu)}{dz} \right] dz + \frac{\text{Surface}}{\text{emission}} + \frac{\text{Surface}}{\text{scattering}} + \frac{\text{Clouvain}}{\text{scattering}} + \dots$$

$$L(v) = \int_0^\infty B(v, T(z)) \left[ \frac{d\tau(v)}{dz} \right] dz + \frac{\text{Surface}}{\text{emission}} + \frac{\text{Surface}}{\text{scattering}} + \frac{\text{Cloud/rain}}{\text{contribution}} + \dots$$

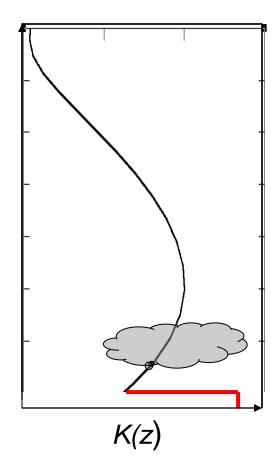
#### Sensitivity to the surface skin and clouds

By placing sounding channels in parts of the spectrum where the absorption is weak we obtain temperature (and humidity) information from the lower troposphere (low peaking weighting functions).

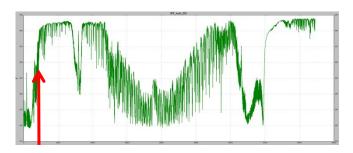
#### BUT ...

These channels (obviously) become more sensitive to surface emission and the effects of cloud and precipitation.

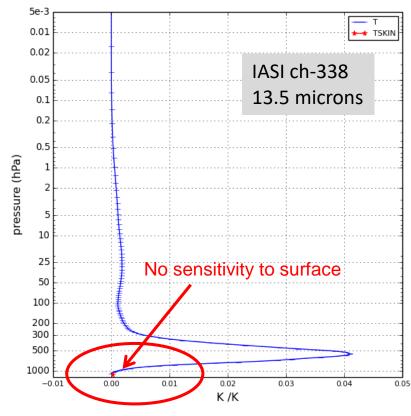
In most cases surface or cloud contributions will dominate the atmospheric signal in these channels and it is difficult to use the radiance data safely (i.e. we may alias a cloud signal as a temperature adjustment)

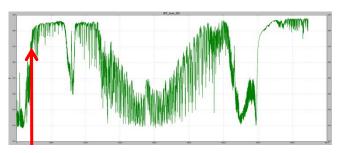


#### **Lower Tropospheric Channels**

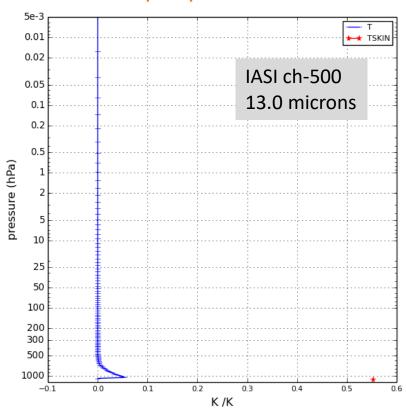


Tropospheric channel

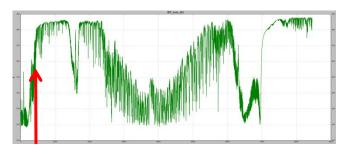




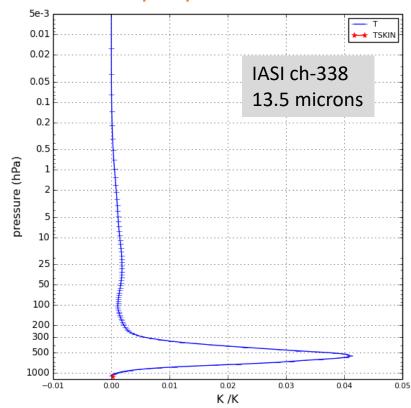
Low tropospheric channel

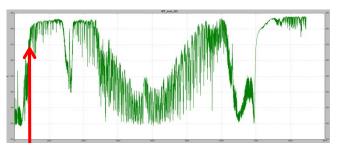


#### **Lower Tropospheric Channels**

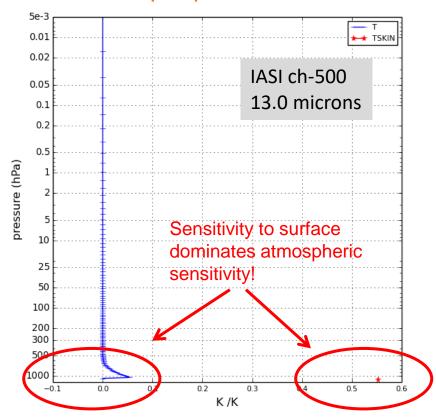


Tropospheric channel





Low tropospheric channel



#### Challenge 2: When the absorbing gas is itself a variable ...

## When the absorbing gas is itself a variable ...

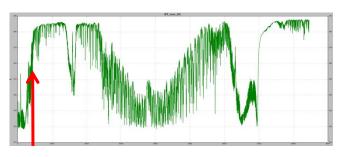
When the primary absorber in a sounding channel is a well mixed gas (e.g. oxygen or carbon dioxide) the radiance essentially gives information about variations in the atmospheric temperature profile only.

$$L(v) = \int_0^\infty B(v, T(z)) \left[ \frac{d\tau(v)}{dz} \right] dz$$

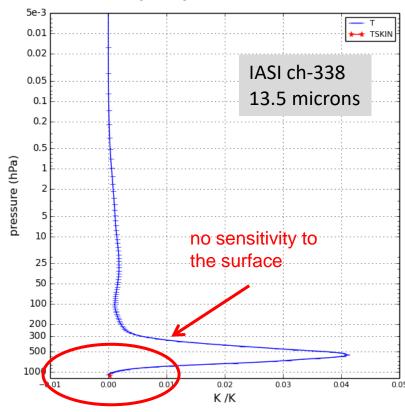
When the primary absorber is **not well mixed** (e.g. water vapour, ozone) the weighting functions **depend on the state** of the atmosphere and radiance gives **ambiguous** information about the temperature profile and the absorber distribution. This ambiguity must be resolved by:

- differential channel sensitivity
- •synergistic use of well mixed channels (constraining the temperature)
- the background error covariance (+ physical constraints)

## Temperature Channels sensitive to water vapour

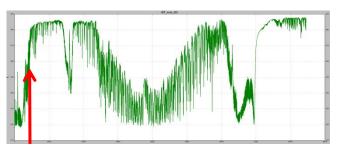


Tropospheric channel

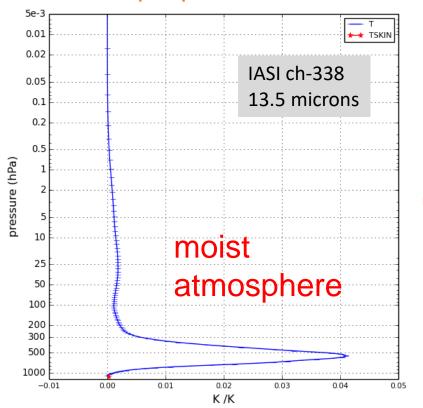


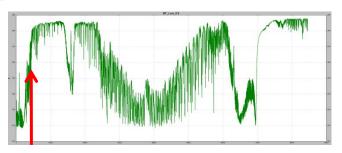
What happens in a very **dry** atmosphere?

## Temperature Channels sensitive to water vapour

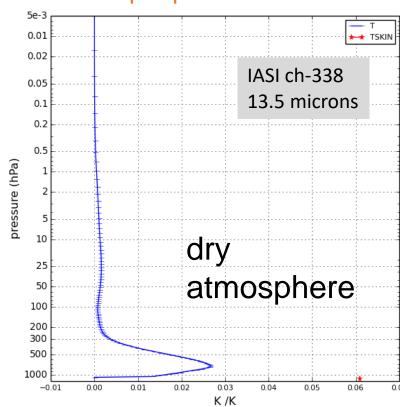


Tropospheric channel

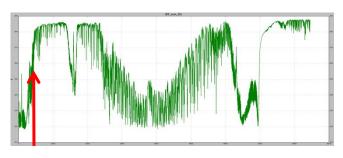




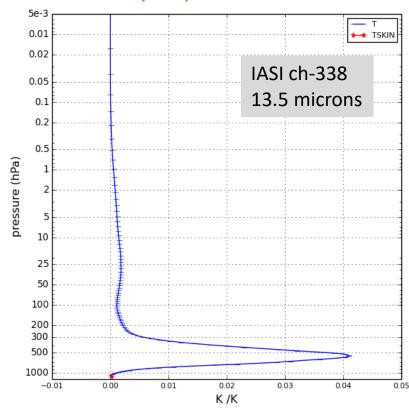
Tropospheric channel

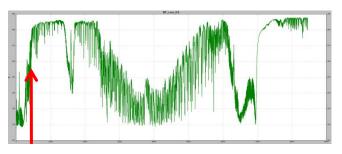


## Temperature Channels sensitive to water vapour

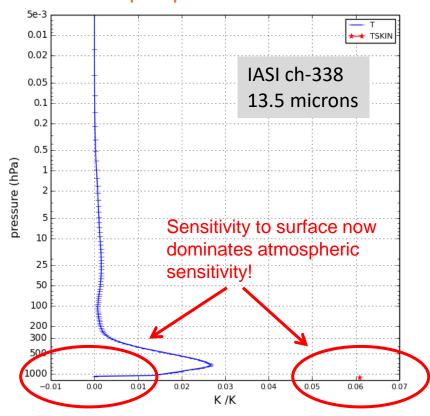


#### Tropospheric channel



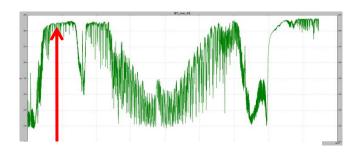


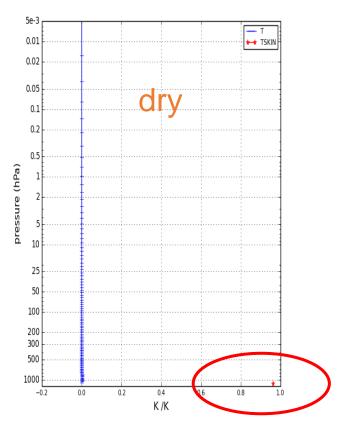
Tropospheric channel

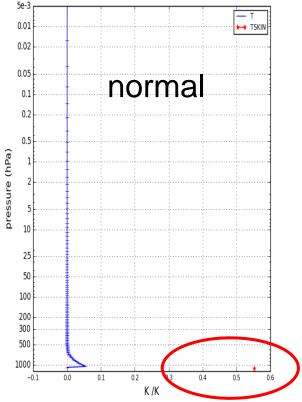


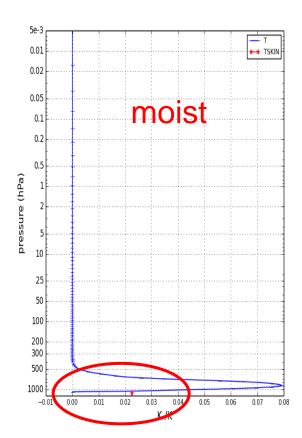
## Window Channels sensitive to water vapour

IASI ch-1200 10.5 microns



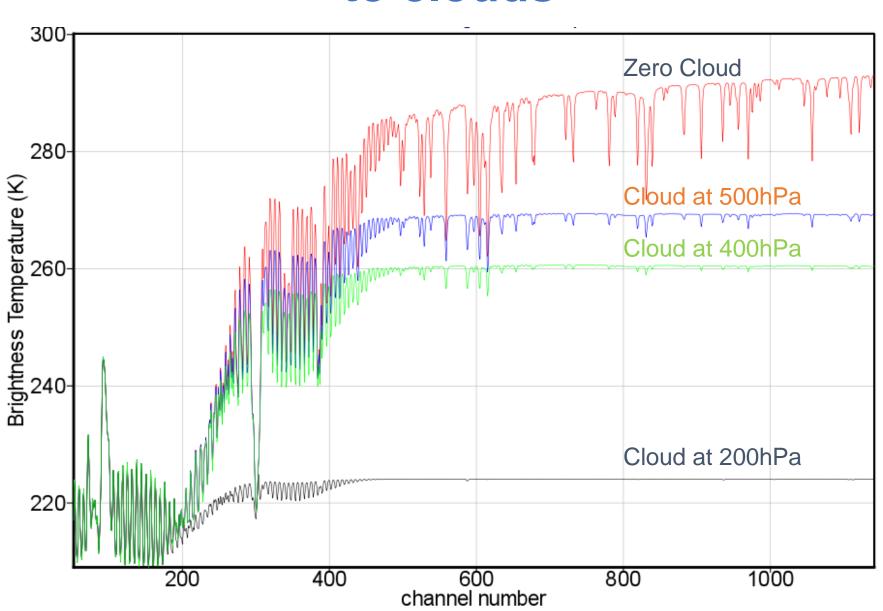




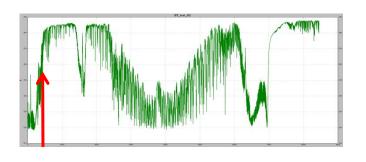


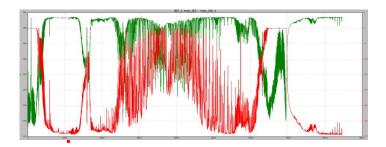
## Challenge 3: Clouds...

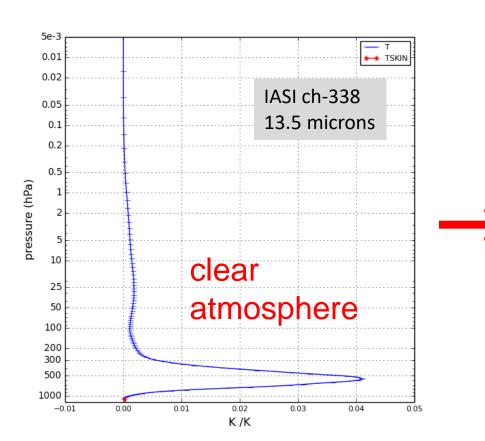
### Temperature Channels sensitive to clouds

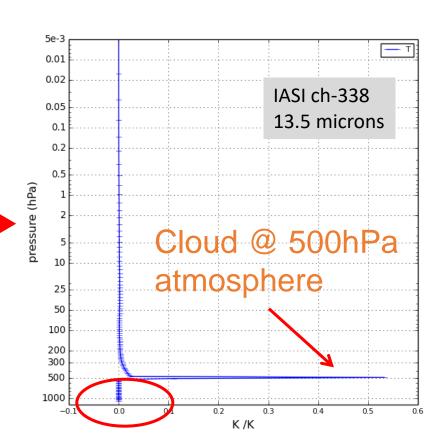


### Temperature Channels sensitive to clouds









## Challenge 3: Clouds...

Dedicated lecture on clouds in IR next.....

### Summary

We now have excellent high quality (and <u>high spectral</u> <u>resolution</u>) measurements of atmospheric infrared radiation.

Instruments such as IASI provide good vertical coverage of the atmosphere by sampling absorption of variable strength. Higher <u>vertical resolution</u> is achieved due to the <u>high number of channels</u>.

Channels in the <u>lower troposphere</u> are also sensitive to the <u>surface</u> (and clouds).

Channels affected by absorption by <u>variable species</u> (e.g. humidity) provide different information in different atmospheres.

### Questions?

#### Units

Radiance = 
$$W \cdot sr - 1 \cdot m - 2 \cdot Hz - 1$$

#### Brightness temperature (K)

$$T_b = \frac{h\nu}{k} \ln^{-1} \left( 1 + \frac{2h\nu^3}{I_{\nu}c^2} \right)$$