

Representing model uncertainty

Stochastic perturbations

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Introduction: Model Uncertainty

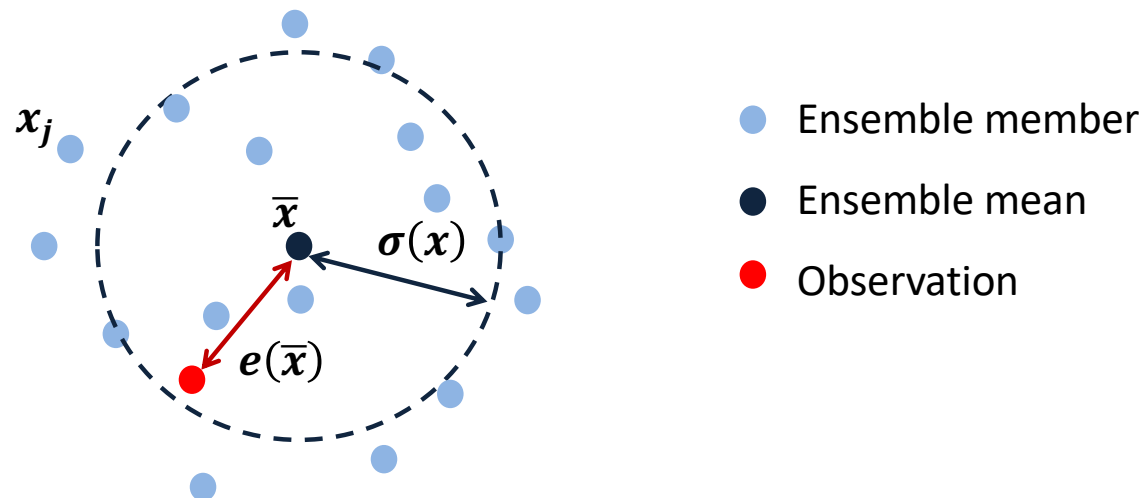
- Ensemble forecasts enable a quantification of the confidence in a forecast, e.g. 10% chance of rain
- An ensemble forecast is made from multiple forecasts or “members”, each member perturbed with respect to the others
- The perturbations comprise
 - a) different initial conditions for each member, to sample the uncertainty in our description of the initial state (*Simon Lang’s lecture*); and
 - b) a different forecast model for each member, to sample the uncertainty due to the model integrations or the “*model uncertainty*”
- To date, much effort has been focused on **model uncertainty** due to the parametrization schemes that describe sub-grid atmospheric physics --- representing this with stochastic perturbations gives rise to “*stochastic physics*”

Using stochastic physics to represent model uncertainty

- Why do we represent model uncertainty in an ensemble forecast?
- What are the sources of model uncertainty?
- How do we currently represent model uncertainty in the IFS?
- Ongoing work towards process-level simulation of model uncertainty

Ensemble reliability

- In a reliable ensemble, **ensemble spread** is a predictor of **ensemble error**



i.e. averaged over many ensemble forecasts,

$$e(\bar{x}) \approx \sigma(x)$$

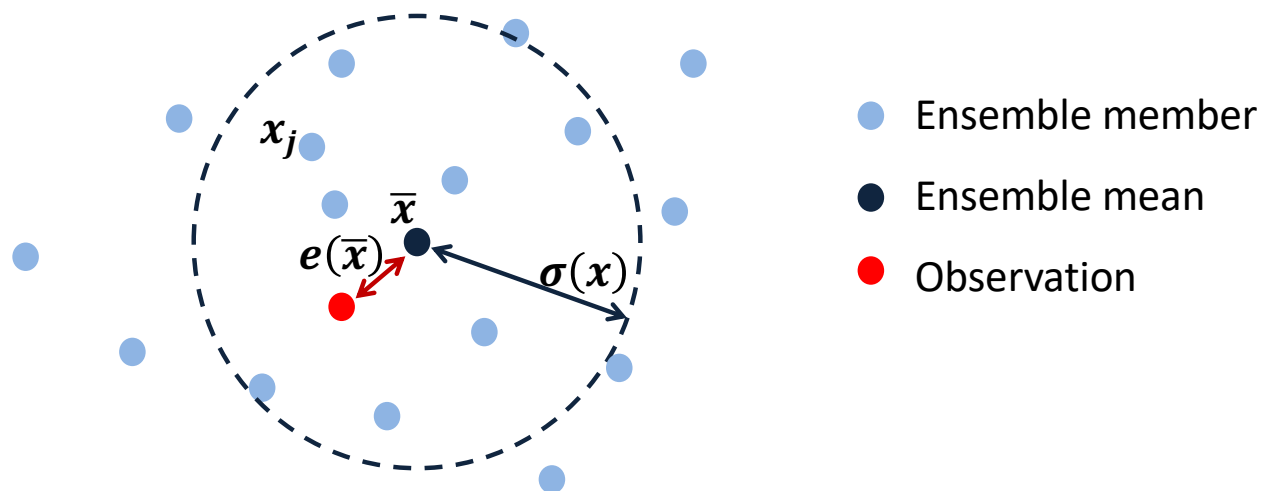
For a thorough discussion of this relationship:

Martin Leutbecher's lectures

Ensemble reliability

- In an **over**-dispersive ensemble,

$$e(\bar{x}) \ll \sigma(x)$$



and ensemble spread does not provide a good estimate of error.

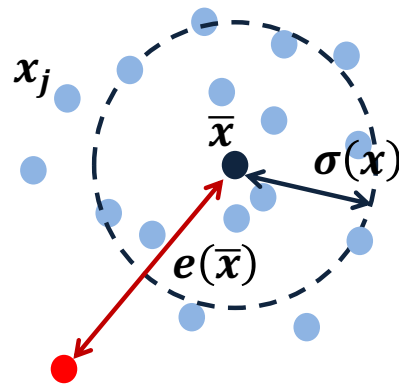
The relatively large spread implies large uncertainty and hence, likely large error:

an “under-confident forecast”

Ensemble reliability

- In an **under**-dispersive ensemble,

$$e(\bar{x}) \gg \sigma(x)$$



- Ensemble member
- Ensemble mean
- Observation

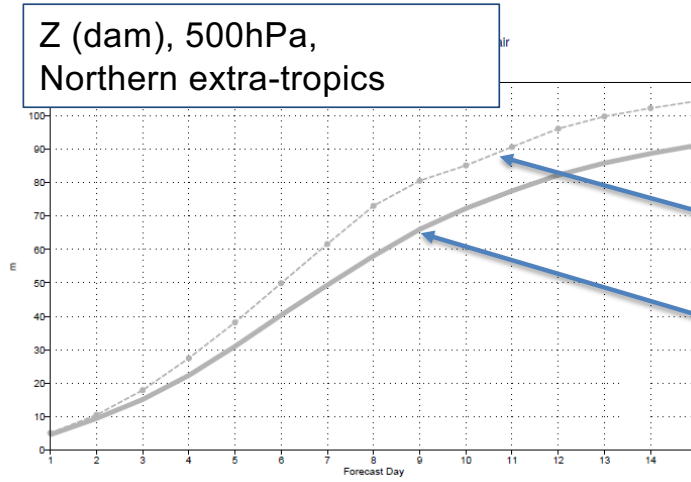
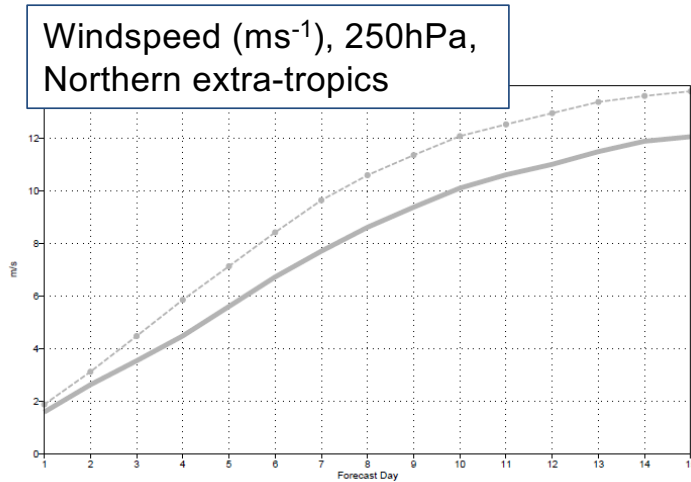
The small spread implies low uncertainty and hence, small errors:

an “over-confident forecast”

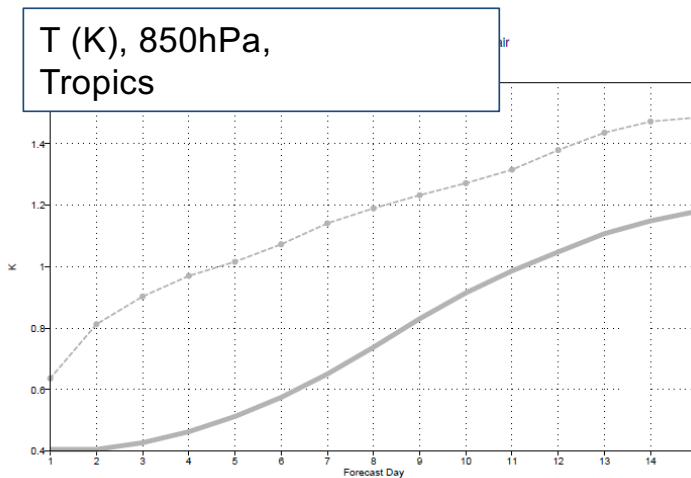
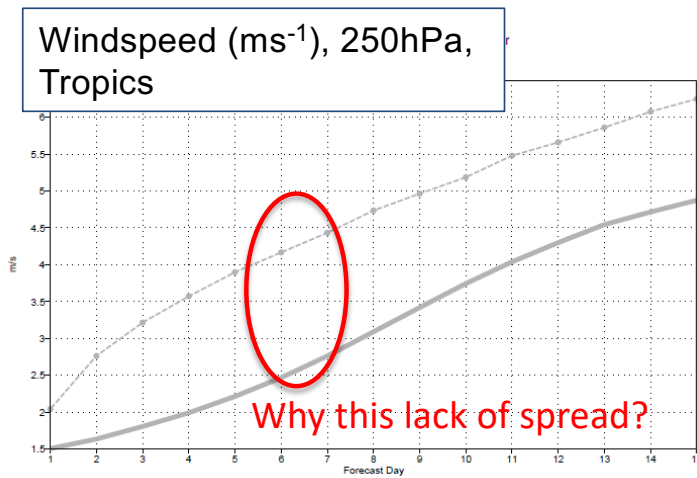
What happens when the ensemble includes no representation of model uncertainty?

Ensemble forecasts with only initial conditions perturbations

Ensemble mean RMSE ("Error") & standard deviation ("Spread")

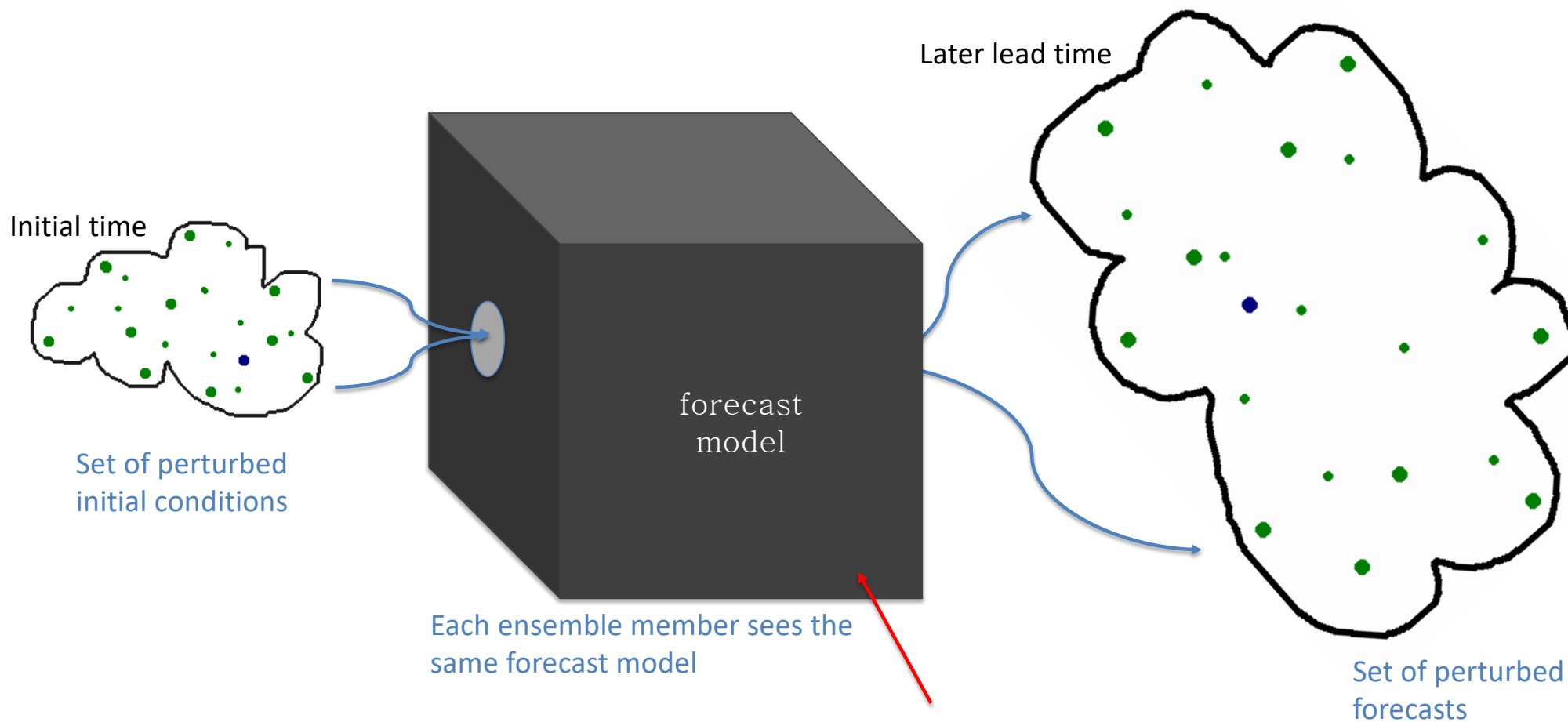


RMSE ensemble mean
RMS ensemble stdev



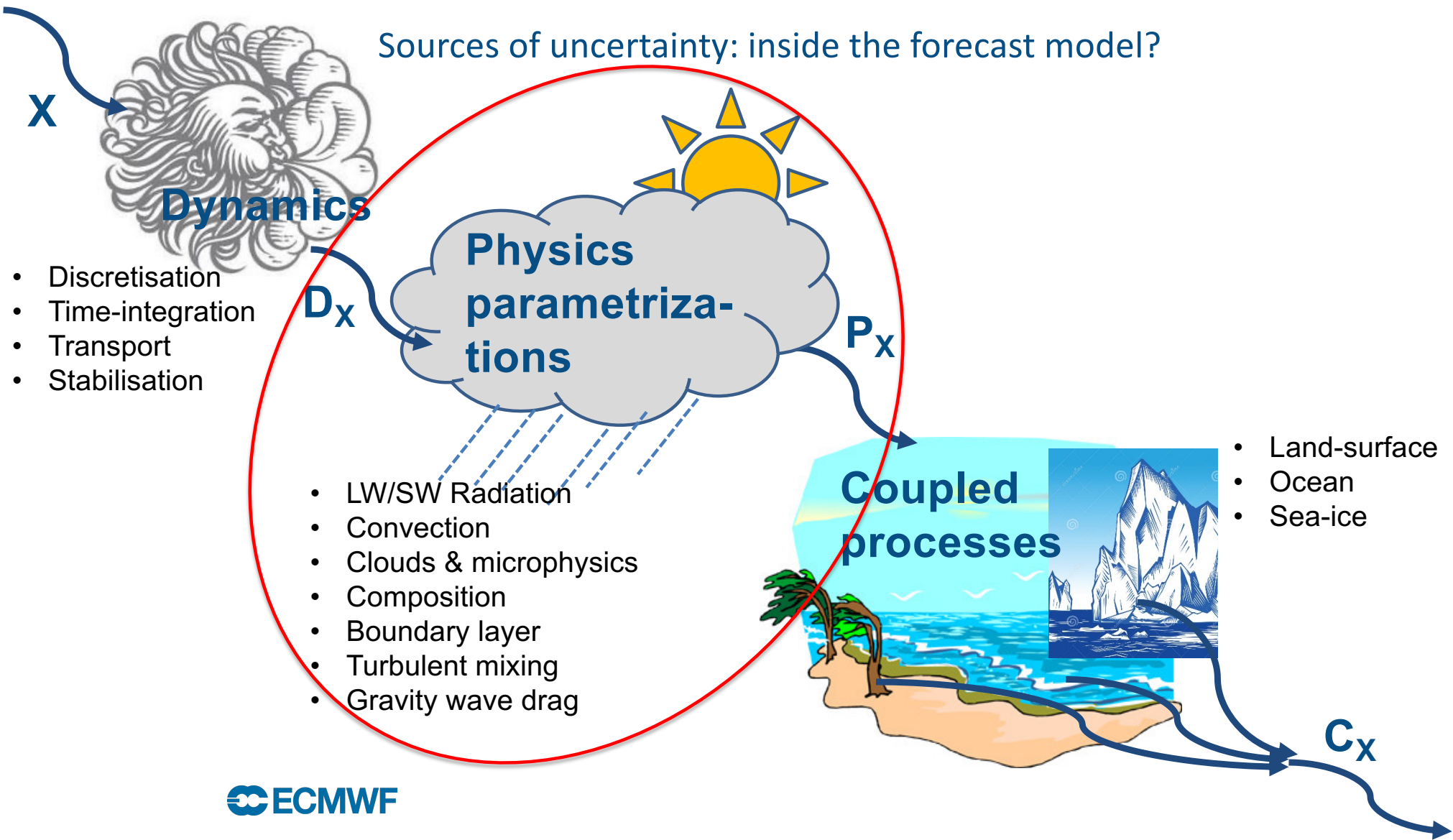
CY47R3
TCo399L137, dt=1200s
30 dates (Dec 2019)
8 perturbed fcs

Sources of uncertainty: initial conditions



What about "model uncertainty"?

Sources of uncertainty: inside the forecast model?

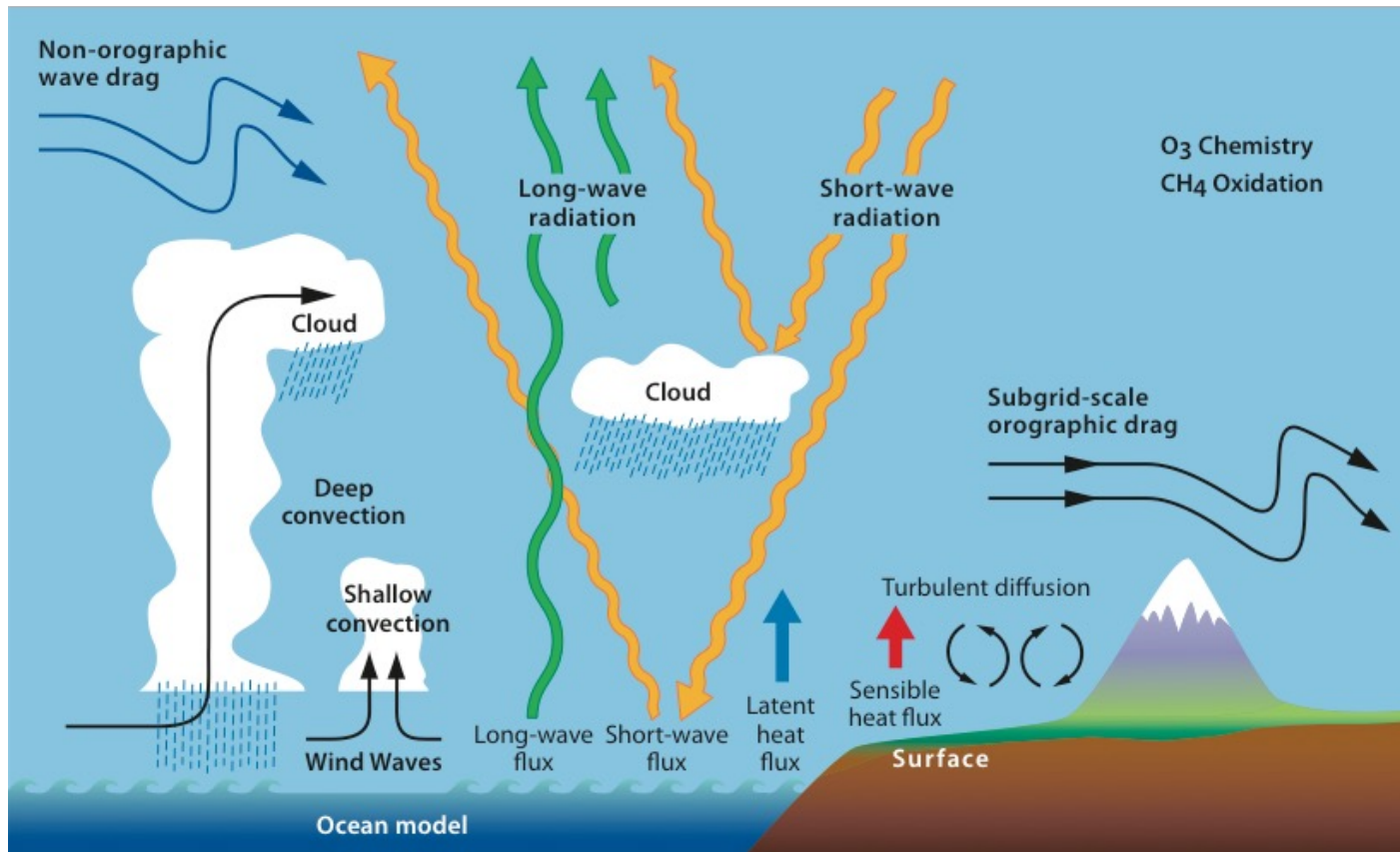


- Discretisation
- Time-integration
- Transport
- Stabilisation

- LW/SW Radiation
- Convection
- Clouds & microphysics
- Composition
- Boundary layer
- Turbulent mixing
- Gravity wave drag

- Land-surface
- Ocean
- Sea-ice

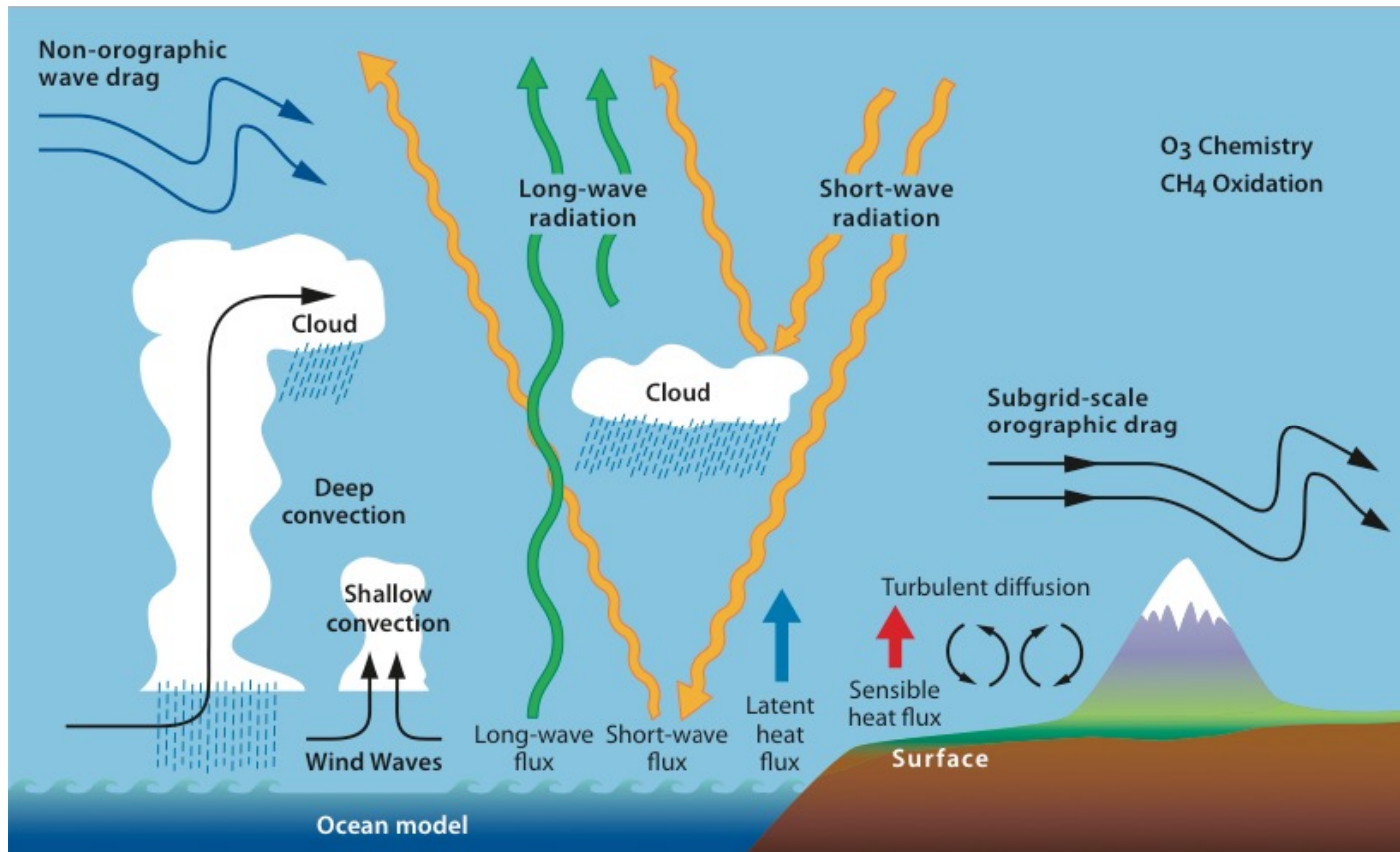
Model uncertainty: parametrized atmospheric physics processes



Uncertainties arise due to:

- Inability to resolve sub-grid scales, e.g.
 - Surface drag (orography/waves)
 - Convection rates (occurrence / en/detrainment)
 - Phase transitions
 - Radiation transfer in cloudy skies
- Poorly constrained parameters, e.g.
 - Vertical cloud-overlap (radiation)
 - Composition
 - Non-orographic drag

Model uncertainty: parametrized atmospheric physics processes



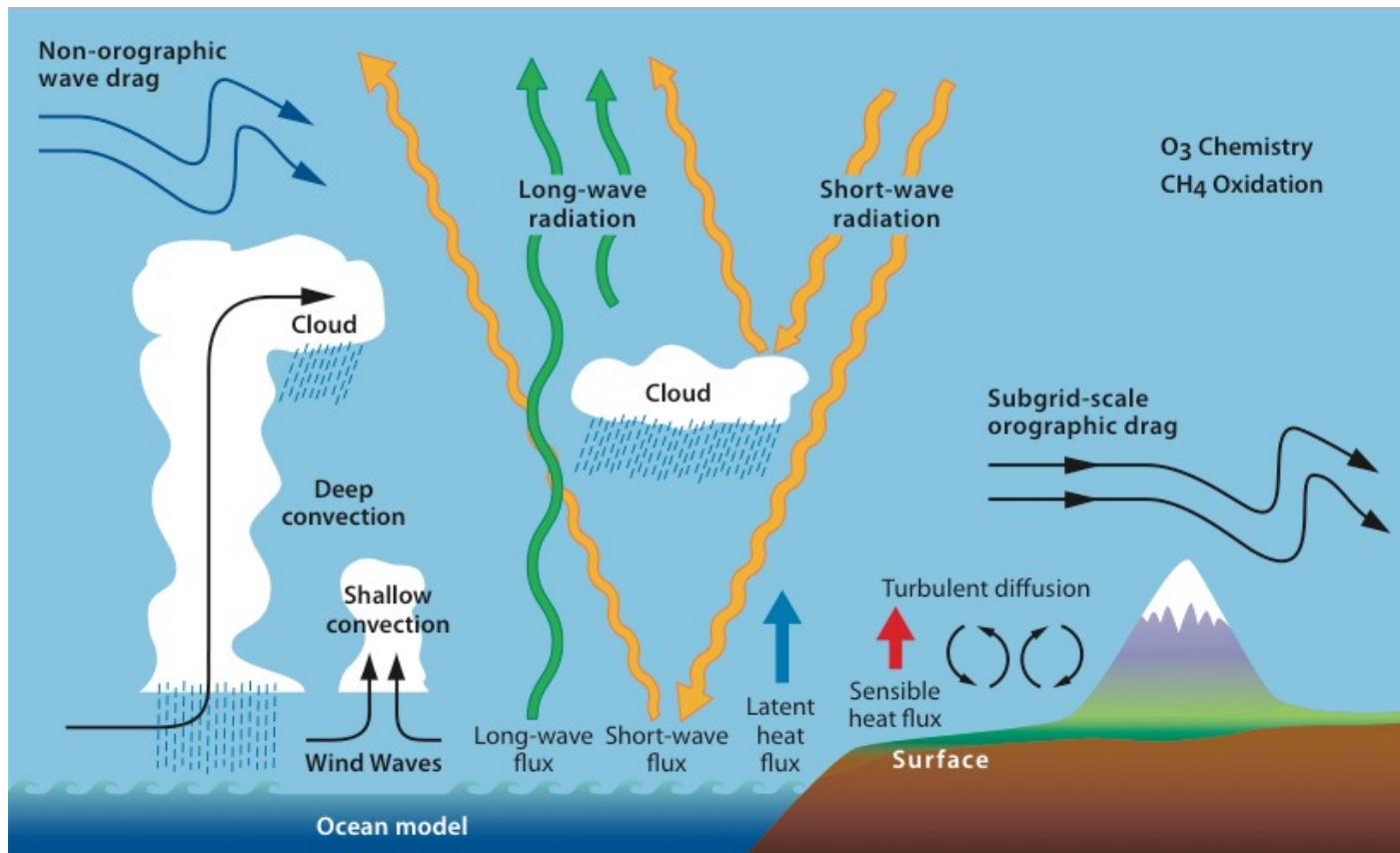
“Let’s take the positives”

Parametrisation schemes:

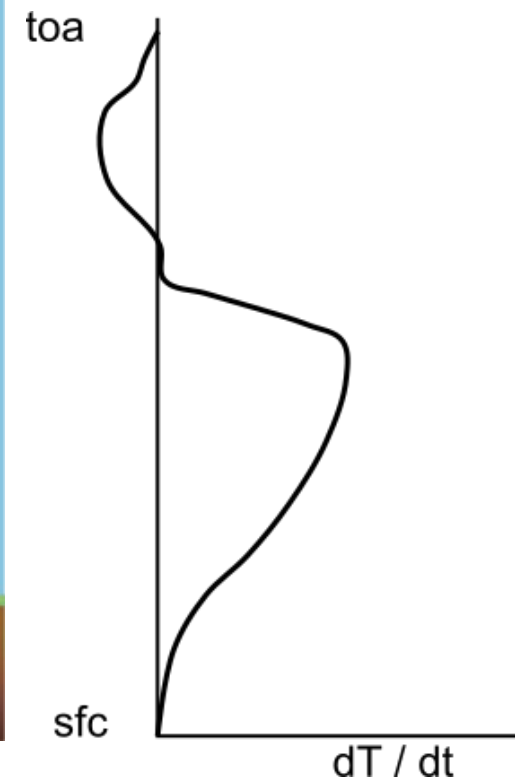
- developed/operate together
- highly tuned for best performance

Seek a description of uncertainty that retains consistencies of the representation of the physical processes.

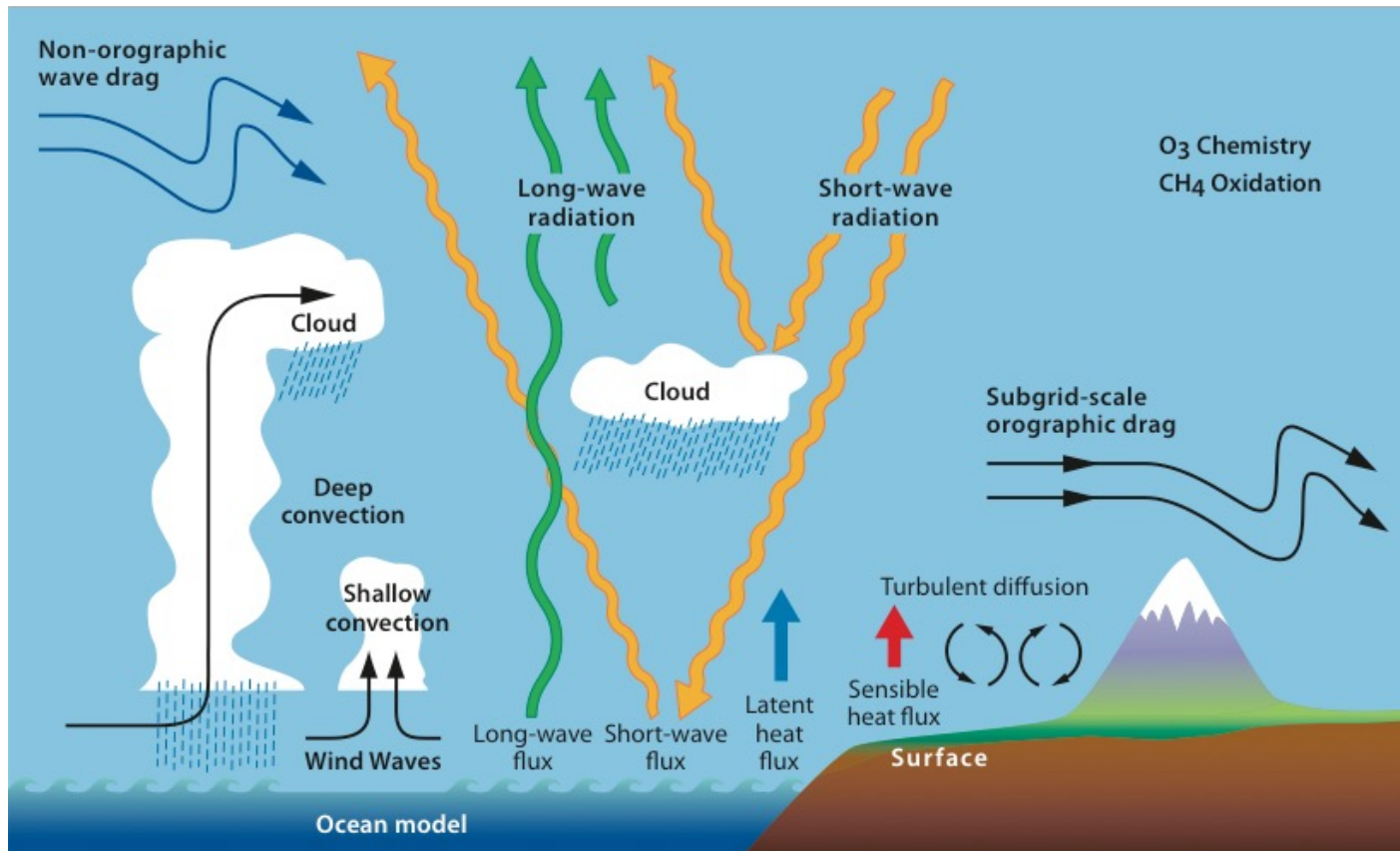
Model uncertainty: parametrized atmospheric physics processes



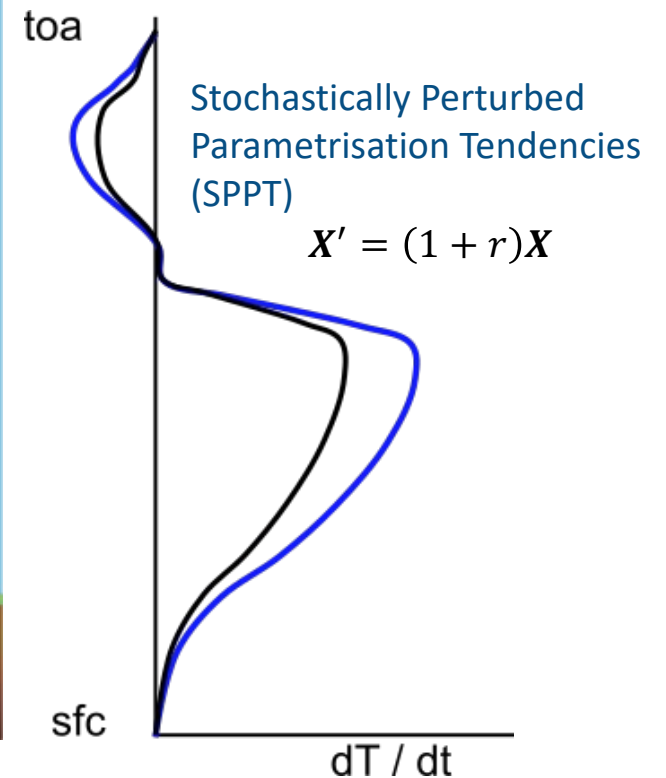
e.g. profile of heating rates from physics parametrisations:



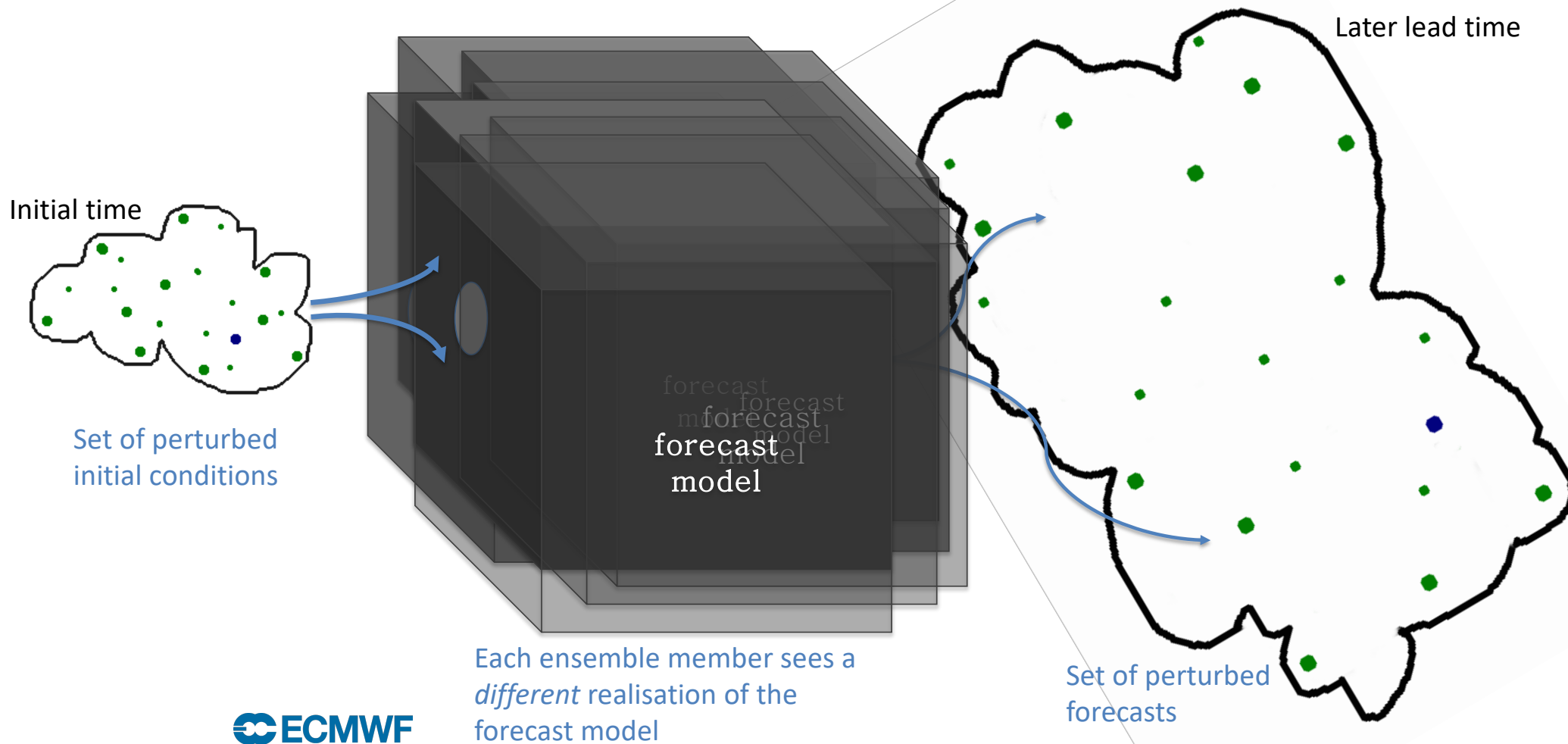
Model uncertainty: parametrized atmospheric physics processes



Proposal: represent uncertainties with a perturbation proportional to the profile of net physics tendencies



Sources of uncertainty: accounting for model uncertainty

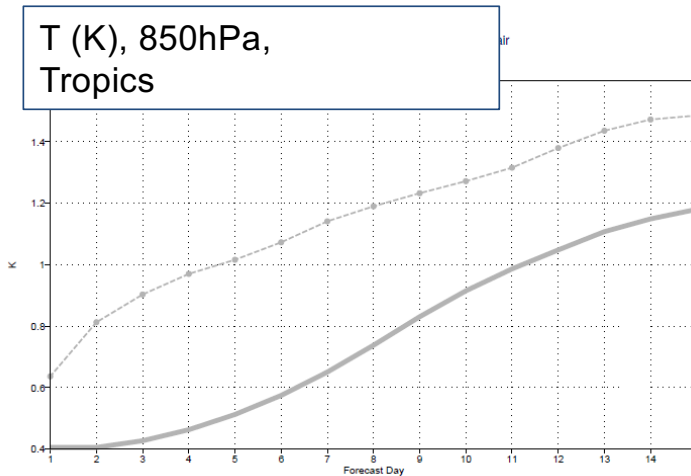
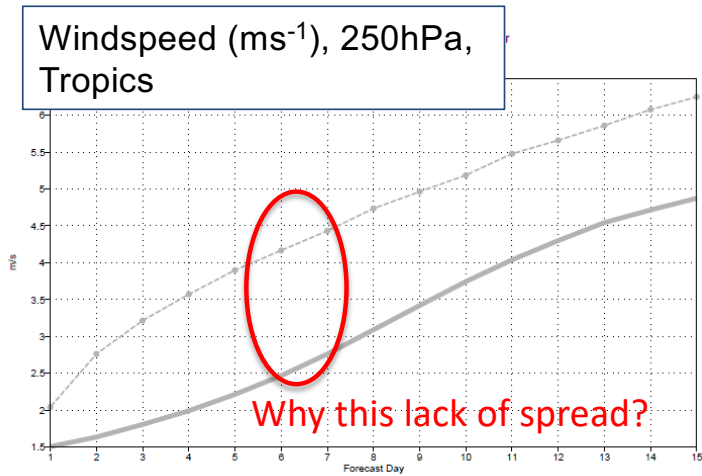
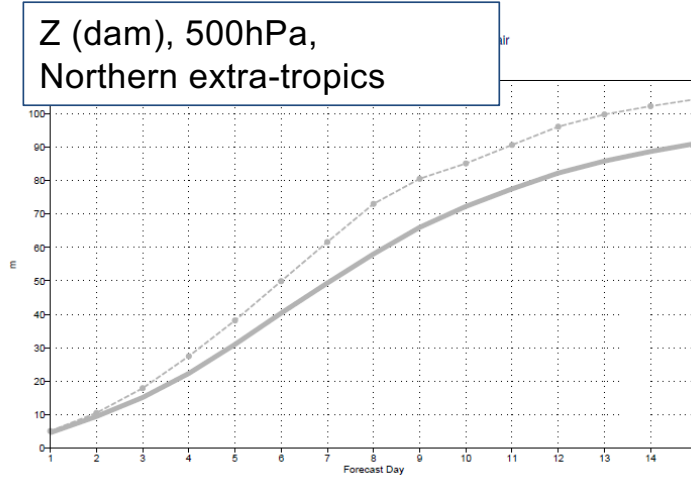
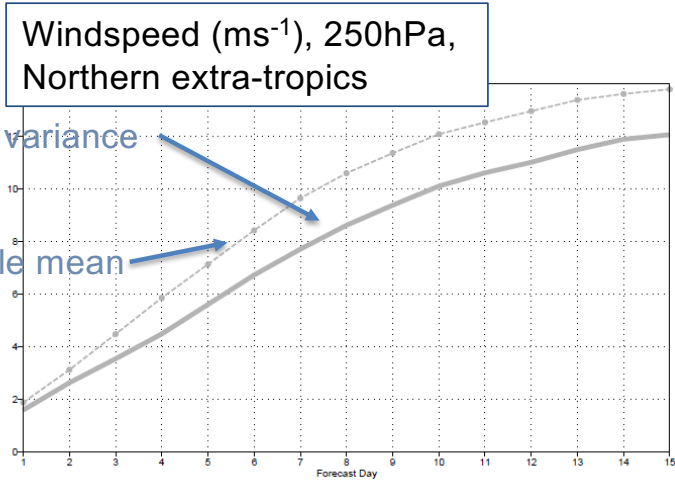


Recall: Ensemble forecasts: with initial conditions perturbations (IP) only

Ensemble mean RMSE ("Error") & standard deviation ("Spread")

RMS ensemble variance
("spread")

RMSE ensemble mean
("error")

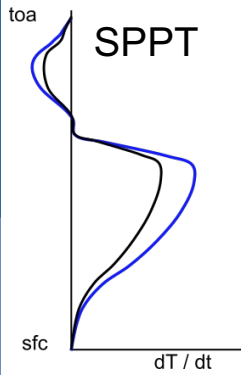


CY47R3

TCo399L137, dt=1200s

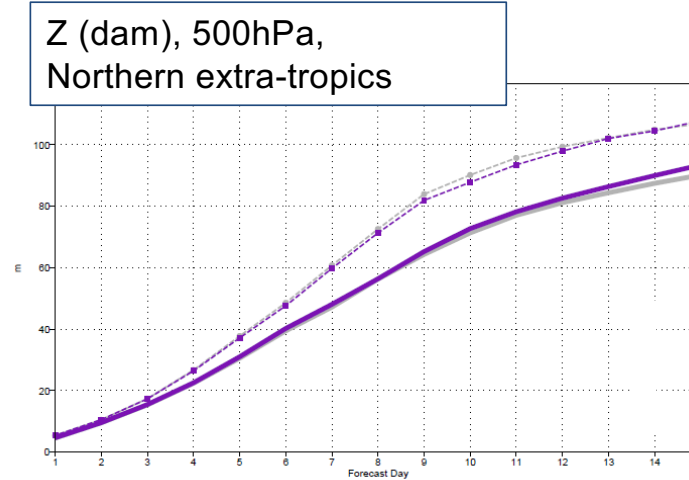
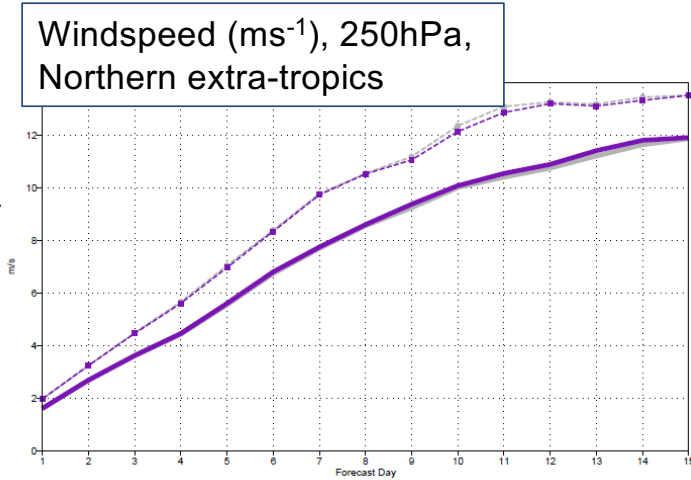
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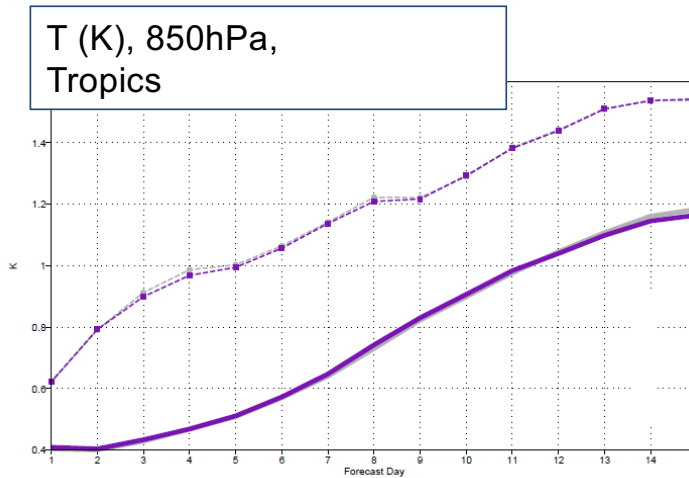
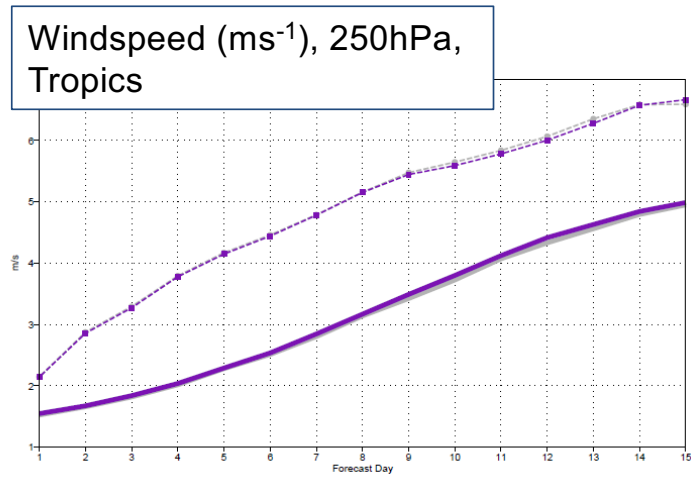


Ensemble forecasts: with **grid-scale** model uncertainty perturbations (SPPT)

Ensemble mean RMSE ("Error") & standard deviation ("Spread")

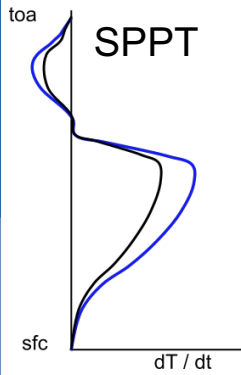


— IP only
 — IP + SPPT*
 (*white noise wrt time/horizontal)



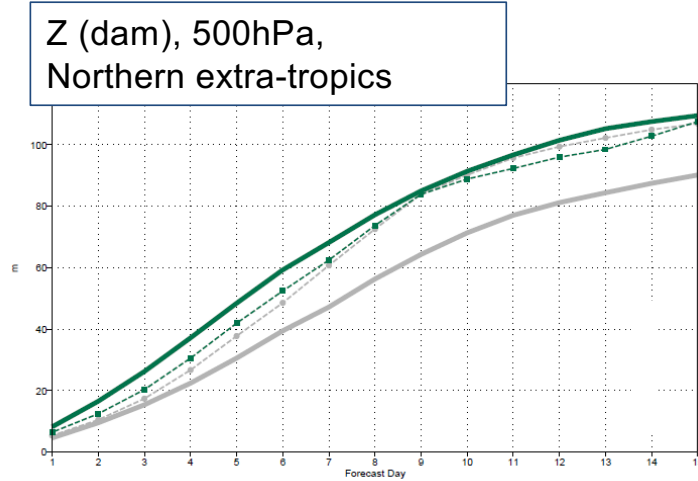
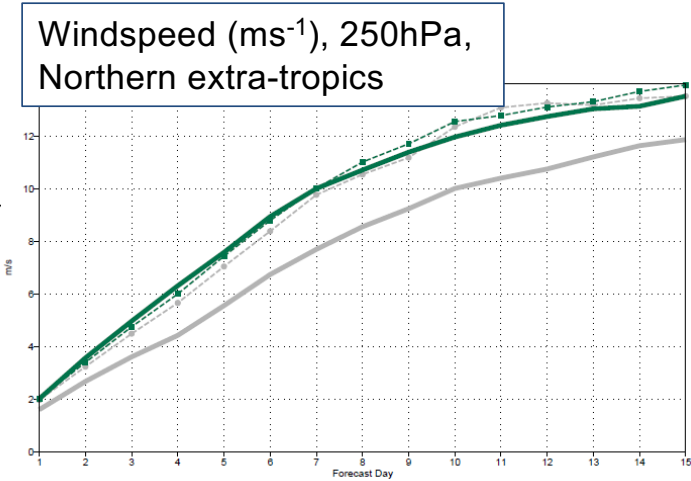
Uncorrelated noise yields little benefit

CY47R3
 TCo399L137, dt=1200s
 11 dates (Dec 2019/Jan 2020)
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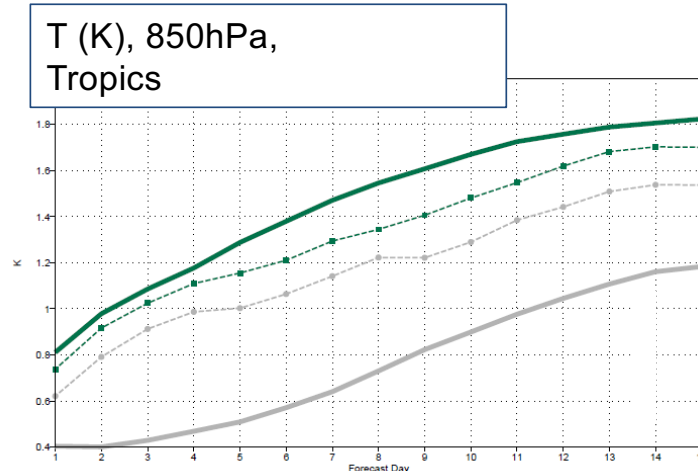
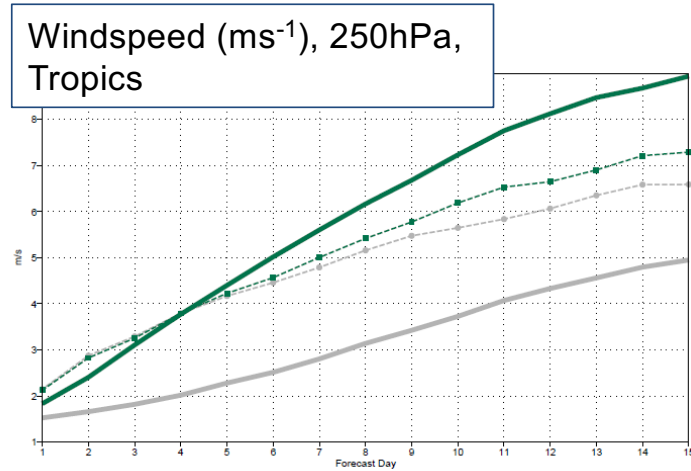


Ensemble forecasts: with **fixed** model uncertainty perturbations (SPPT)

Ensemble mean RMSE ("Error") & standard deviation ("Spread")



— IP only
 — IP + SPPT*
 (*fixed perts wrt time/horizontal)

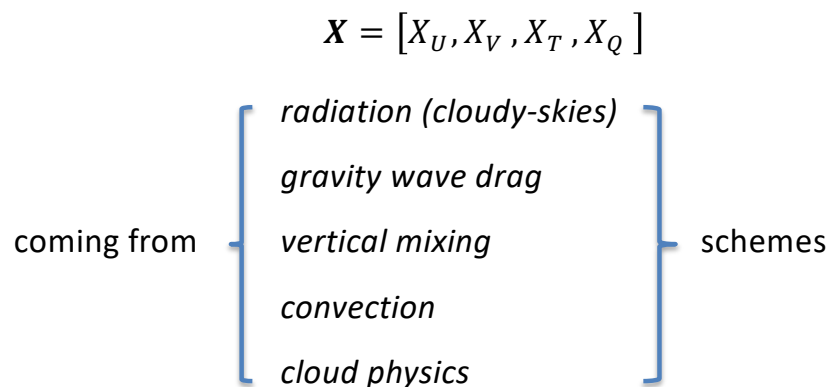


Fixed perturbations yield increased errors

CY47R3
 TCo399L137, dt=1200s
 11 dates (Dec 2019/Jan 2020)
 8 perturbed fcs

Stochastically Perturbed Parametrisation Tendencies (SPPT) scheme

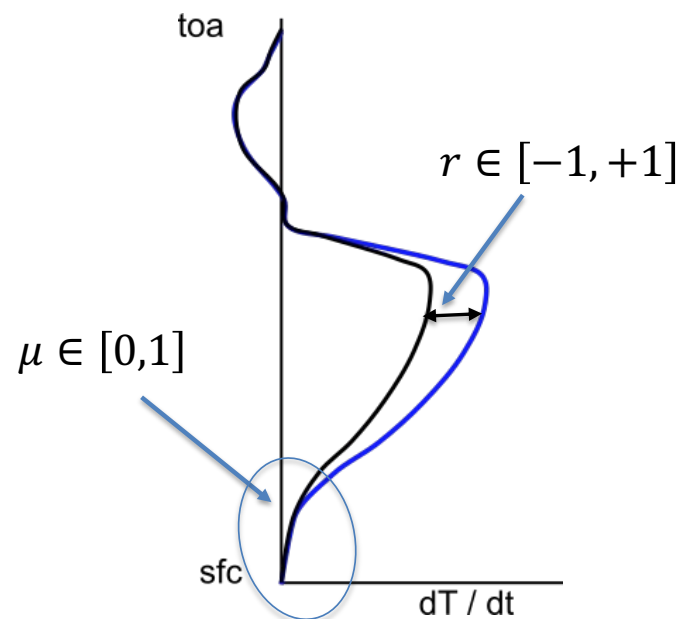
- History (IFS): implemented, 1998 (Buizza et al., 1999); revised, 2009 (Palmer et al., 2009), 2019 (Lock et al., 2019):
- Simulates model uncertainty due to *physics parameterisations* by
 - taking the net tendencies from the physics parametrisations (excl. clear-sky heating rates):



- and perturbing with multiplicative noise $r \in [-1, +1]$ as:

$$\mathbf{X}' = (1 + \mu r)\mathbf{X}$$

where $\mu \in [0,1]$ tapers the perturbations to zero near the surface.

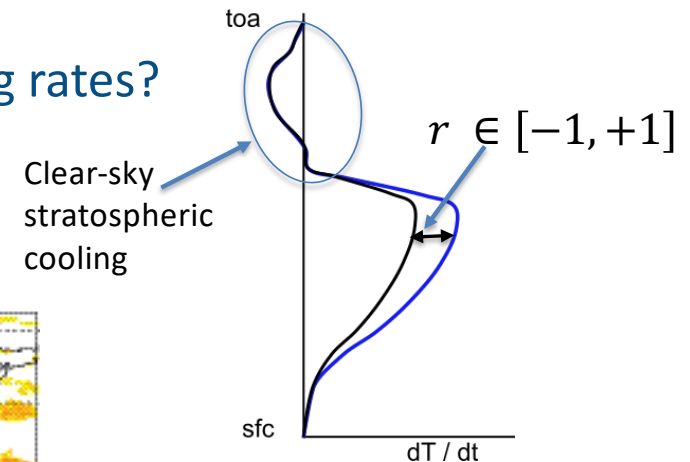


Shutts et al. (2011, ECMWF Newsletter);
 Palmer et al., (2009, ECMWF Tech. Memo.);
 Lock et al., (2019, QJRMS)

SPPT perturbations: why exclude clear-sky heating rates?

- SPPT perturbs net physics tendencies (*excluding clear-sky heating rates*) with multiplicative noise $r \in [-1, +1]$ as:

$$X' = (1 + \mu r)X$$

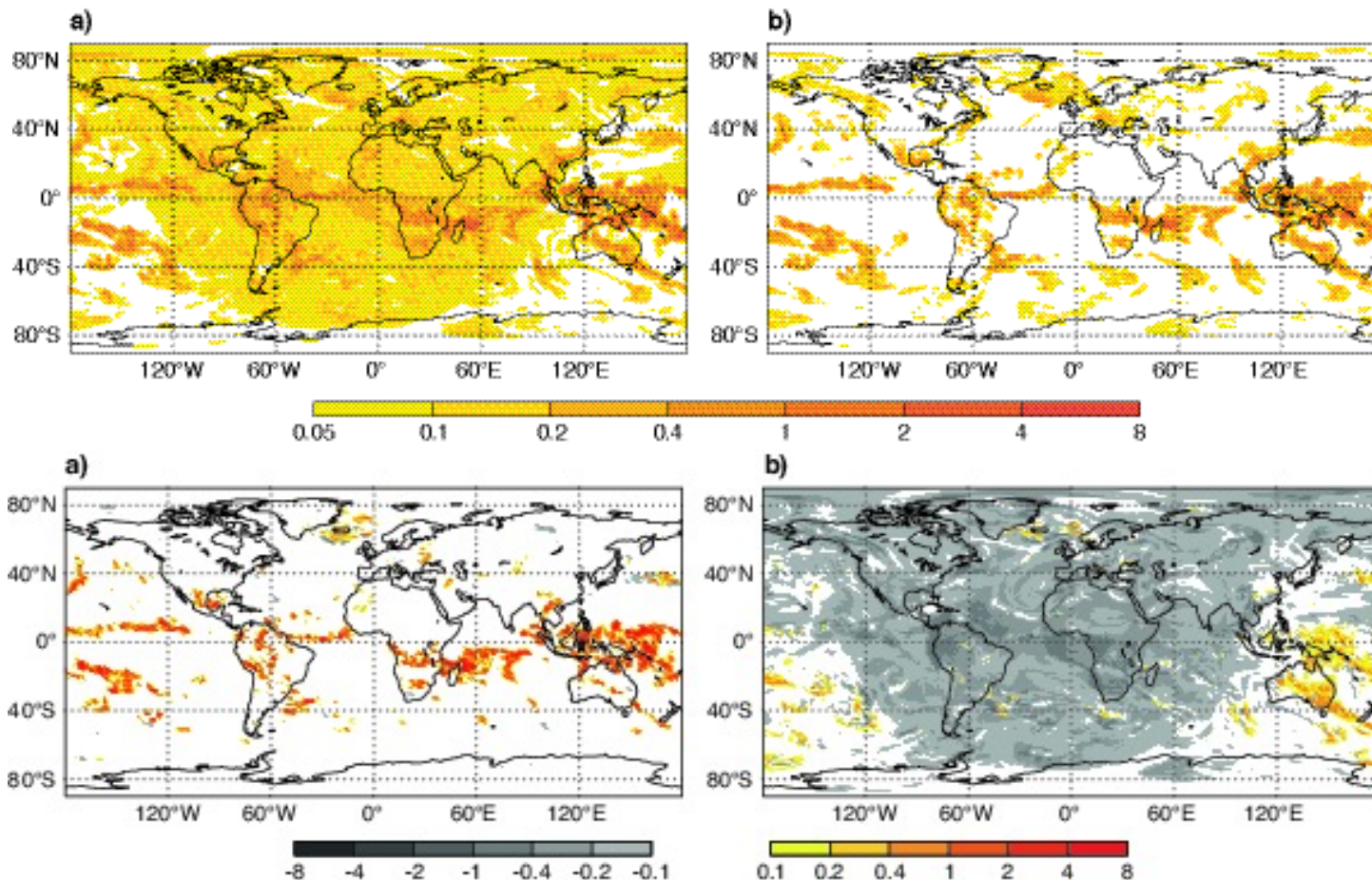


T tendencies from a model level in mid-troposphere accumulated during t+0-3h (K/3h):

Top: Ensemble stdev with SPPT perturbations with (a) clear-sky HRs (a) included & (b) excluded.

Bottom: From control forecast, from (a) convection & (b) radiation schemes

Figure 2 & Figure 1, from Lock et al. (2019, QJRMS)



SPPT random pattern

- 2D random pattern in spectral space:

- First-order auto-regressive [AR(1)] process for evolving spectral coefficients \hat{r}

$$\hat{r}(t + \Delta t) = \phi \hat{r}(t) + \rho \eta(t)$$

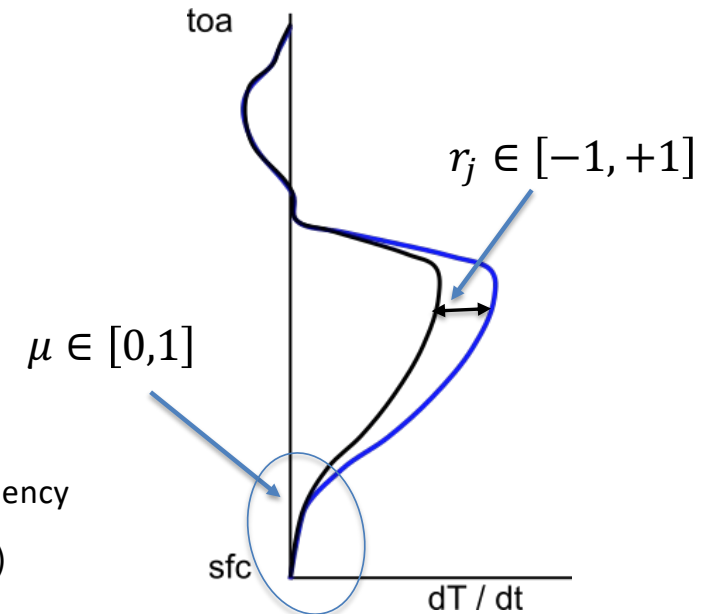
where $\phi = \exp(-\Delta t/\tau)$ controls the correlation over timestep Δt ;

and spatial correlations (Gaussian around the globe) for each wavenumber

define ρ for random numbers, η

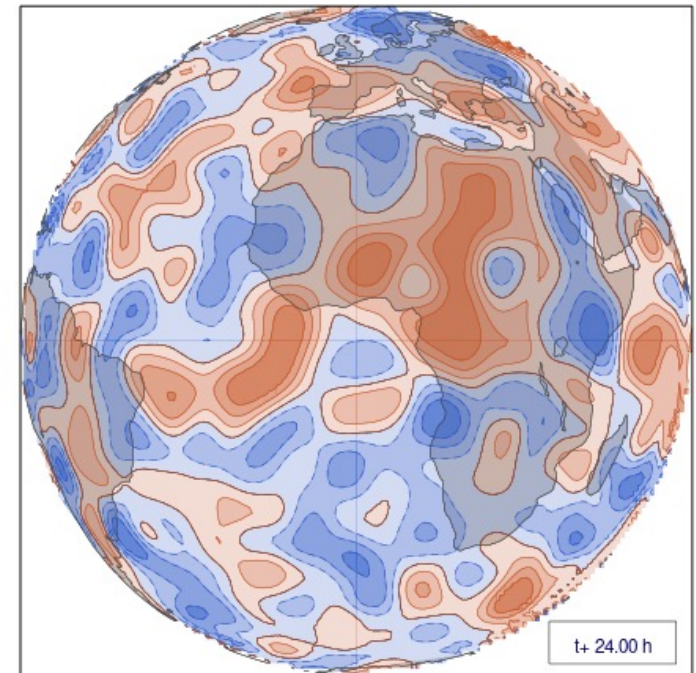
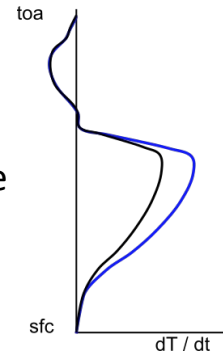
- Resulting pattern mapped into grid-point space r :

- clipped such that $r \in [-1, +1]$ --- prevents perturbation *changing the sign* of the tendency
- same pattern is applied to T, Q, U, V (*excluding clear-sky heating rates from radiation*)
- applied at all model levels to preserve vertical structures**
- ***Except*: tapered to zero at model bottom, to avoid:
 - excessive spread in the boundary layer caused by applying perturbations to large wind tendencies.



SPPT random pattern

- 2D random pattern, r :
 - Time-correlations: AR(1)
 - Spatial-correlations: Gaussian shape around the globe
 - Clipped such that $r \in [-1, +1]$
- Applied at all model levels to preserve vertical structures**
***Except*: tapered to zero at model bottom

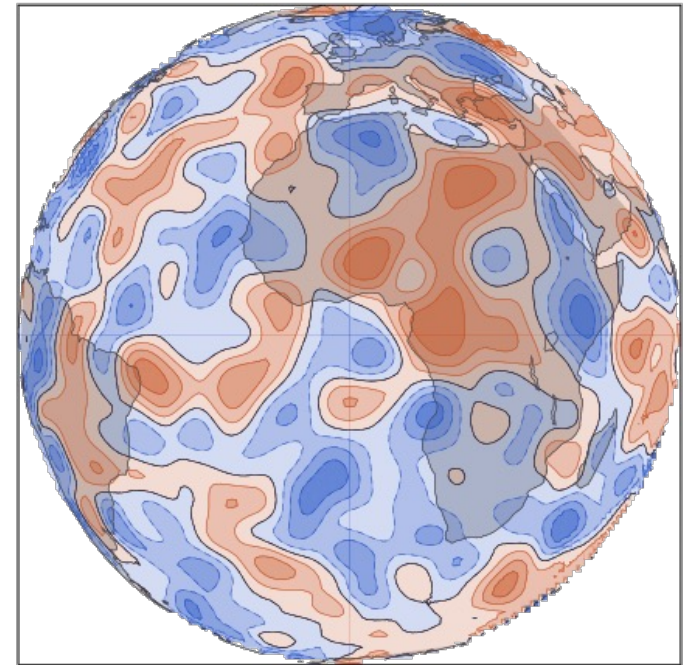
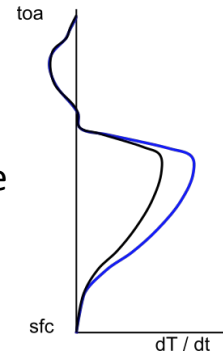


Example random pattern:

- Perturbed member, number 1
- Pattern at $t = 24h$
- Colours: blues = $[-1,0)$, reds = $(0,1]$

SPPT random pattern

- 2D random pattern, r :
 - Time-correlations: AR(1)
 - Spatial-correlations: Gaussian shape around the globe
 - Clipped such that $r \in [-1, +1]$
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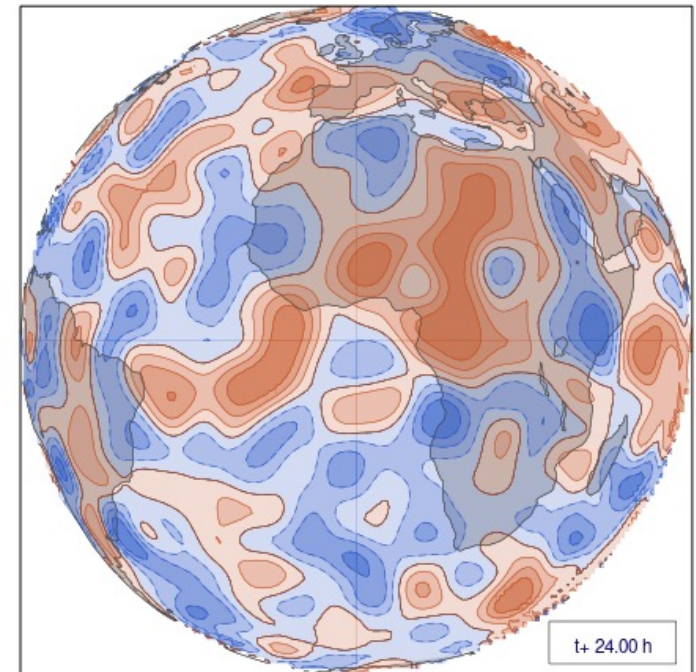


Example random pattern:

- Perturbed member, number 1
- Pattern at $t = 0 \dots 48\text{h}$ ($dt = 15 \text{ min}$)
- Colours: blues = $[-1,0)$, reds = $(0,1]$

SPPT random pattern

- 2D random pattern, r :
 - Time-correlations: AR(1)
 - Spatial-correlations: Gaussian shape around the globe
 - Clipped such that $r \in [-1, +1]$
- Applied at all model levels to preserve vertical structures**
 - ***Except*: tapered to zero at model bottom



- Multi-scale pattern:

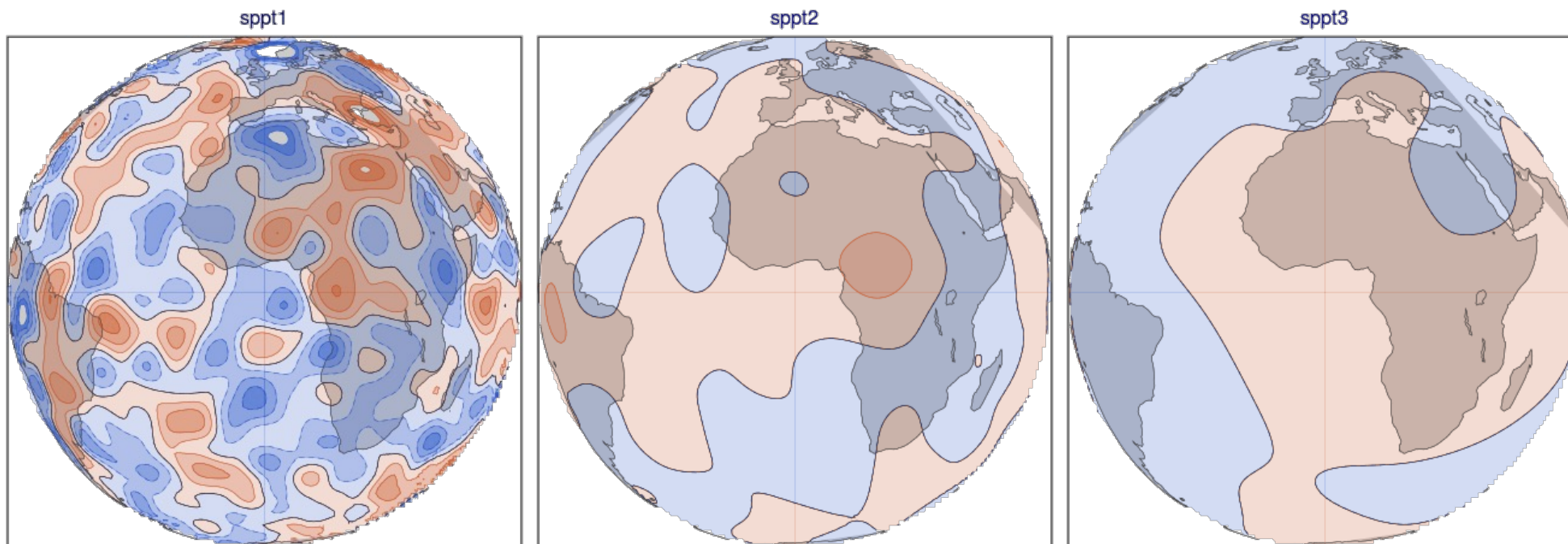
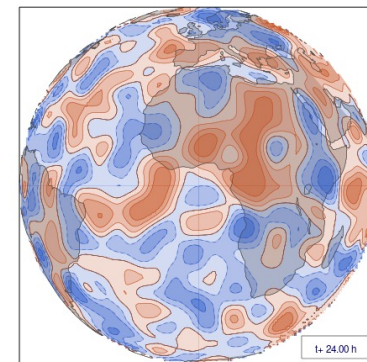
- 3 time/space scales	sppt1	6 hours,	500 km,	$\sigma = 0.42$
- Shortest scales dominate	sppt2	3 days,	1 000 km,	$\sigma = 0.14$
- $\sigma_{3-scale} = 0.4453$	sppt3	30 days,	2 000 km,	$\sigma = 0.048$

SPPT random pattern: multi-scale

sppt1	6 hours,	500 km,	$\sigma = 0.42$
sppt2	3 days,	1 000 km,	$\sigma = 0.14$
sppt3	30 days,	2 000 km,	$\sigma = 0.048$

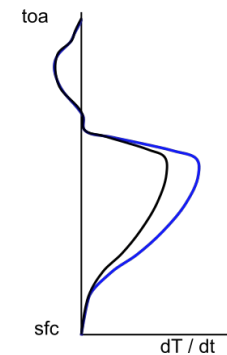
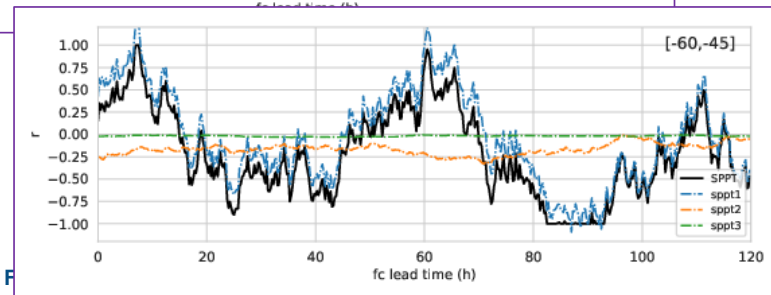
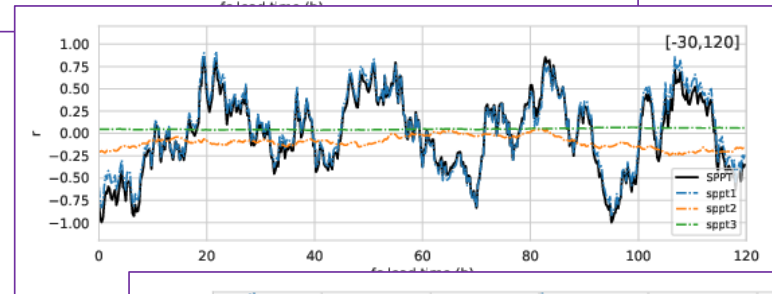
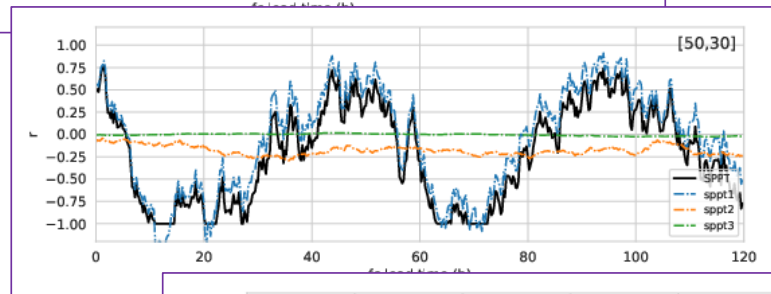
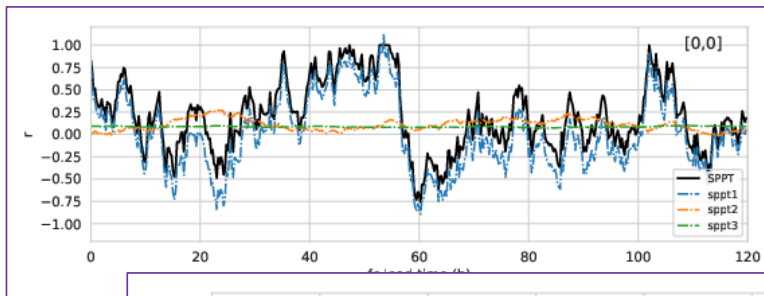
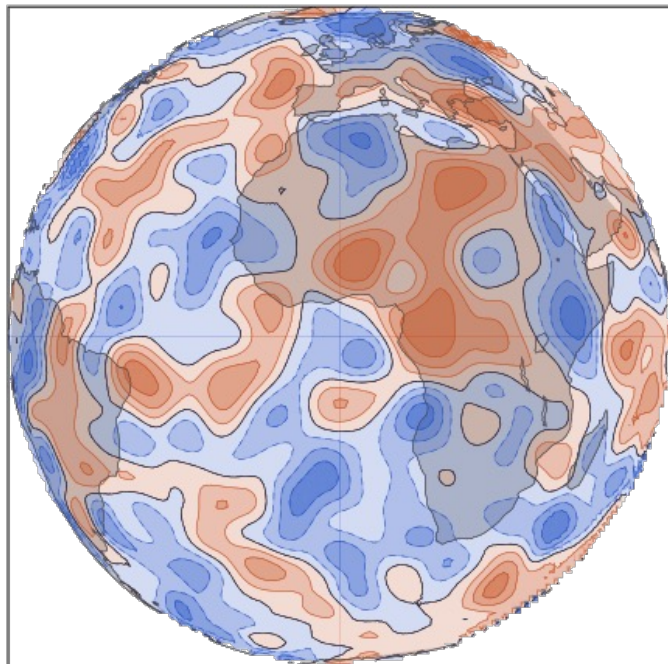
Example random patterns:

- Perturbed member, number 1
- Patterns at $t = 0 \dots 48\text{h}$ ($dt = 15\text{ min}$)
- Colours: blues = $[-1,0)$, reds = $(0,1]$



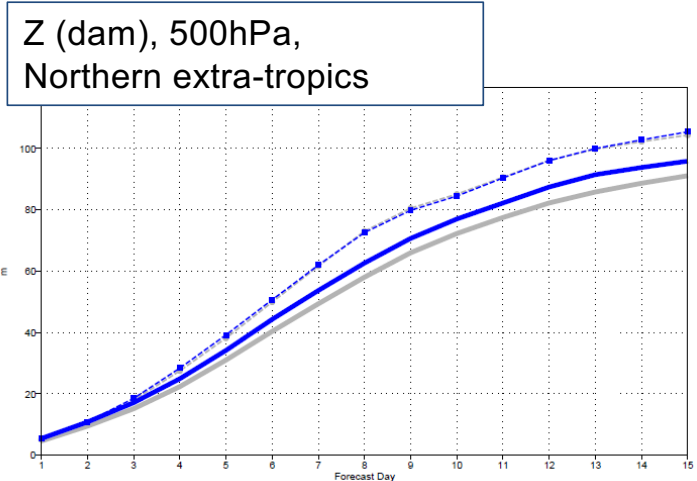
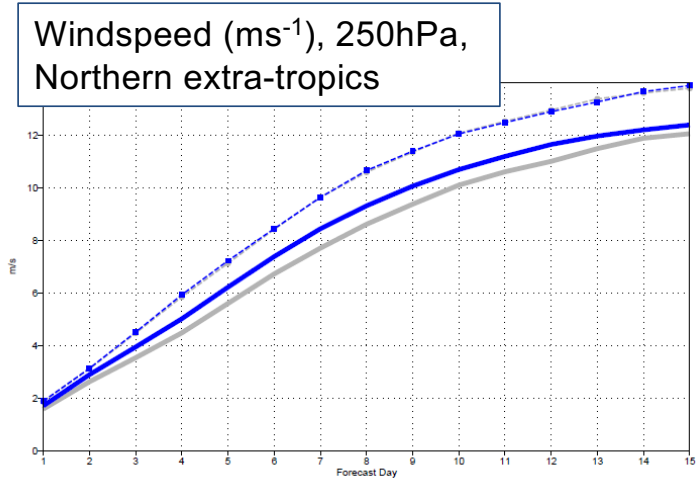
SPPT random pattern

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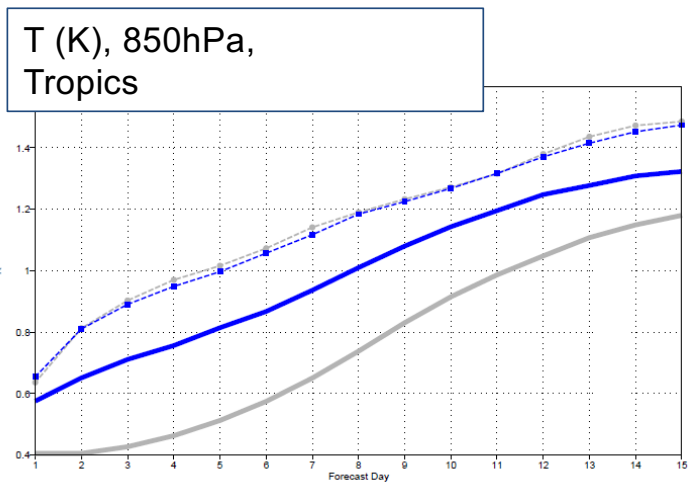
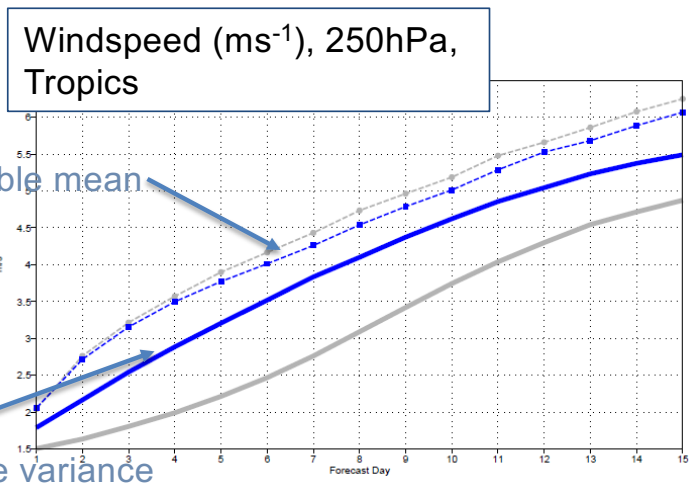


Ensemble forecasts: with **multi-scale** model uncertainty perturbations (SPPT)

Ensemble mean RMSE (“Error”) & standard deviation (“Spread”)



— IP only
 - - IP + SPPT3*
 (*3 scales)

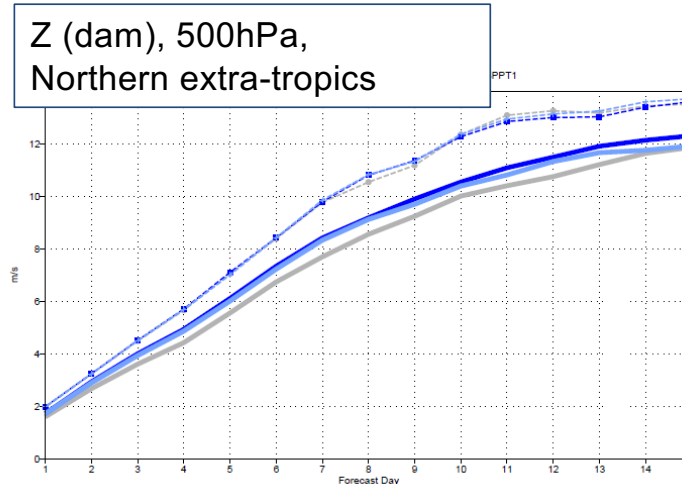
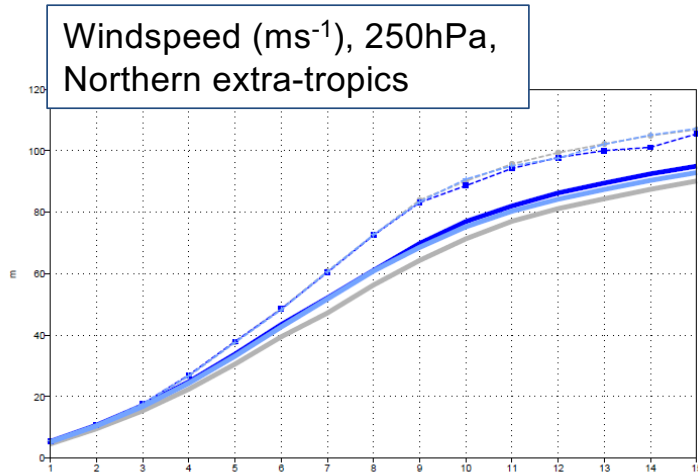


RMSE ensemble mean
 (“error”)
 RMS ensemble variance
 (“spread”)

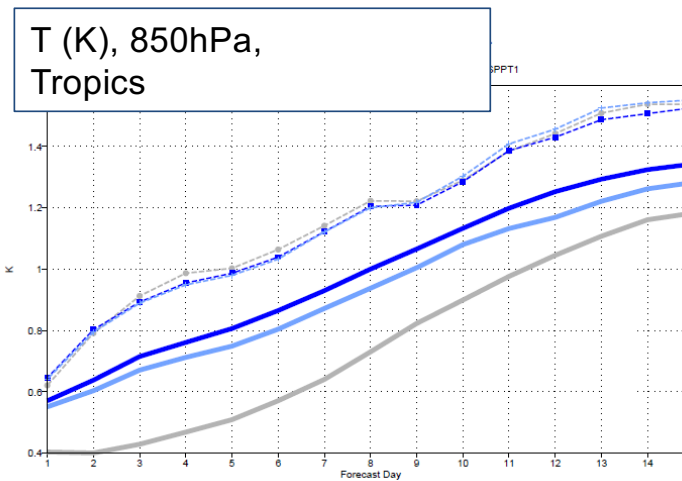
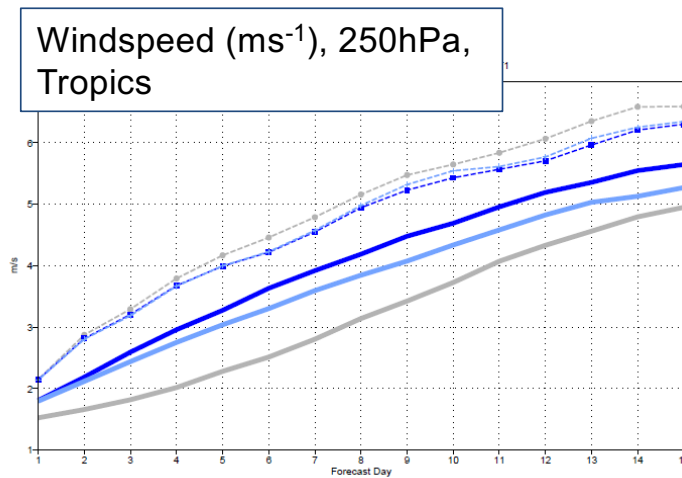
CY47R3
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Ensemble forecasts: with **multi-scale** model uncertainty perturbations (SPPT)

Ensemble mean RMSE (“Error”) & standard deviation (“Spread”)



- IP only
- IP + SPPT3* (*3 scales)
- IP + SPPT1** (**shortest scale only)



Some additional spread from **SPPT3** - 3rd scale deemed important for longer-range forecasts

CY47R3

TCo399L137, dt=1200s

11 dates (Dec 2019/Jan 2020)

8 perturbed fcs

Ensemble forecasts: with **multi-scale** model uncertainty perturbations (SPPT)

Scorecard of probabilistic skill (“fCRPS”) & ensemble standard deviation (“Spread”)

verified
against
analysis

verified
against
observations



Scorecard (summary):

IP + SPPT3* *versus* IP only
(*3 scales)

Spread:

Purple = more spread / Green = less spread

fCRPS:

Blue = more skillful / Red less skillful

Framed cell indicates statistically significant differences at the 95% confidence interval

CY47R3

TCo399L137, dt=1200s

30 dates (Dec 2019)

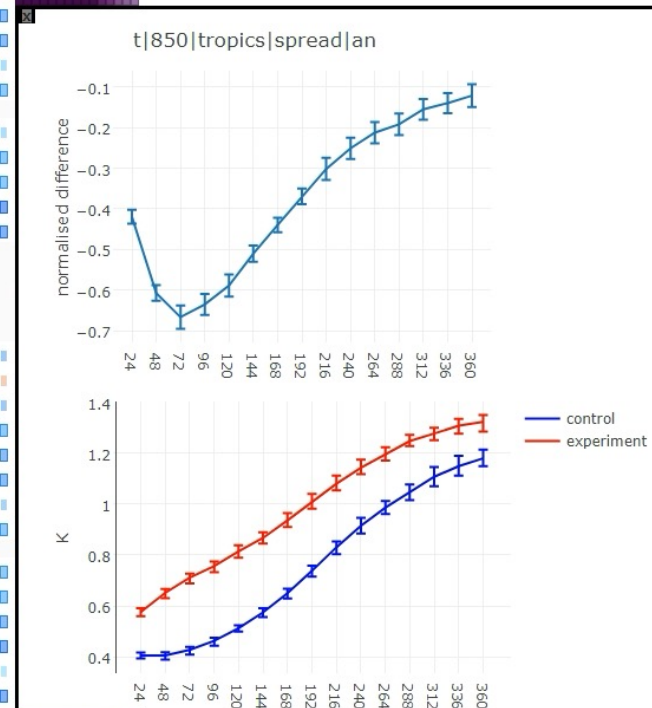
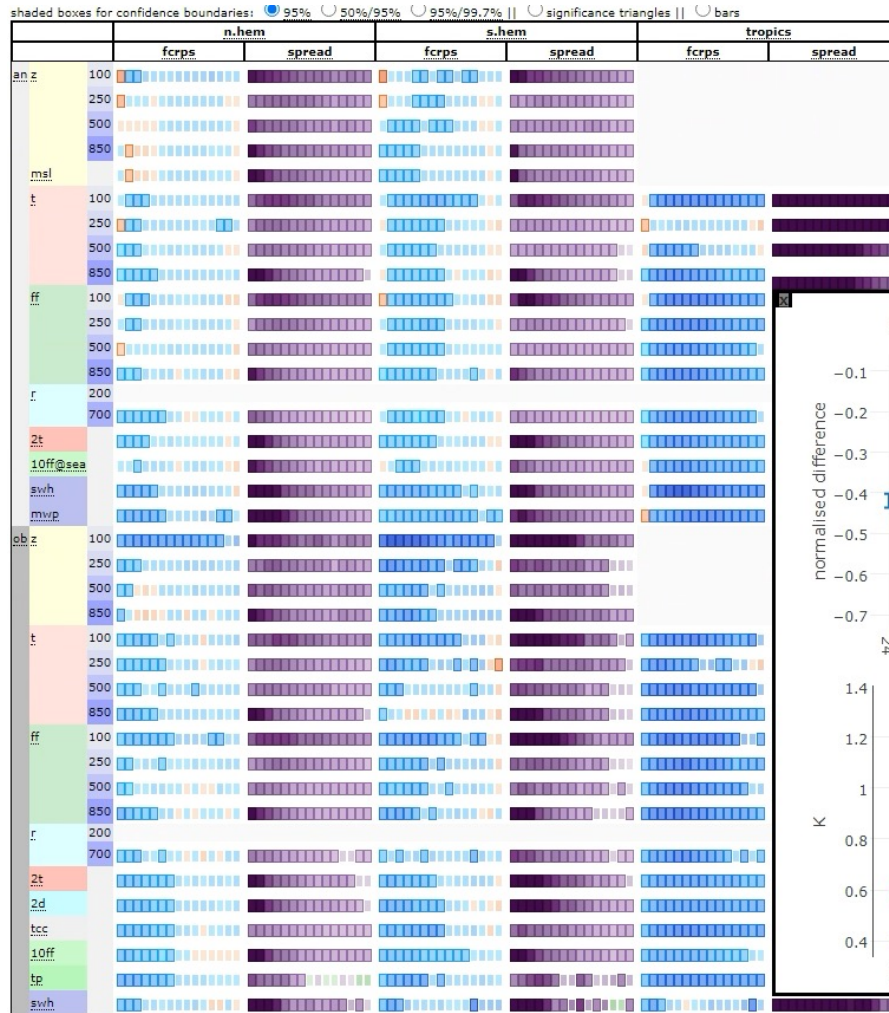
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Ensemble forecasts: with **multi-scale** model uncertainty perturbations (SPPT)

Scorecard of probabilistic skill ("fCRPS") & ensemble standard deviation ("Spread")

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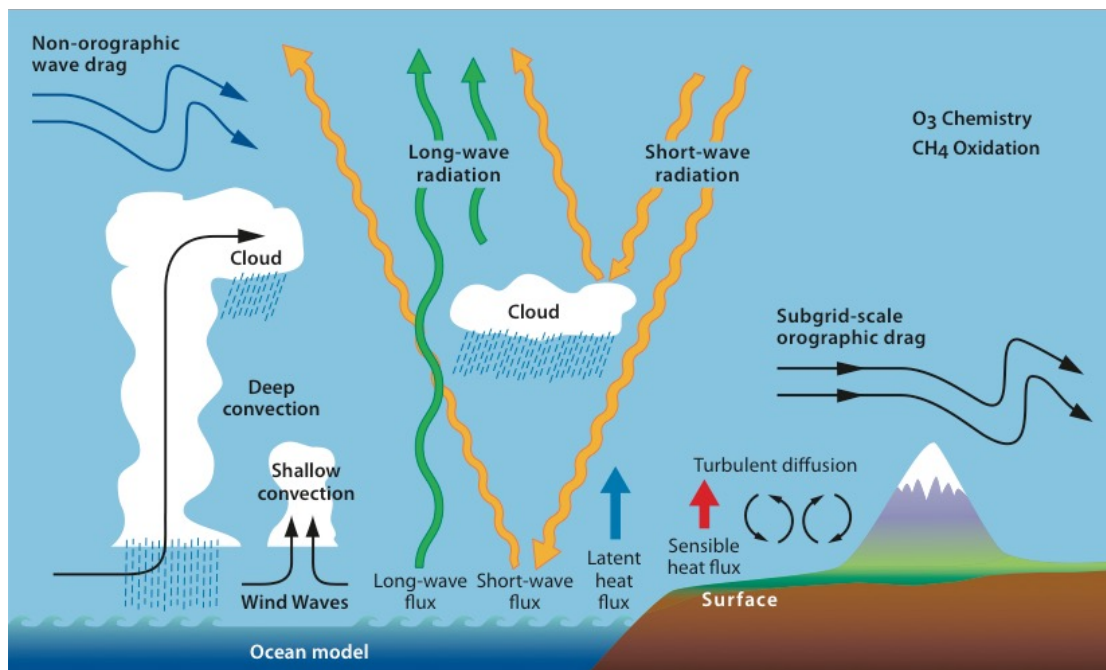
8 perturbed fcs

Summary: stochastic representation of model uncertainty in IFS

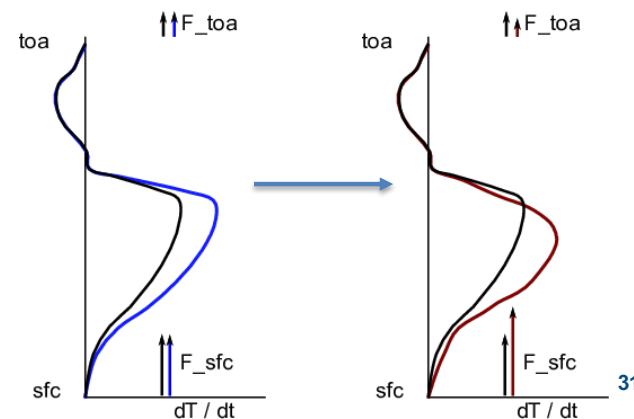
- Model uncertainty (MU) due to unresolved and misrepresented processes
- Without representing MU, ensemble forecasts are under-dispersive => over-confident
- Stochastic representations of model uncertainty can **improve ensemble reliability**
- SPPT: represents uncertainty due to sub-grid atmospheric physics parameterisations
 - **Medium-range:** increased ensemble spread, greater probabilistic skill
 - **Seasonal:** reduction in biases; better representation of MJO, ENSO, PNA regimes (Weisheimer et al., 2014, Phil. Trans. R. Soc. A)
- *Difficult to characterise sources of model uncertainty due to their small scales*

Stochastic representations of model uncertainty: outlook for IFS

Towards process-level model uncertainty representation

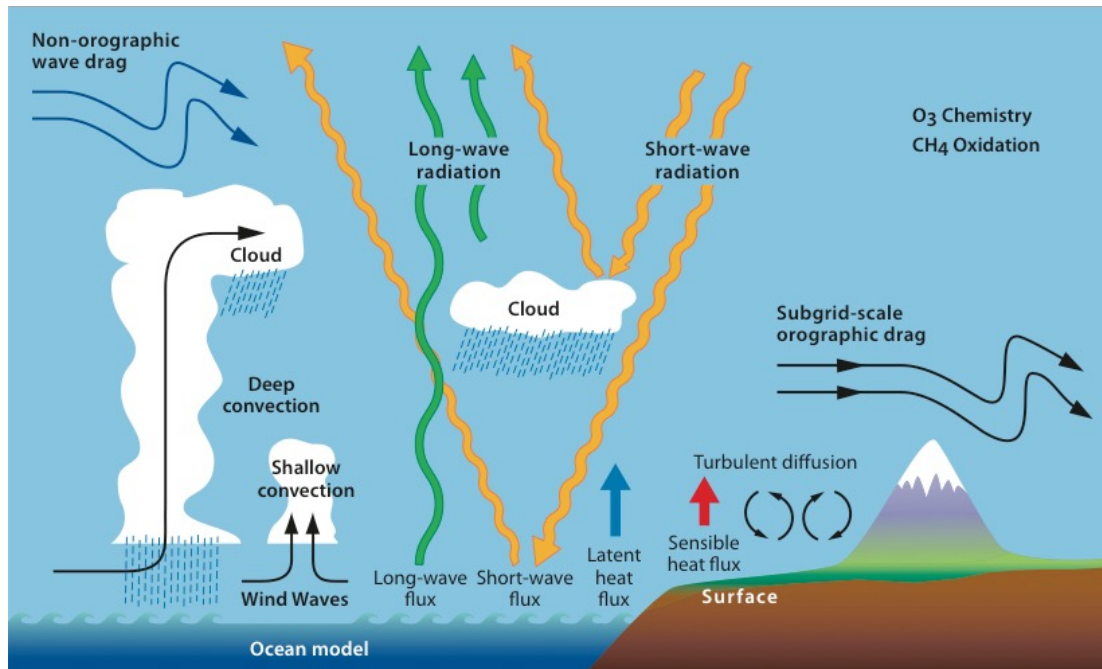


- **Aim:** to improve the physical consistency
- Local conservation of moisture, momentum, energy
- Generate flux perturbations at the top of atmosphere (TOA) and surface that are consistent with tendency perturbations within the atmospheric column
- Remove ad hoc tapering in boundary layer
- Include multi-variate aspects of uncertainties



Stochastic physics: outlook for IFS

Towards process-level model uncertainty representation



Stochastically Perturbed Parametrisations (SPP)

(Lang et al., 2021, QJRMS; Ollinaho et al., 2017, QJRMS)

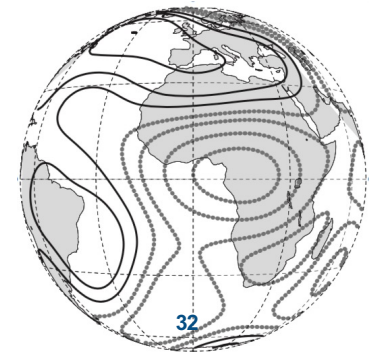
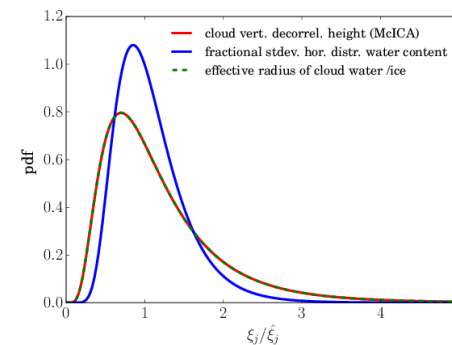
- Embed stochasticity inside IFS parametrisations
- Perturb parameters/variables directly
- Specify spatial/temporal correlations
- Target uncertainties that matter (level of uncertainty and impact)
- Require that stochastic schemes converge to deterministic schemes in limit of vanishing variance

Stochastically perturb parameters/variables in the physics parametrisations ($\hat{\xi}_j$):

$$\xi_j = \hat{\xi}_j \exp(\Psi_j)$$

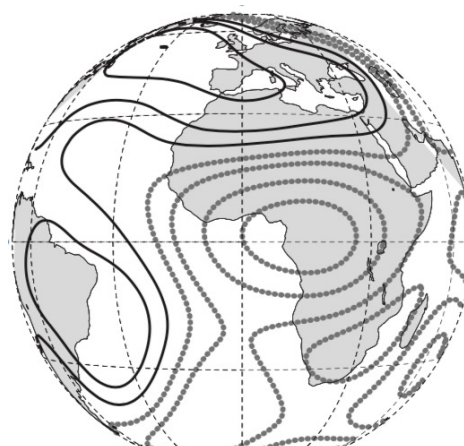
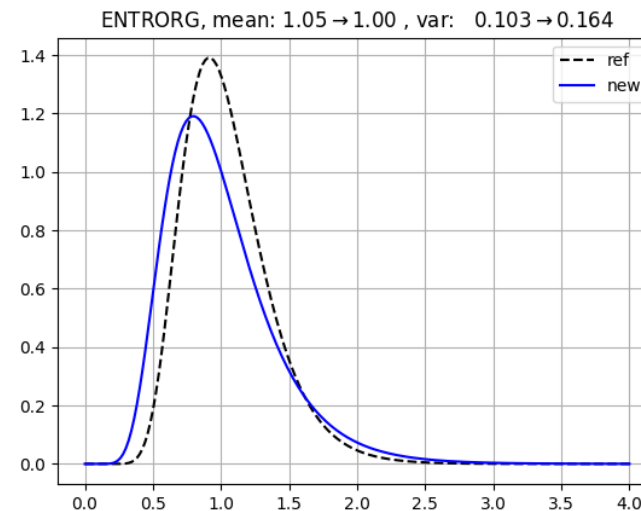
where

$$\Psi_j \sim \mathcal{N}(\mu_j, \sigma_j^2)$$

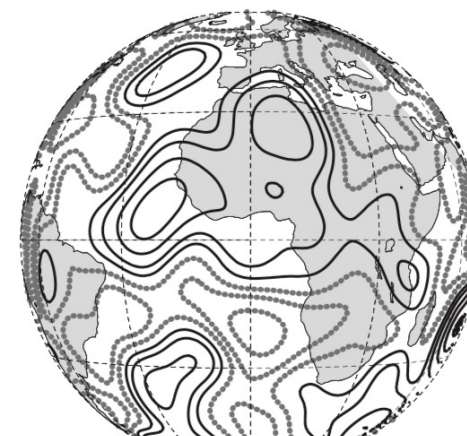


Revisions to SPP

- Work on a revision of SPP completed: Lang et al (2021), <https://doi.org/10.1002/qj.3978>
- Summary of changes
 - probability distributions (mean and variance)
 - correlation scale
 - additional perturbed quantities (total 27):
 - Cloud scheme
 - Convection scheme
 - BL scheme

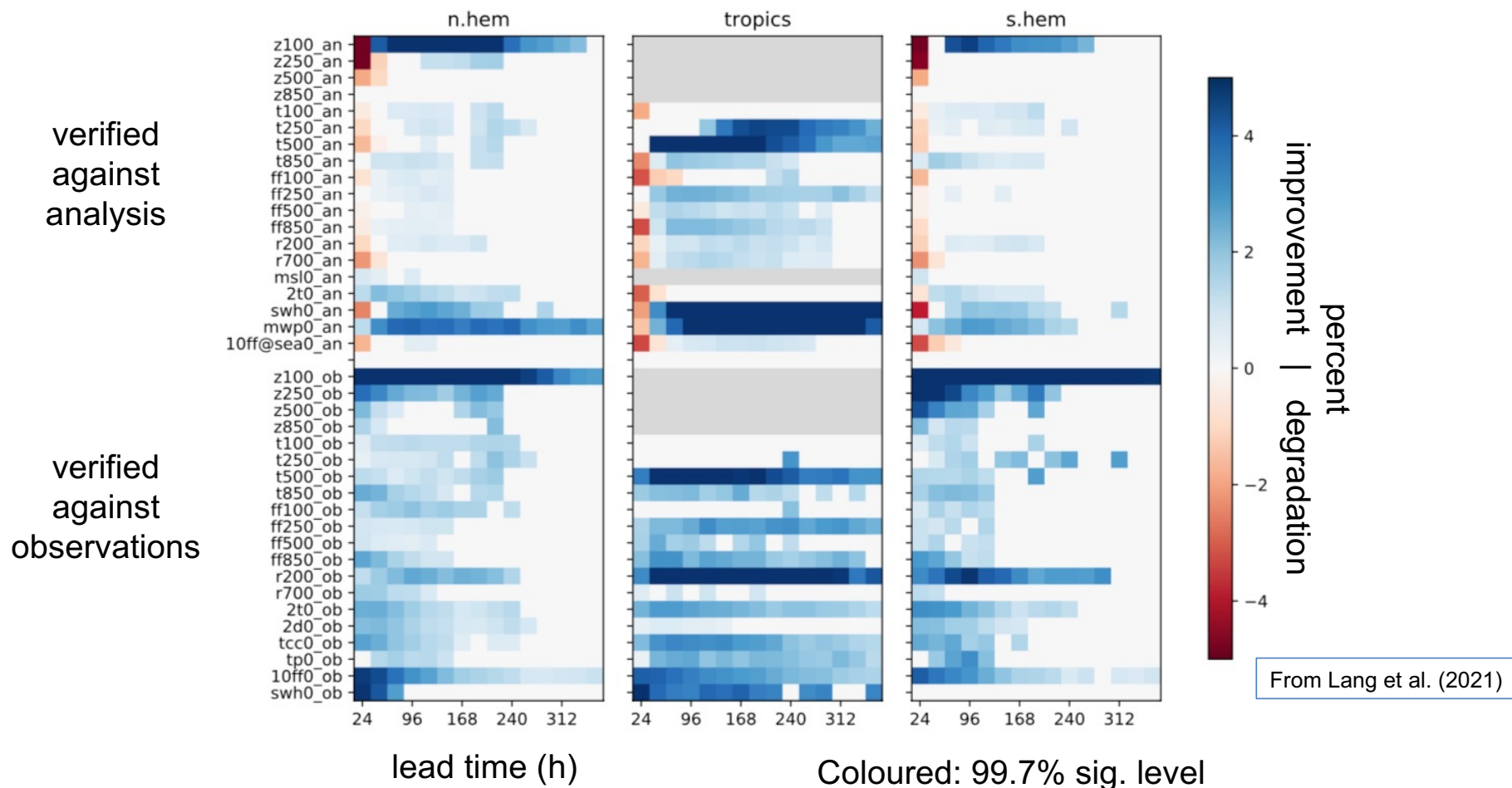


2000 km

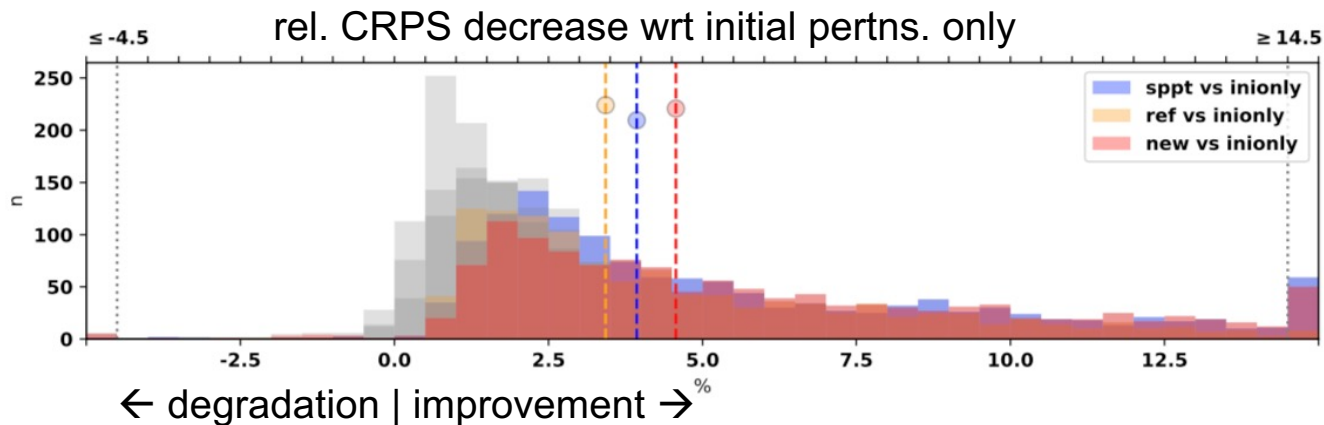


1000 km

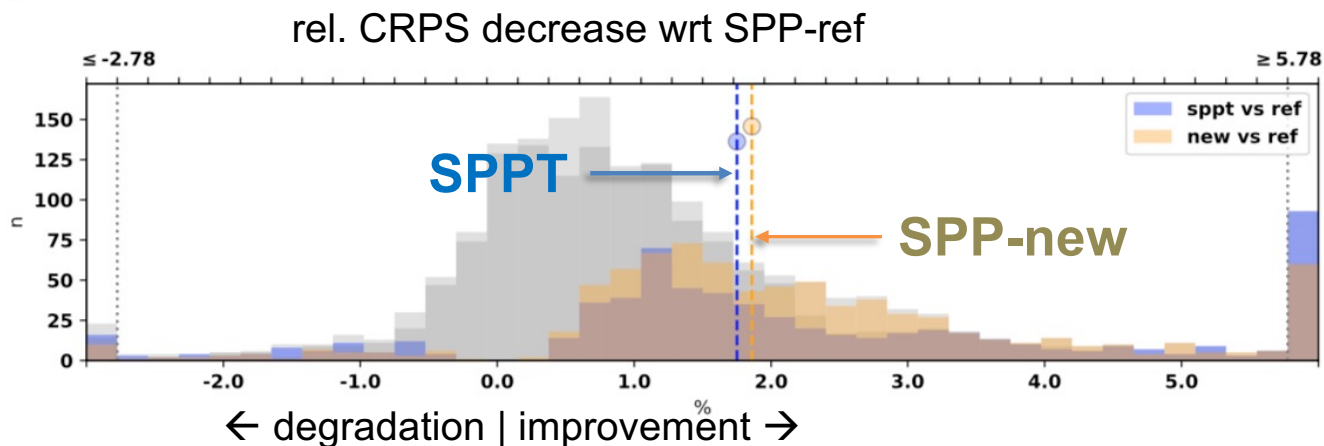
SPP-new versus SPP-ref scorecard showing fCRPS changes



Histogram of relative CRPS changes: SPPT, SPP



b)



lead times 24 h ... 360 h
 combination of variables and
 levels evaluated in scorecard
 3 regions: N-Hem, S-Hem,
 tropics

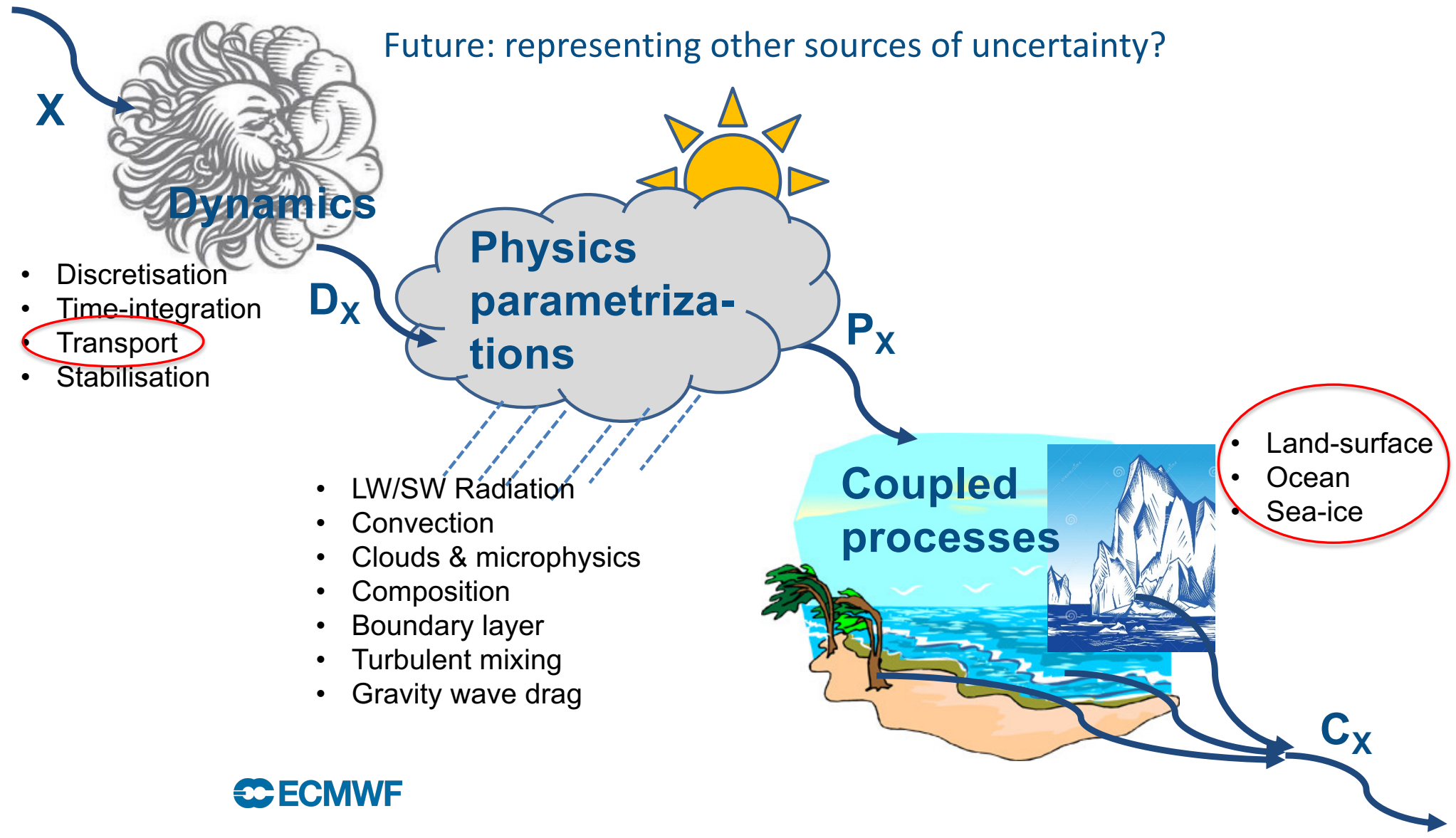
vertical lines: median change

based on 8 members and fair
 CRPS, boreal summer + boreal
 winter, 212 start dates, TCo399

colour: stat. significance 99.7%,
 grey otherwise

**Enters operations
 in CY49R1 (2024)!**

Future: representing other sources of uncertainty?



STOCHDP:

Stochastically perturbed semi-Lagrangian (SL) departure point (DP) estimates

Diamantakis & Magnusson (2016):

- Explored convergence rate of the iterative DP estimate
- Slowest convergence \leftrightarrow most complex flow (strong shear / curvature)
- Example: Typhoon Neoguri:
 - HRES forecast: initialised: 2014-07-05, 00UTC

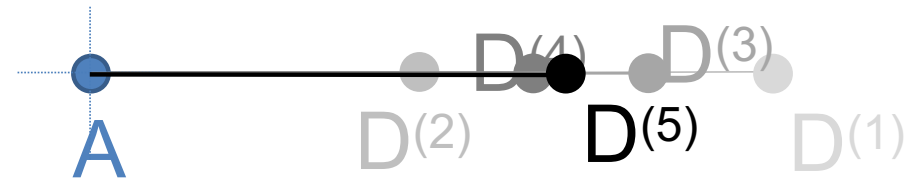


Fig. 1c: $t+96h$, 850hPa windspeeds

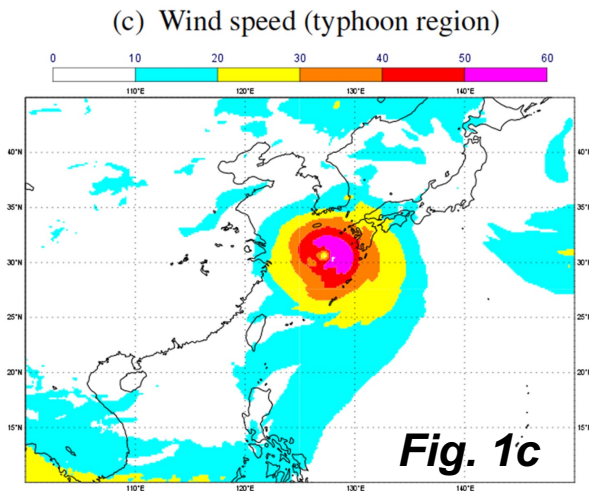
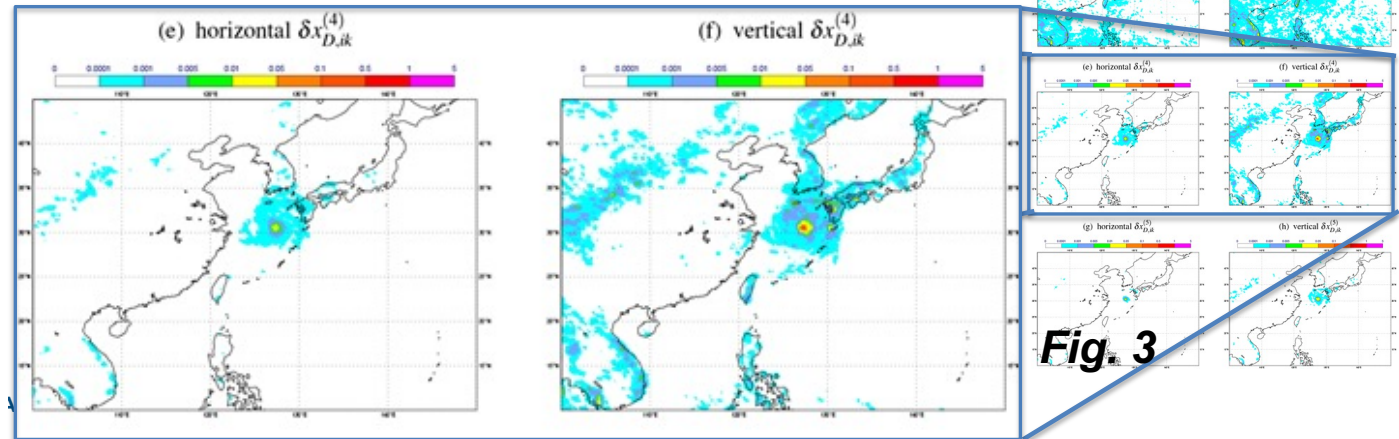


Figure 3: difference in DP estimate between consecutive iterations (scaled)



STOCHDP:

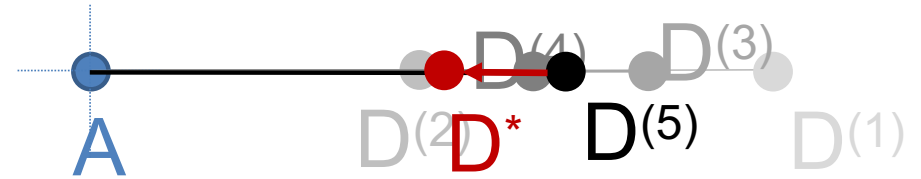
Stochastically perturbed semi-Lagrangian (SL) departure point (DP) estimates

Model uncertainty scheme, "STOCHDP":

- use the DP estimate convergence rate to attribute MU:

$$D^* = D^{(5)} + r(D^{(5)} - D^{(5-i)}), i = 1..4$$

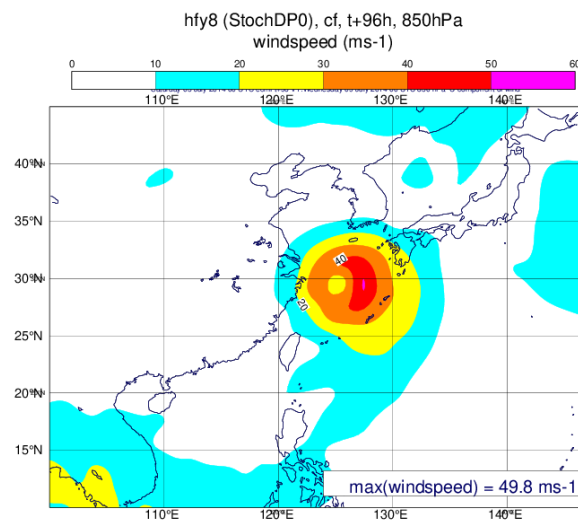
where D^* is the perturbed DP and r is a random number



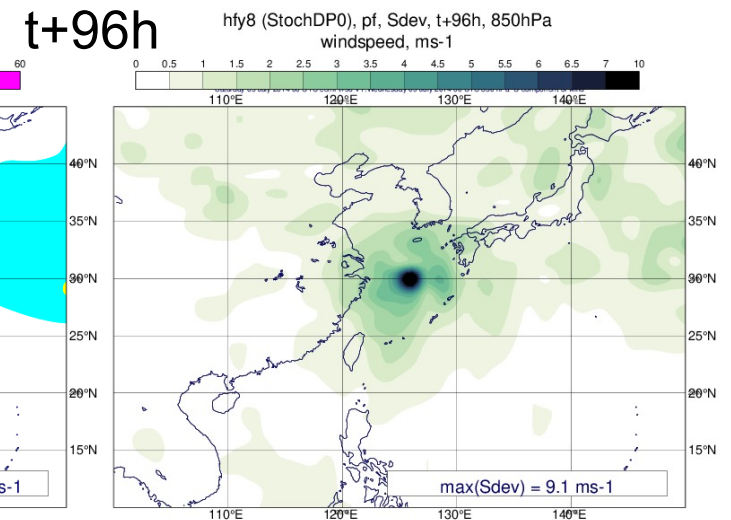
➤ STOCHDP represents MU from SL advective winds

Early results, e.g.:

- Typhoon Neoguri case
- ENS: STOCHDP only
- TCo639L91, dt=720s
- 20+1 members
- Peak ENS stdev develops and tracks with TC



Control forecast



Ensemble stdev

Summary

- Including a representation of model uncertainty can improve the reliability of ensemble forecasts
- "Model uncertainty" describes inaccuracies due to the model integrations
- Using stochastic physics schemes enables representation of the model uncertainty arising from the parametrization of unresolved atmospheric physics
- Current stochastic physics scheme used in the IFS: SPPT
- Outlook: new scheme "SPP" improves the physical consistency of the stochastic physics perturbations
- Ongoing: exploring stochastic perturbations to represent model uncertainty in the dynamics – STOCHDP

References

- Buizza, R., et al. (1999), Stochastic representation of model uncertainties in the ECMWF ensemble prediction system. QJRMS, 125: 2887-2908. <https://doi.org/10.1002/qj.49712556006>
- Diamantakis, M., & Magnusson, L. (2016). Sensitivity of the ECMWF Model to Semi-Lagrangian Departure Point Iterations, MWR, 144(9), 3233-3250. <https://doi.org/10.1175/MWR-D-15-0432.1>
- Lang, STK, et al. (2021), Revision of the Stochastically Perturbed Parametrisations model uncertainty scheme in the Integrated Forecasting System. QJRMS, 147: 1364–1381. <https://doi.org/10.1002/qj.3978>
- Lock, S-J, et al. (2019), Treatment of model uncertainty from radiation by the Stochastically Perturbed Parametrization Tendencies (SPPT) scheme and associated revisions in the ECMWF ensembles. QJRMS, 145 (Suppl. 1): 75- 89. <https://doi.org/10.1002/qj.3570>
- Ollinaho, P., et al. (2017), Towards process-level representation of model uncertainties: stochastically perturbed parametrizations in the ECMWF ensemble. QJRMS, 143: 408-422. <https://doi.org/10.1002/qj.2931>
- Palmer et al., 2009: Stochastic parametrization and Model Uncertainty, ECMWF Tech. Mem., **598**, pp. 42
- Weisheimer A., et al. (2014), Addressing model error through atmospheric stochastic physical parametrizations: impact on the coupled ECMWF seasonal forecasting system. *Phil. Trans. R. Soc. A.* **372**: 20130290. <https://doi.org/10.1098/rsta.2013.0290>

Further reading & upcoming workshop

In 2016, we undertook an extensive review of existing and future efforts in model uncertainty representation – as a Special Topic paper for our Scientific Advisory Committee

Report covers:

- Literature review of model uncertainty work
- Descriptions/discussions of SPPT / SKEB / SPP schemes
- Impacts of the schemes in the IFS (EDA; short / medium / extended / longer forecast ranges)
- Proposals for future directions – improvements to SPPT; extensions to SPP; new approaches

Leutbecher, M., Lock, S.-J., Ollinaho, P., Lang, S.T.K., et al. (2017), Stochastic representations of model uncertainties at ECMWF: state of the art and future vision. Q.J.R. Meteorol. Soc, 143: 2315-2339. <https://doi.org/10.1002/qj.3094>